

From Dwarfs to Giants: the history of Star-formation since $z \sim 1$

The Redshift One LDSS3 Emission line Survey

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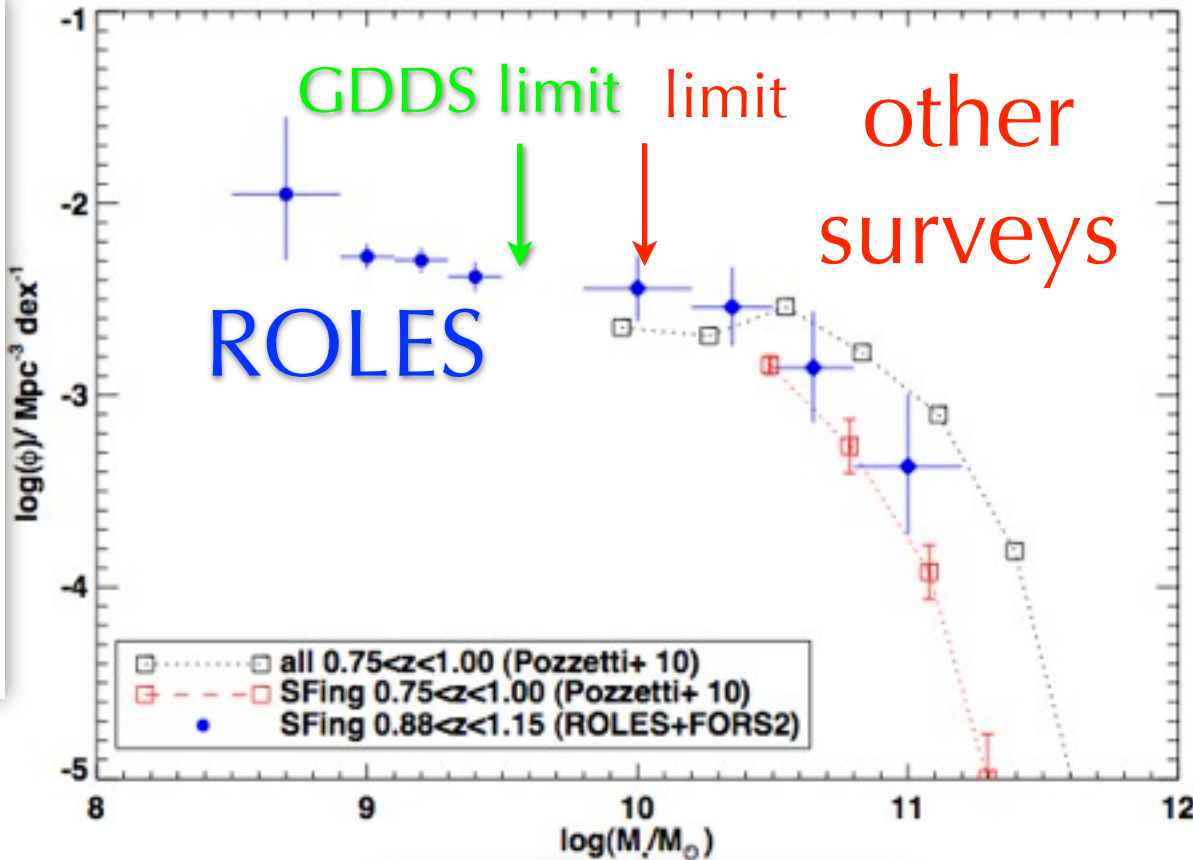
Karl Glazebrook (Swinburne)

Richard Bower (Durham)

I. Baldry, G. Davies, G. Hau, I-H. Li, P. McCarthy, M. Sawicki

Current $z \sim 1$ spectroscopic surveys

Number density



stellar mass

Low mass galaxies
much more
numerous than
their higher mass
counterparts
=> important for
-metal enrichment
of Universe
-reionisation

ROLES

Two fields: **GOODS-S/CDFS**
& **MS1054 FIRES**

$22.5 < K_{AB} < 24.0$

$8.5 < \log(M_s/M_{\text{Sun}}) < 9.5$

7000Å--8000Å

band-limited spectroscopy

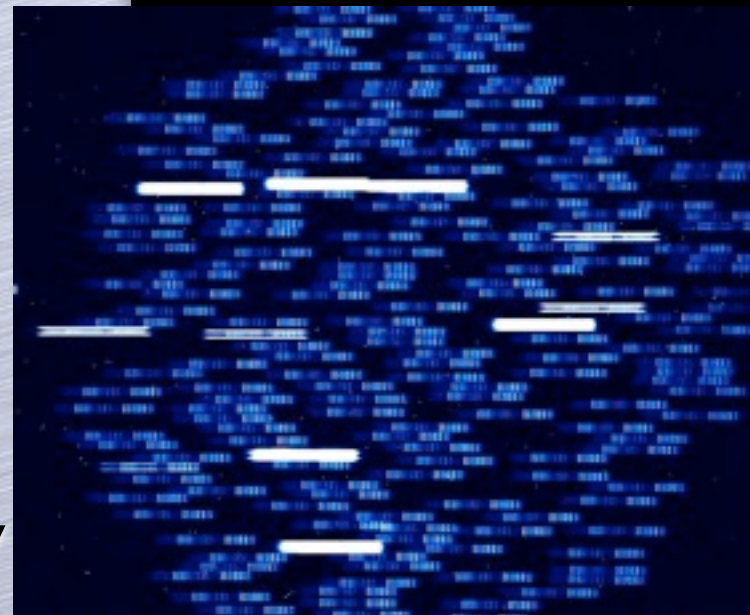
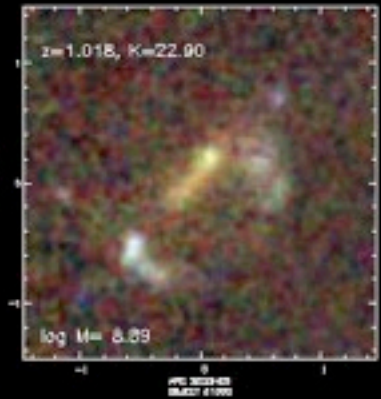
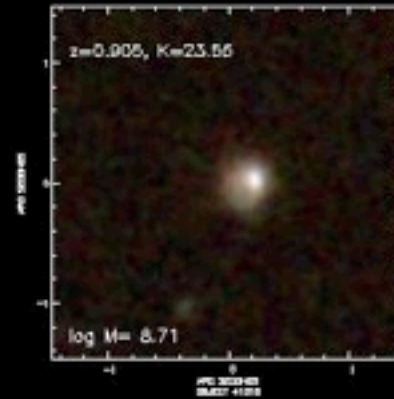
[OII] λ 3727 : $0.889 < z < 1.149$

~4 hours with

LDSS3 (6.5-m Magellan)

~200 slits per mask, 19 masks total

CDFS/GOODS-S



285 $z \sim 1$ dwarf galaxies
with
 $\text{SFR}_{\text{corr}} \approx 0.3 M_{\odot} \text{yr}^{-1}$

Deriving local SFR calibrations from SDSS

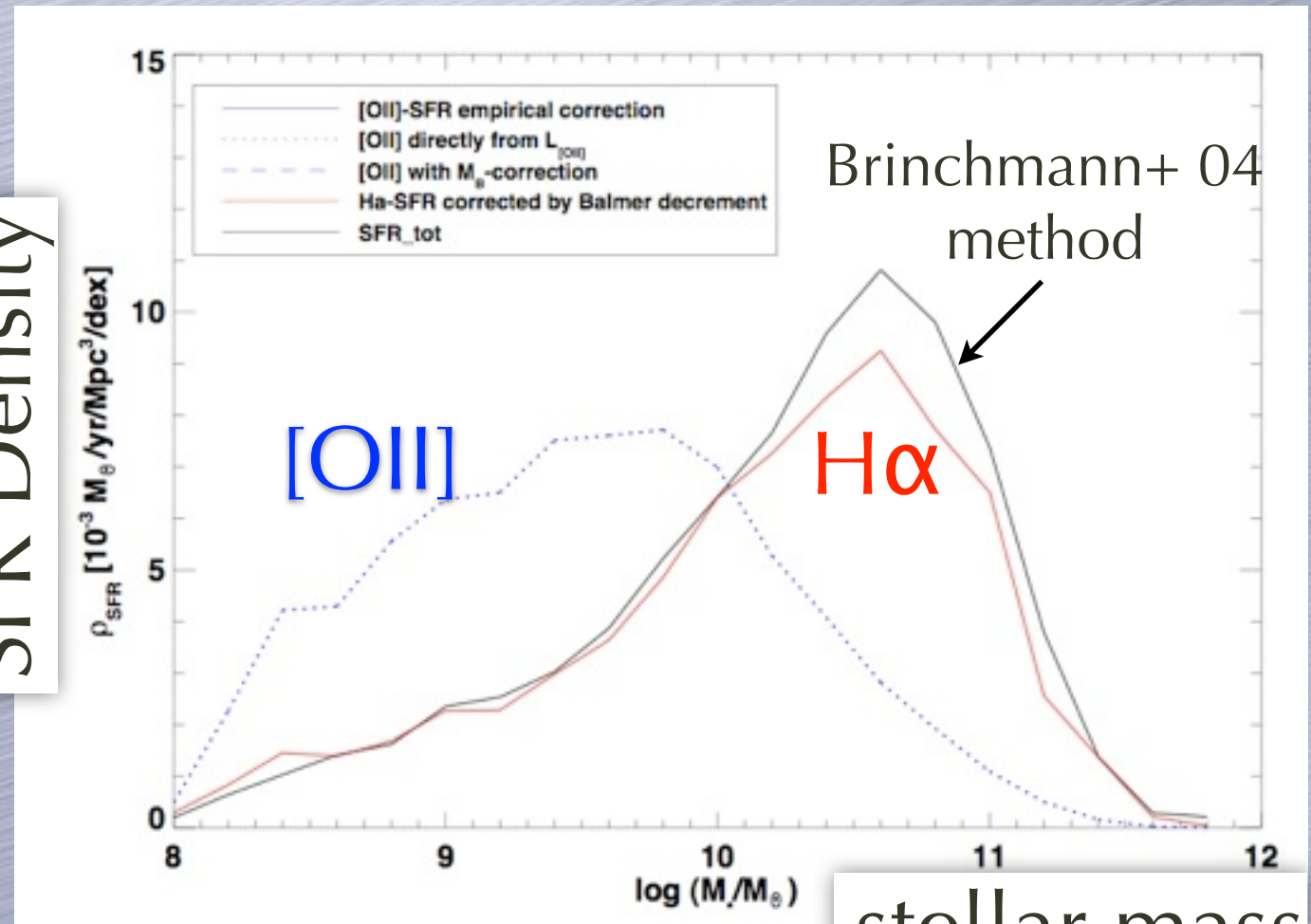
Conversion $L([\text{OII}]) \leftrightarrow \text{SFR}$
depends on metallicity, dust and
ionisation parameter

- Check consistency between **SFR indicators**

(using $\text{H}\alpha$, $[\text{OII}]$, u -band, FUV from GALEX) in Stripe 82

- **SFRD** as a function of **stellar mass**

SFR Density



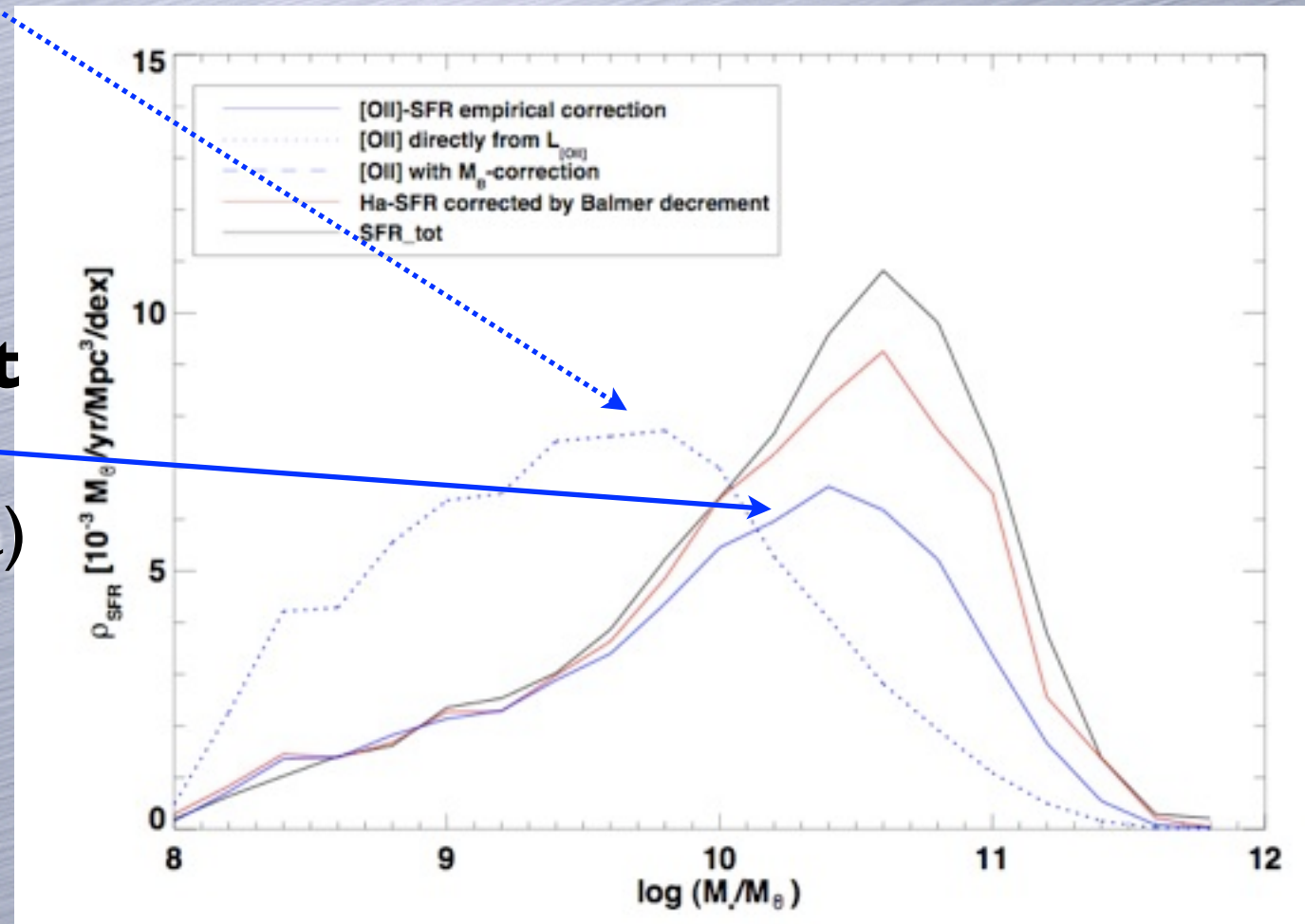
stellar mass

Deriving local SFR calibrations from SDSS

Conversion $L([\text{OII}]) \leftrightarrow \text{SFR}$
depends on metallicity, dust and
ionisation parameter

Constant scaling
from $L([\text{OII}])$ to SFR
(as often assumed)
gives v. different
answer!

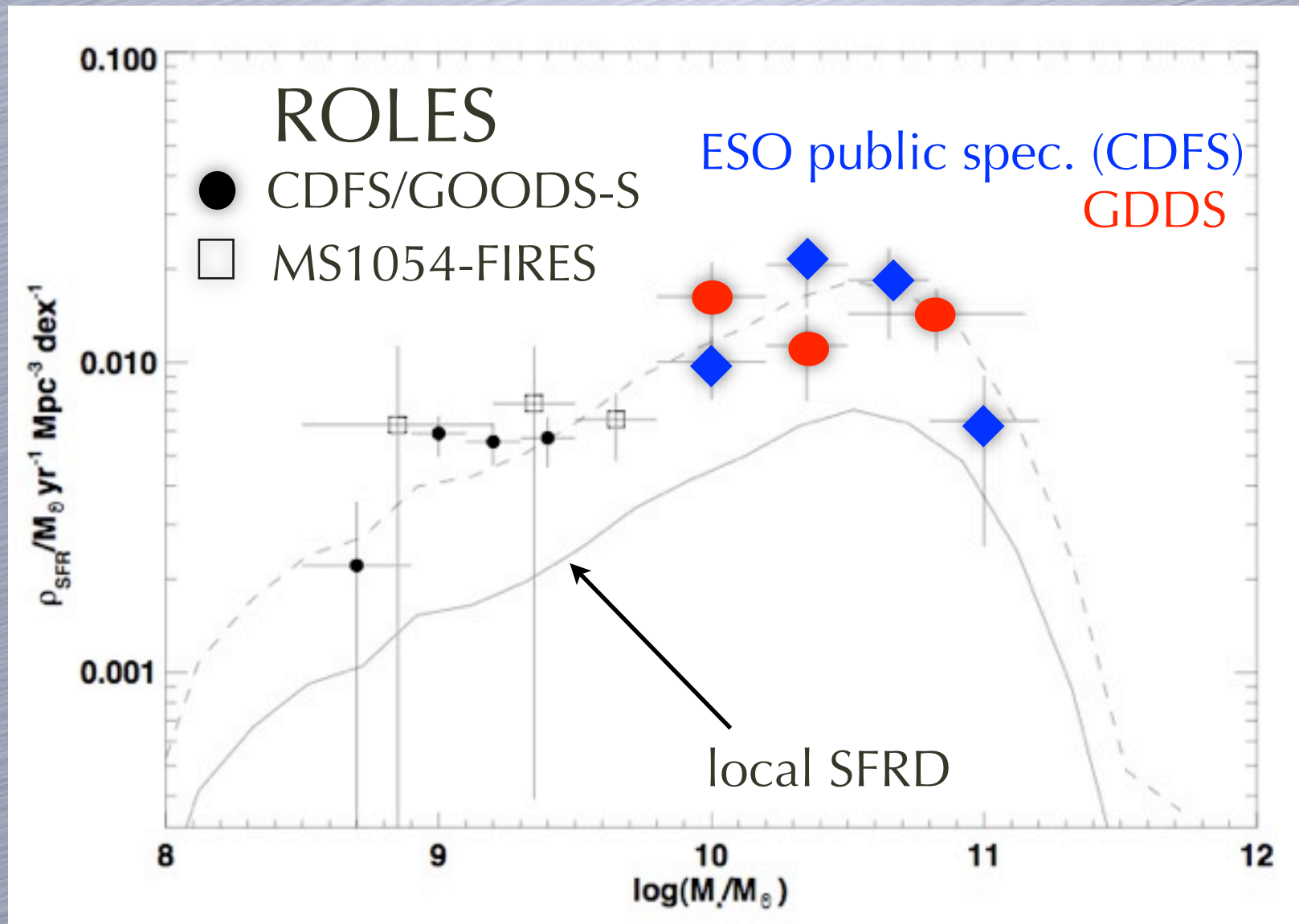
New, **empirical
mass-dependent
correction**
(Gilbank et al. 2010a)



$z \sim 1$ SFRD

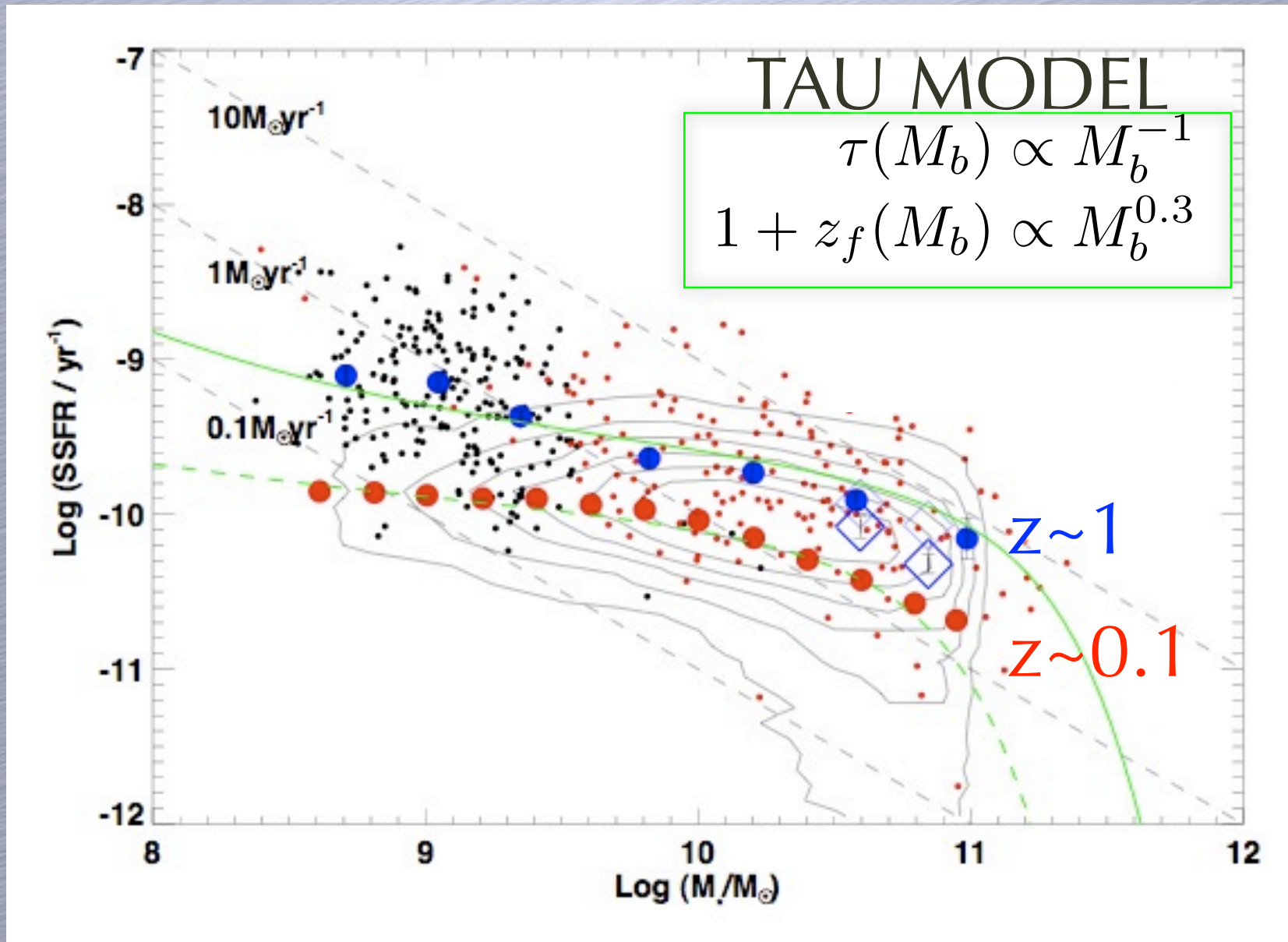
Empirically-corrected [OII] SFRD

Gilbank et al. 2010b, MNRAS 405, 2419



Specific SFR--Mass Relation

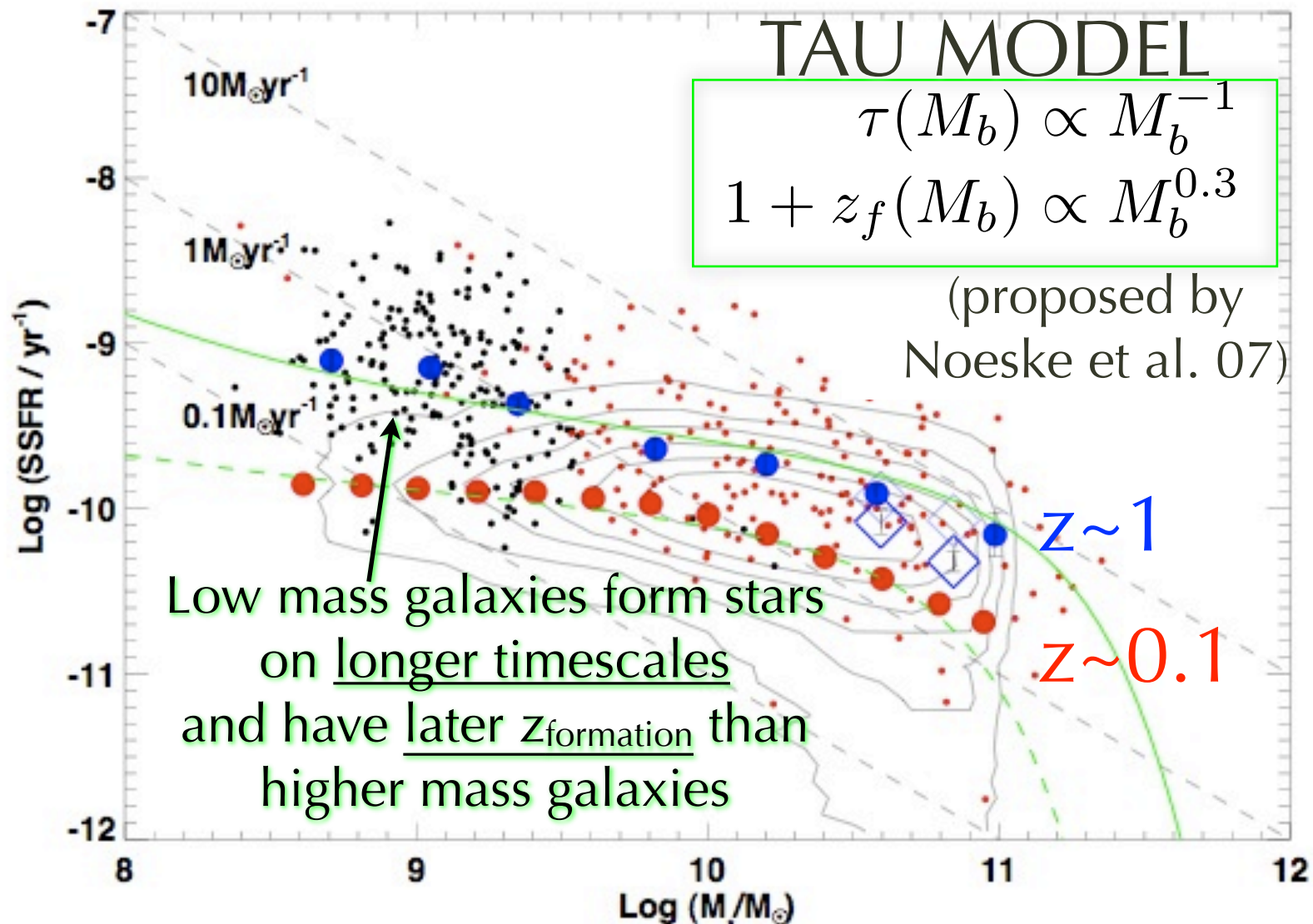
$$\text{SFR}/M_{\text{stellar}} = M/M \sim 1/t$$



Specific SFR--Mass Relation

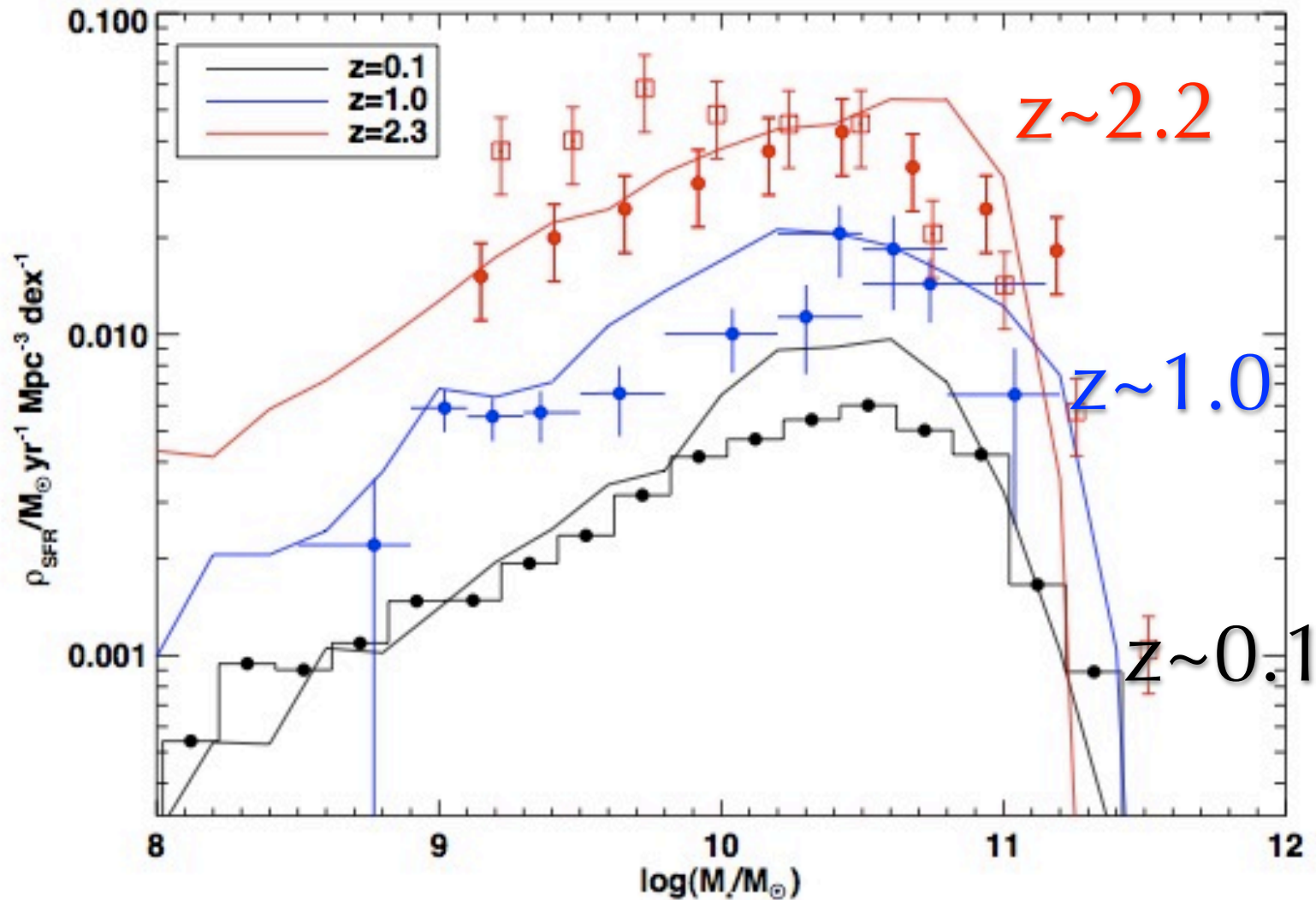
$$\text{SFR}/M_{\text{stellar}} = M/\dot{M} \sim 1/t$$

$1/t_{\text{SFR}}$



Comparison with Galform

[updated version of Bower et al. (2006)]
Gilbank et al. 2011, MNRAS, 414, 304



Summary

- New empirical mass-dependent correction for [OII]-SFR *previous corrections all underestimate*
- ROLES: [OII] SFRs for dwarf gals at $z \sim 1$ (robust!)
Extinction negligible for these low mass sources ($8.5 < \log(M_/M_\odot) < 9.5$), so obscured SF not problematic*
- Evolution of SFRD from $z \sim 1$ to $z \sim 0.1$ is an equal decrease for all mass galaxies
- SSFR--mass potentially more sensitive test
 - \Rightarrow high mass galaxies formed stars earlier and faster (“down-sizing”)
Down-sizing extremely subtle! Must be careful about systematics