

# Probing outflows/inflows from DLA host galaxies with spatially extended $\text{Ly}\alpha$ emission

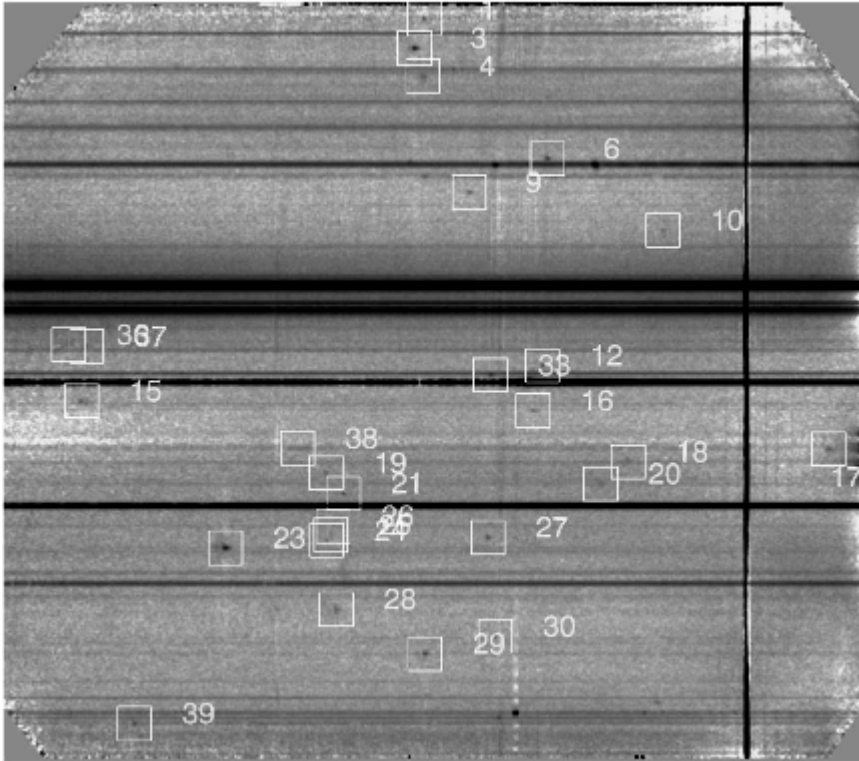
**Martin Haehnelt**

in collaboration with:

**Luke Barnes** , **Michael Rauch**,

**George Becker** , **Wal Sargent**, **Edoardo**

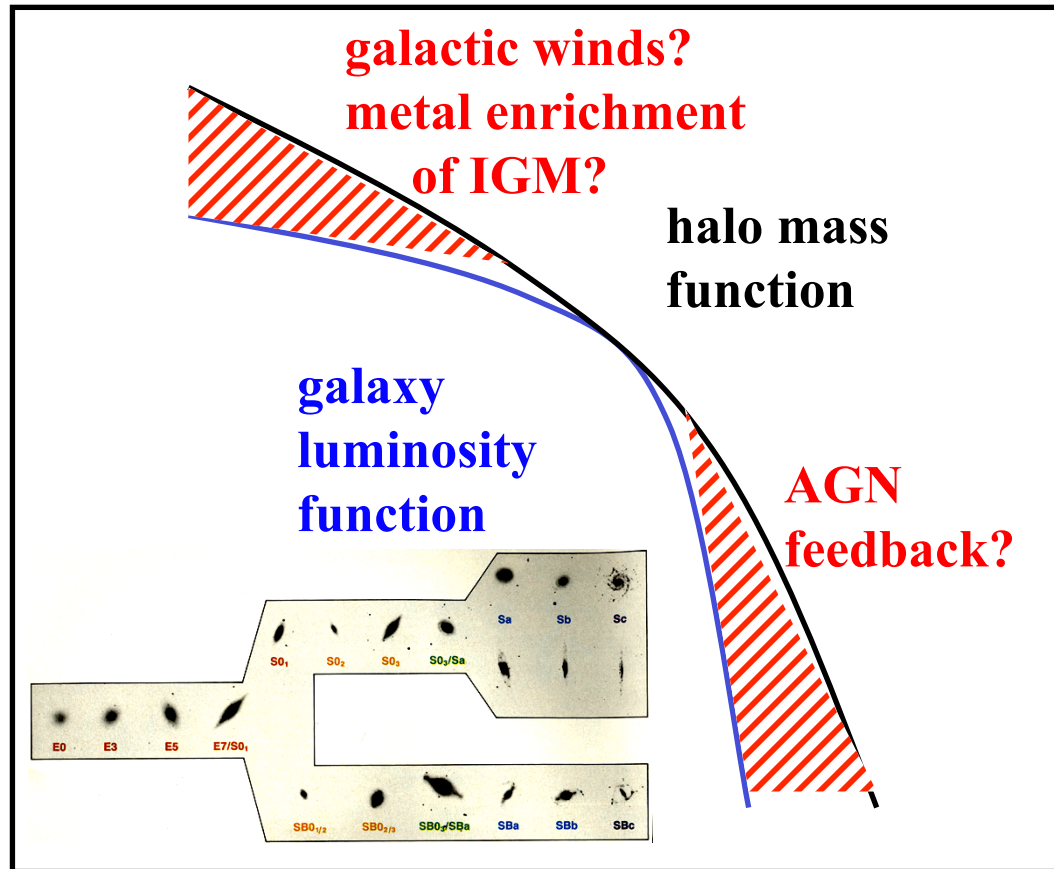
**Tescari**, **Matteo Viel**



Durham, 21 July 2011

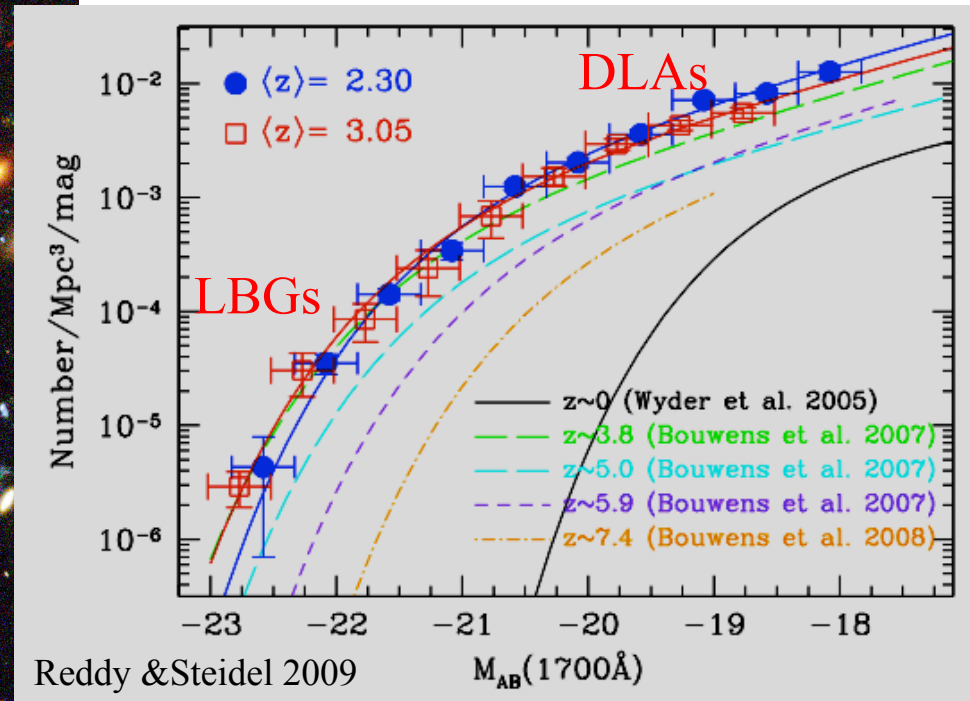
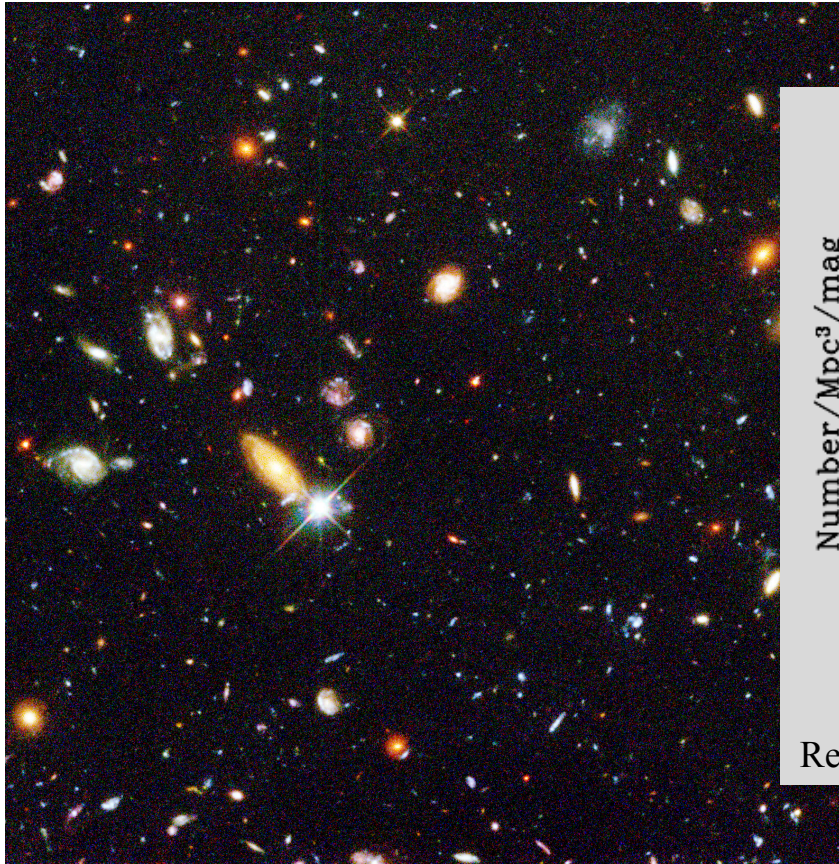


space density



luminosity/mass

Low mass galaxies at high redshift should reveal important clues on  
- how feedback works and  
- how gas gets into galaxies

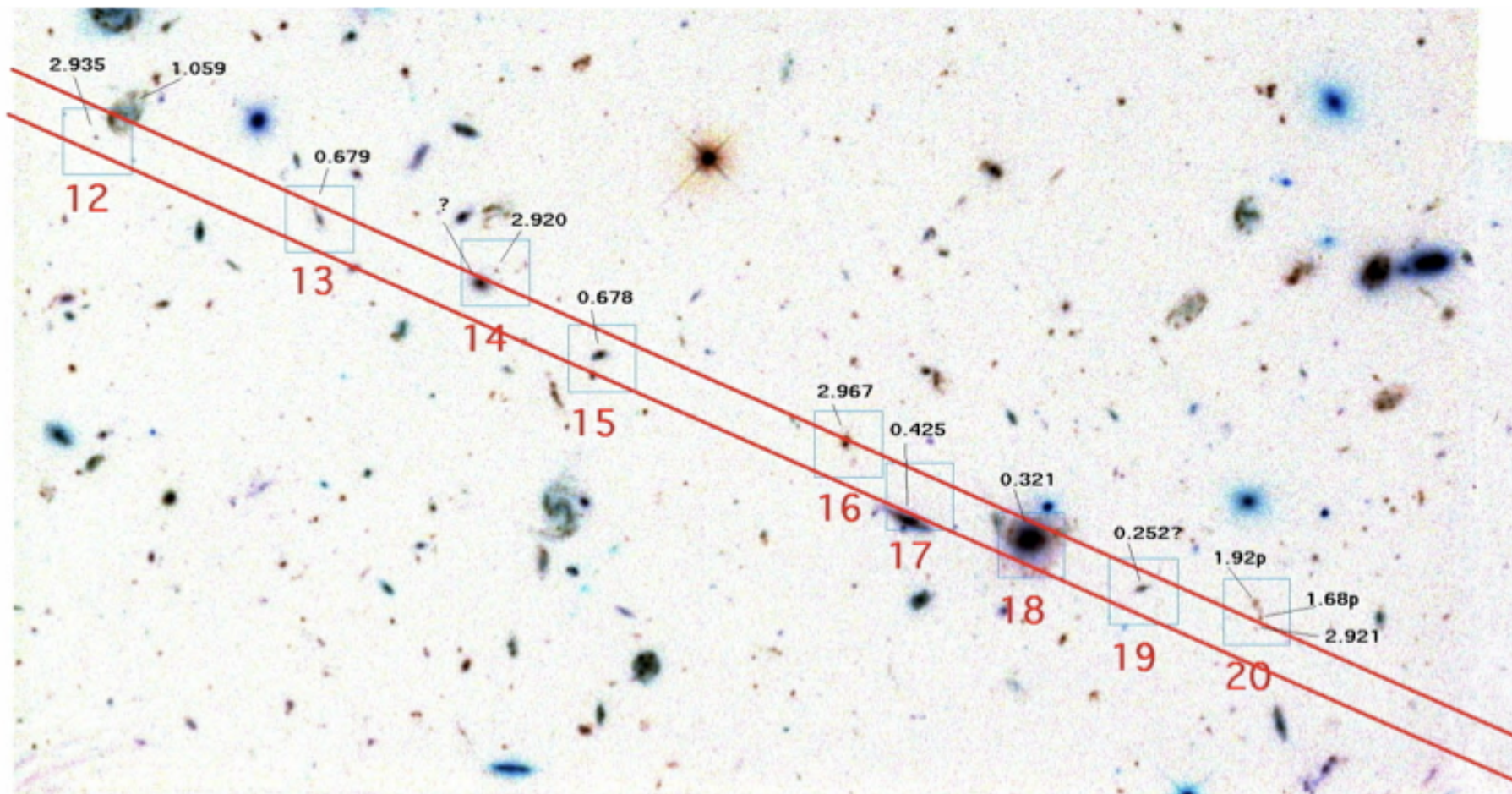


DLAs should populate the extreme faint end of the LBG populations

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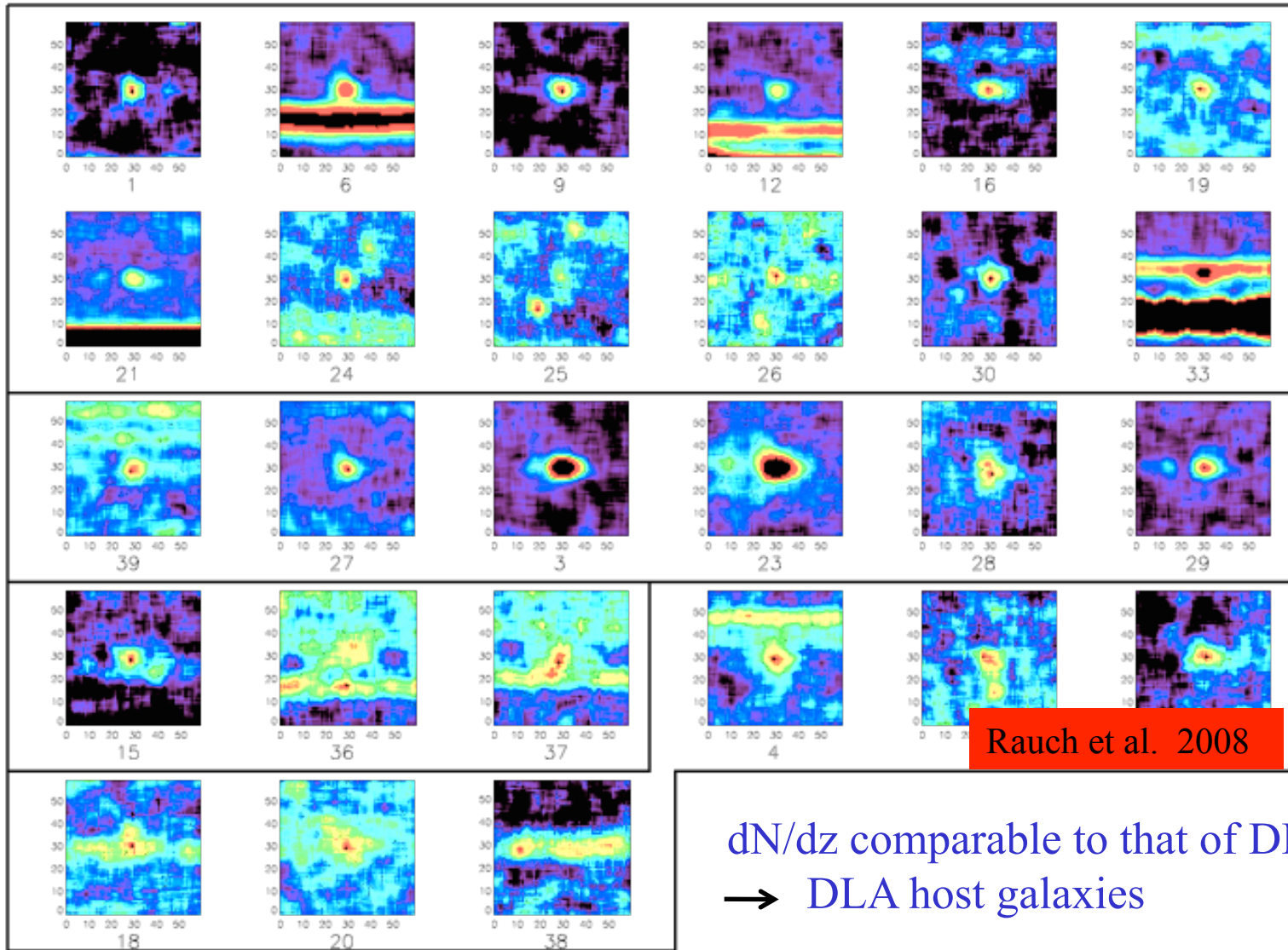
# Blind ultra-deep longslit spectroscopy



Durham, 21 July 2011

Michael Rauch et al.

$\rightarrow$  2300 km/s  $\leftarrow$  **27 spatially extended Ly $\alpha$  emitters at 2.6 < z < 3.6**  
 $\rightarrow$  116 kpc  $\leftarrow$



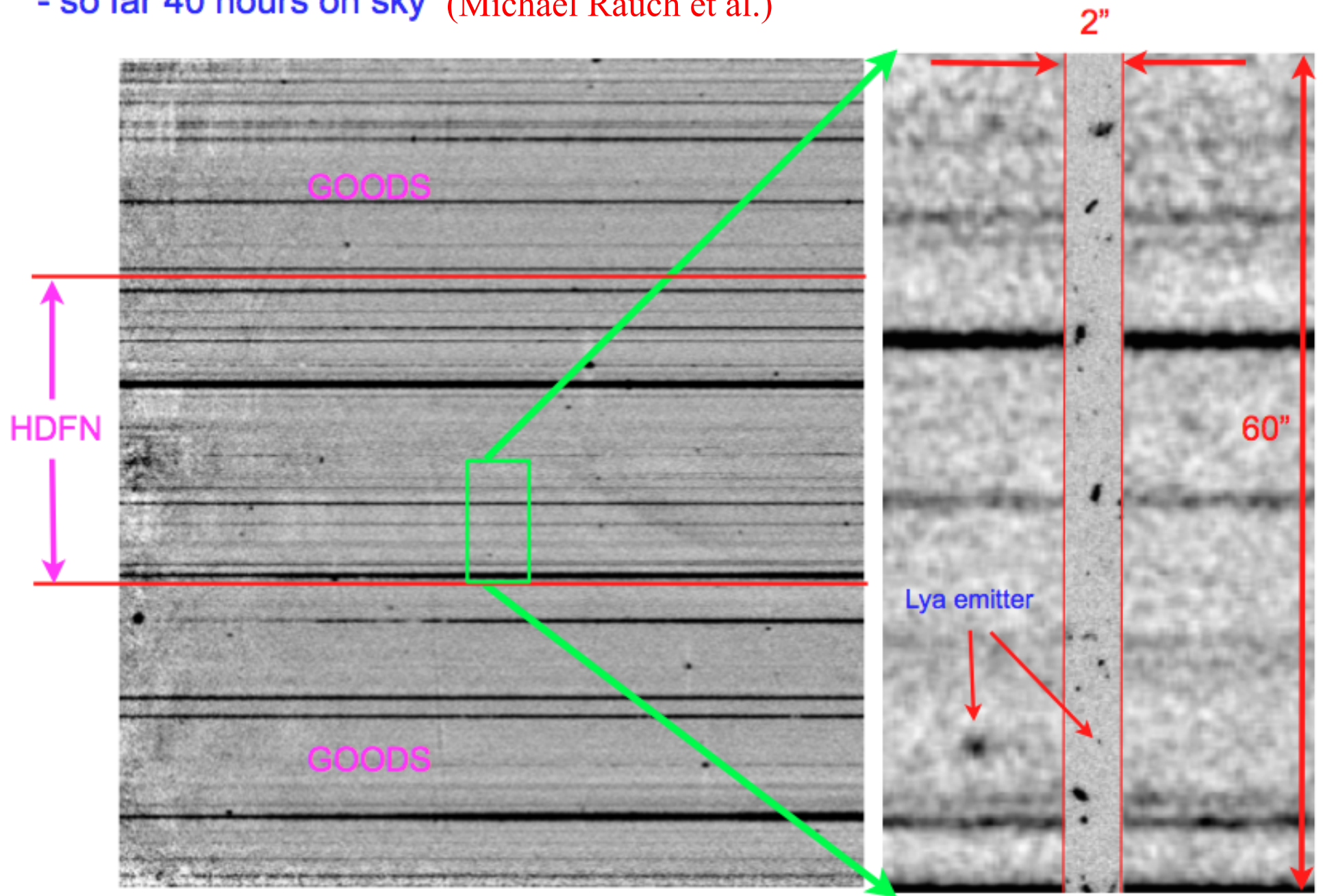
92 hours in a blank field with FORS!

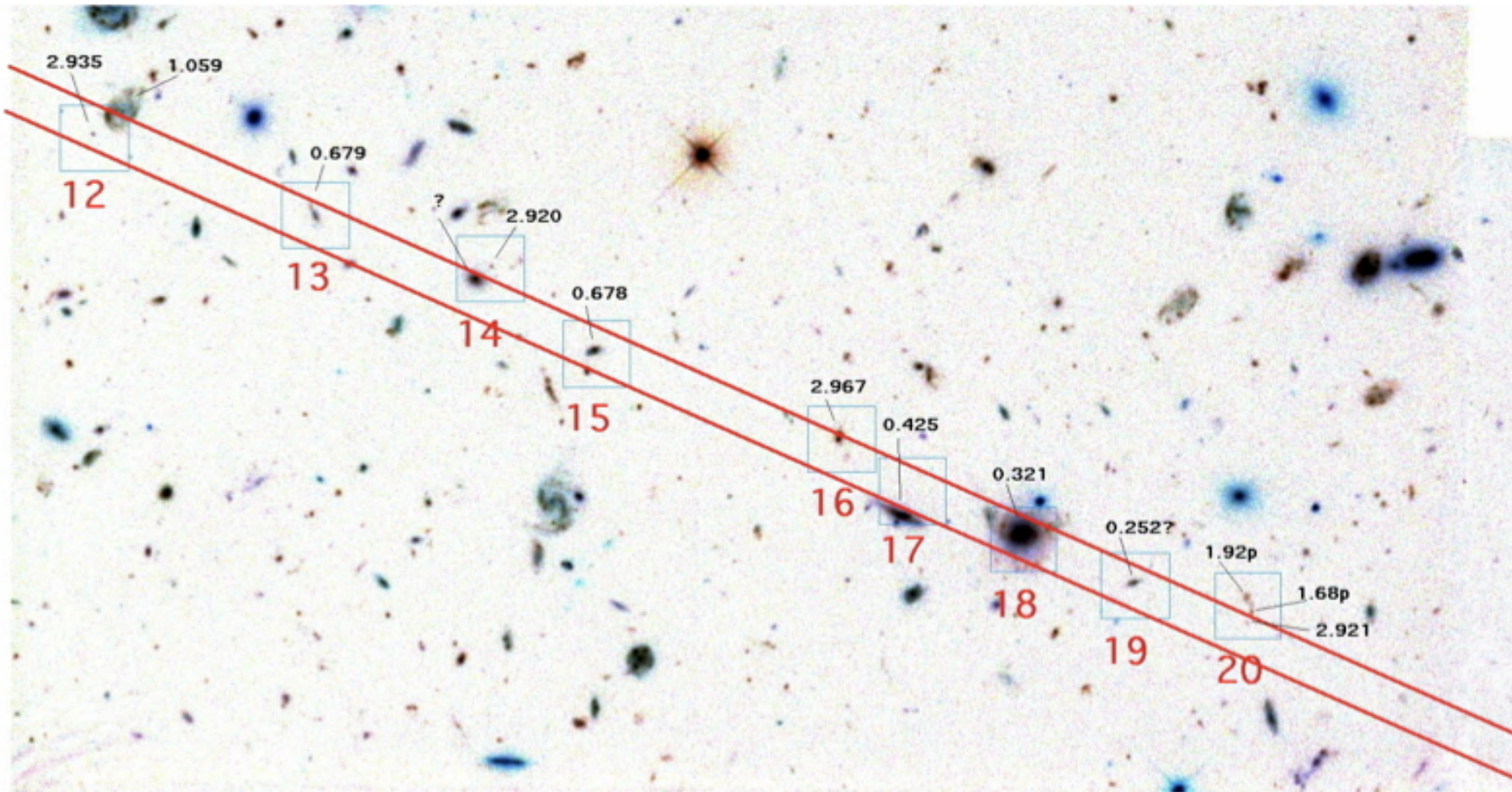
dN/dz comparable to that of DLAs  
 $\rightarrow$  DLA host galaxies

space density:  $3 \times 10^{-2} h_{70}^3 \text{ Mpc}^{-3}$

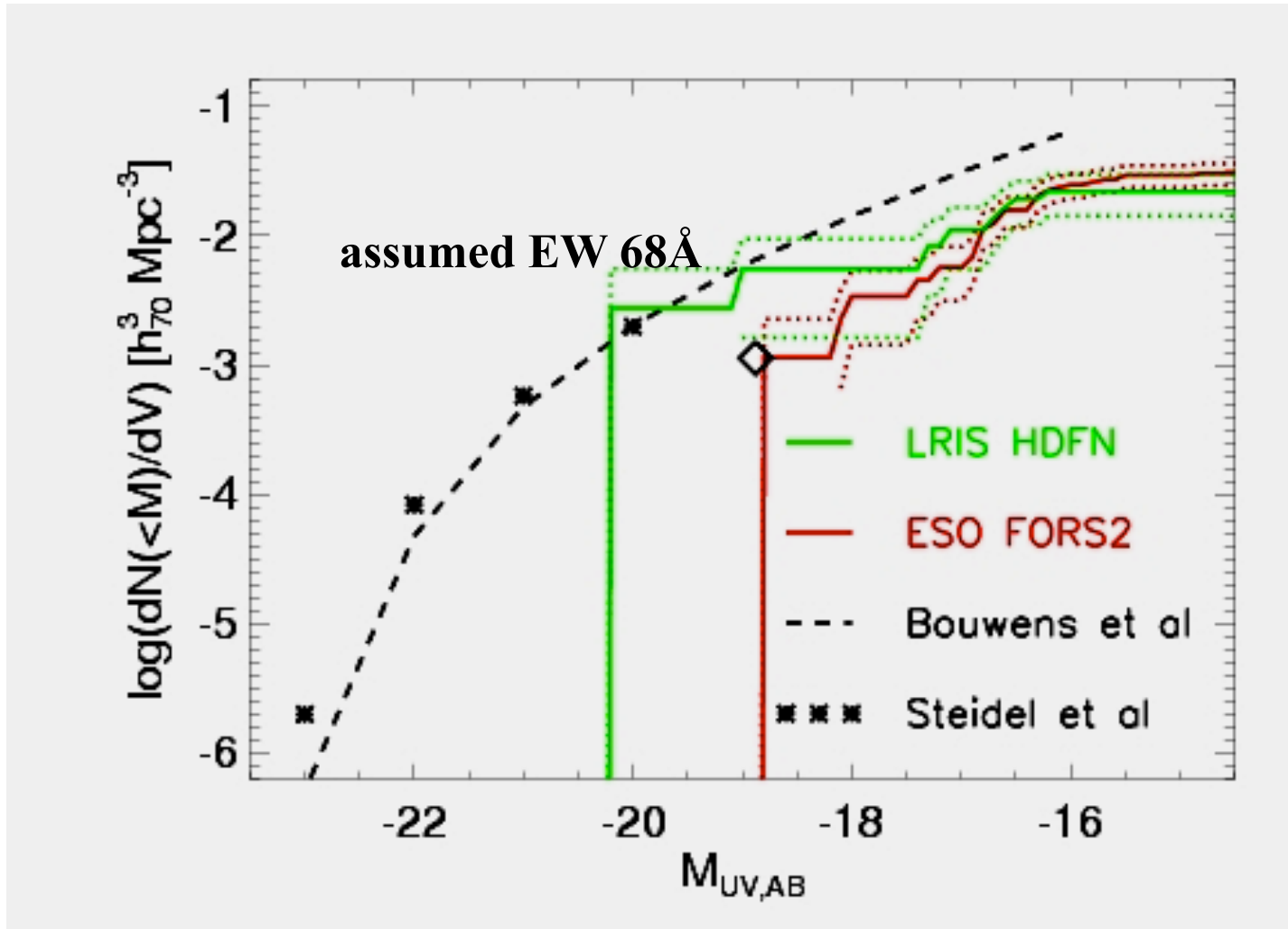


- Keck LRIS LS spectroscopy of the Hubble Deep Field North
- so far 40 hours on sky (Michael Rauch et al.)





**Continuum counterparts are compact and indeed from the very faint end of the LBG population.**



**The emitters are as abundant but fainter as B dropouts in the HUDF!**

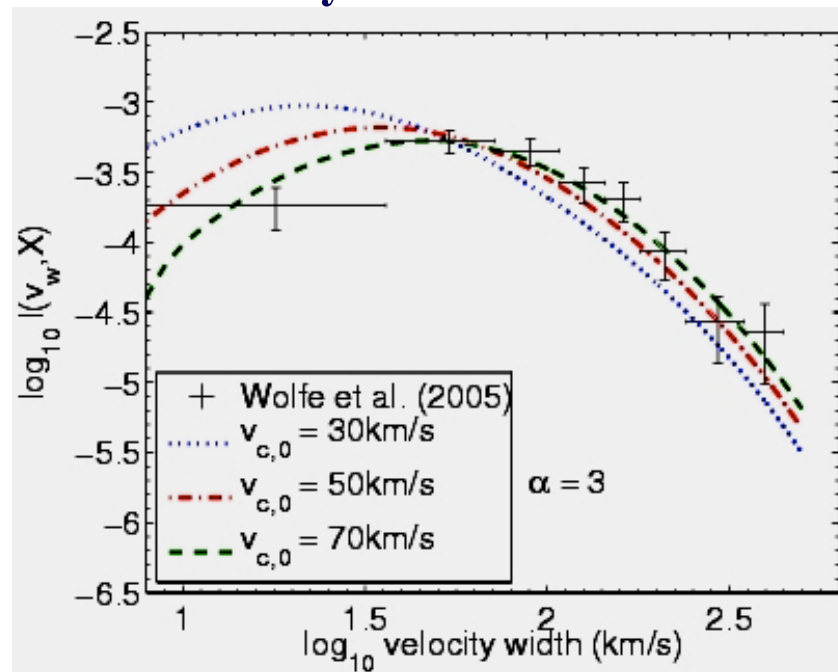


# A joint model for the emission and absorption properties of damped Ly $\alpha$ absorption systems

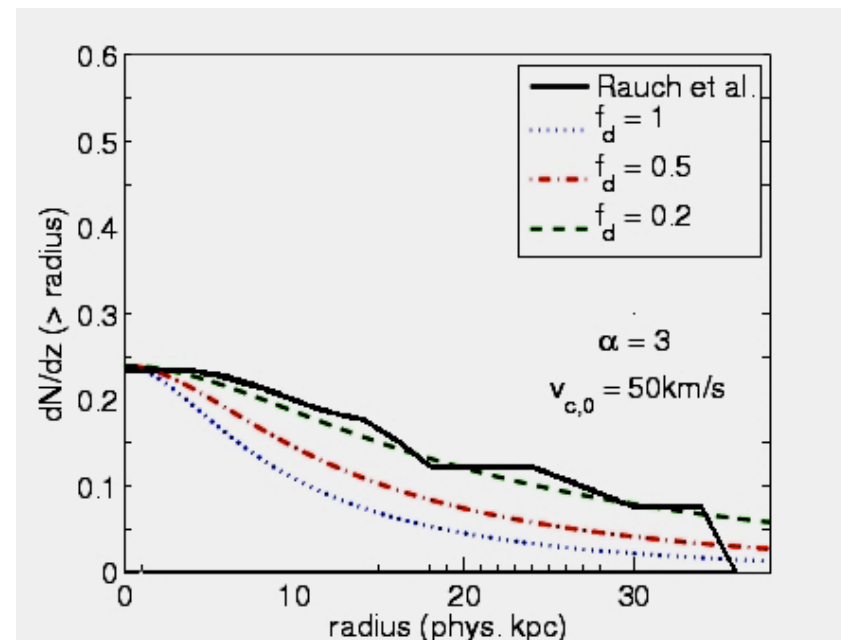
Luke A. Barnes\* and Martin G. Haehnelt  
*Institute of Astronomy, Madingley Road, Cambridge, CB3 0HA*



## velocity width distribution



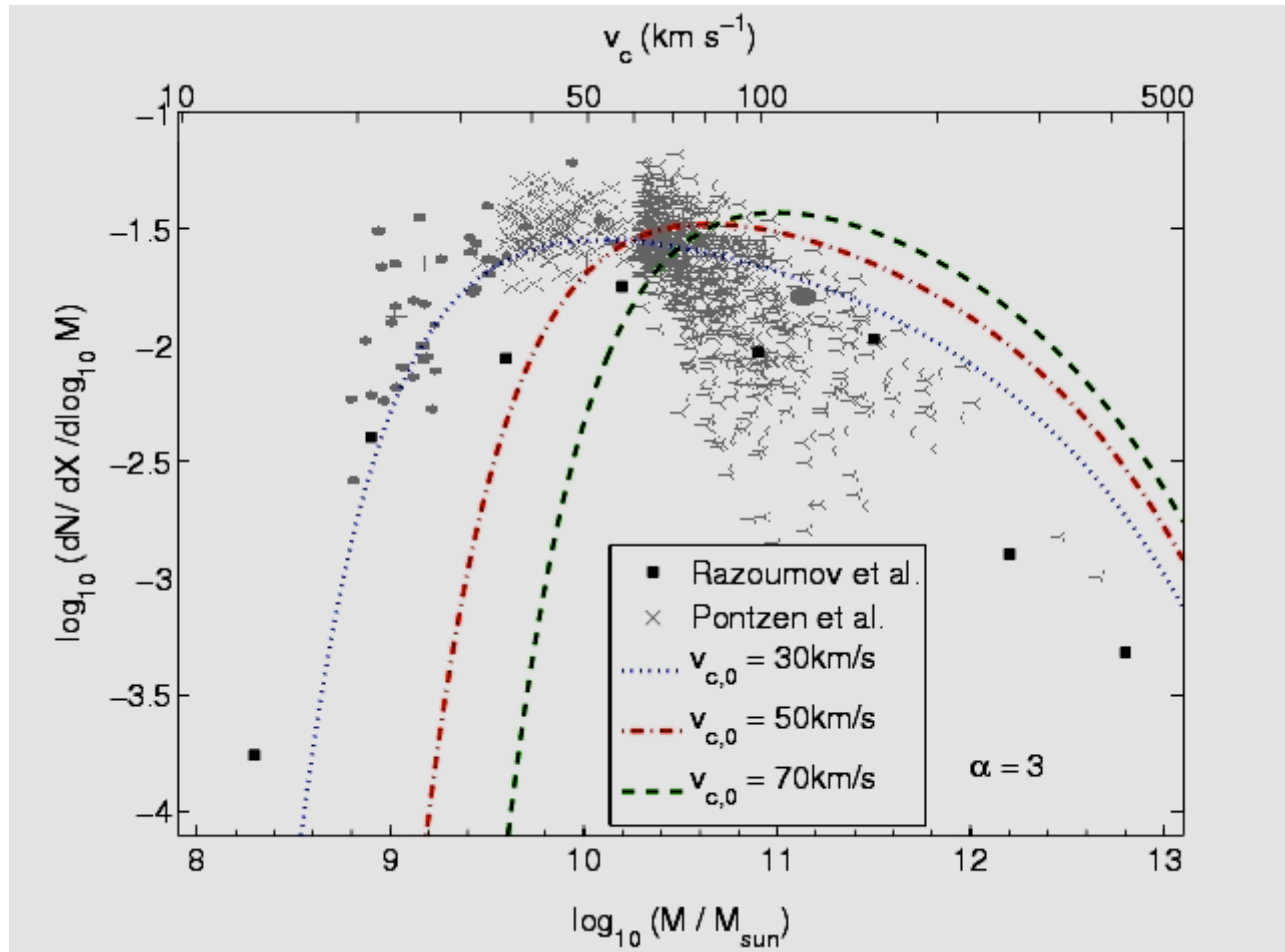
## cumulative incidence rate



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# The differential incidence rate of DLAS

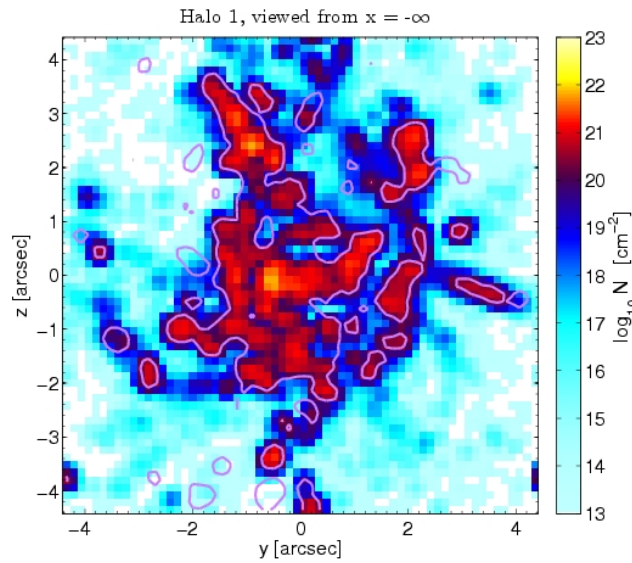


Barnes & Haehnelt 2008

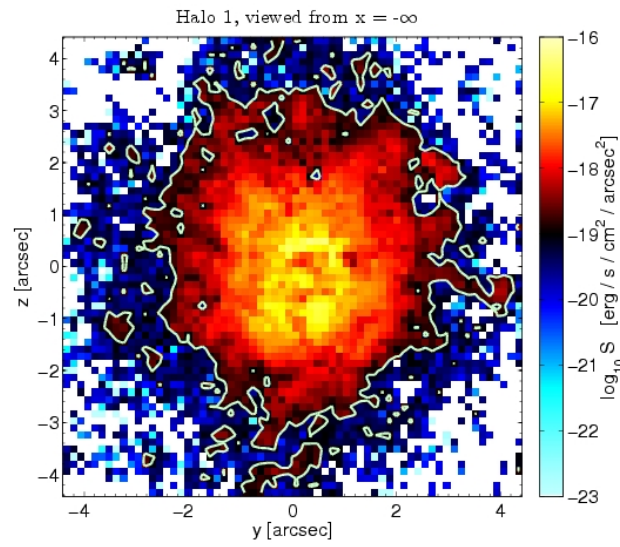
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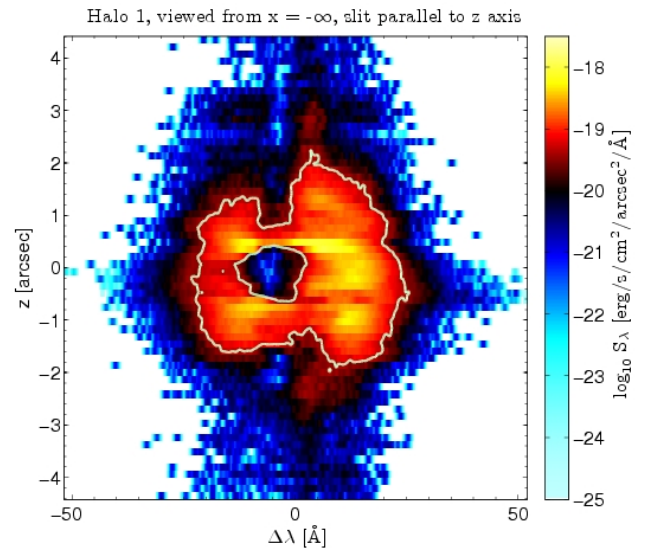
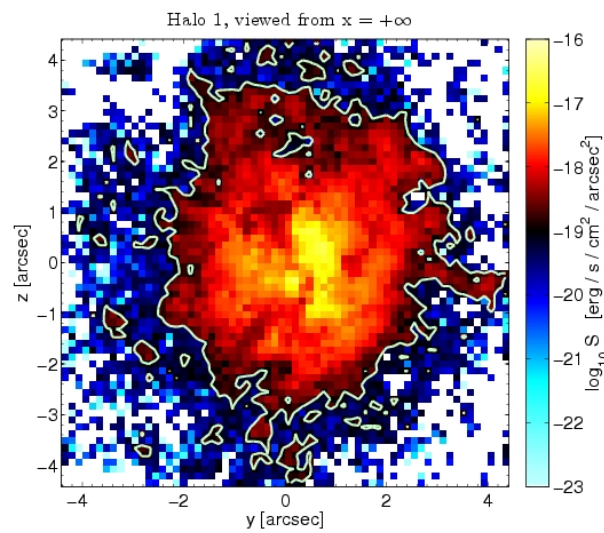
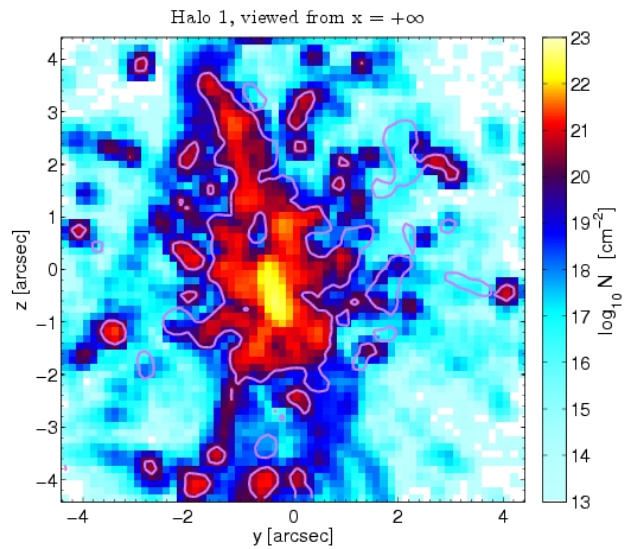
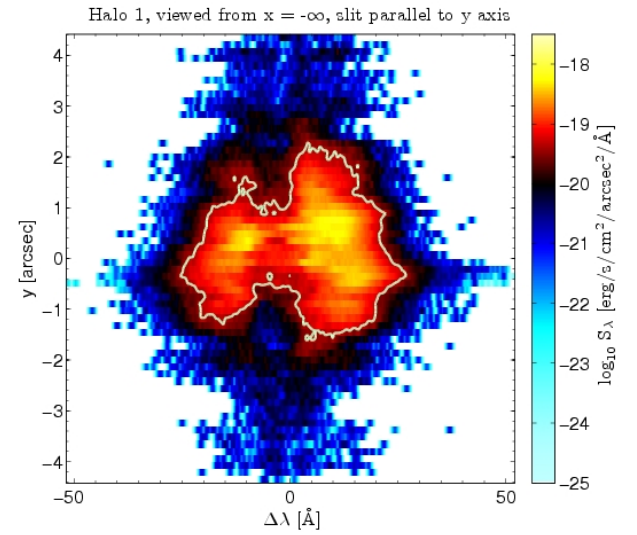
$N_{\text{HI}}$



$\text{Ly}\alpha$  image



2D spectrum

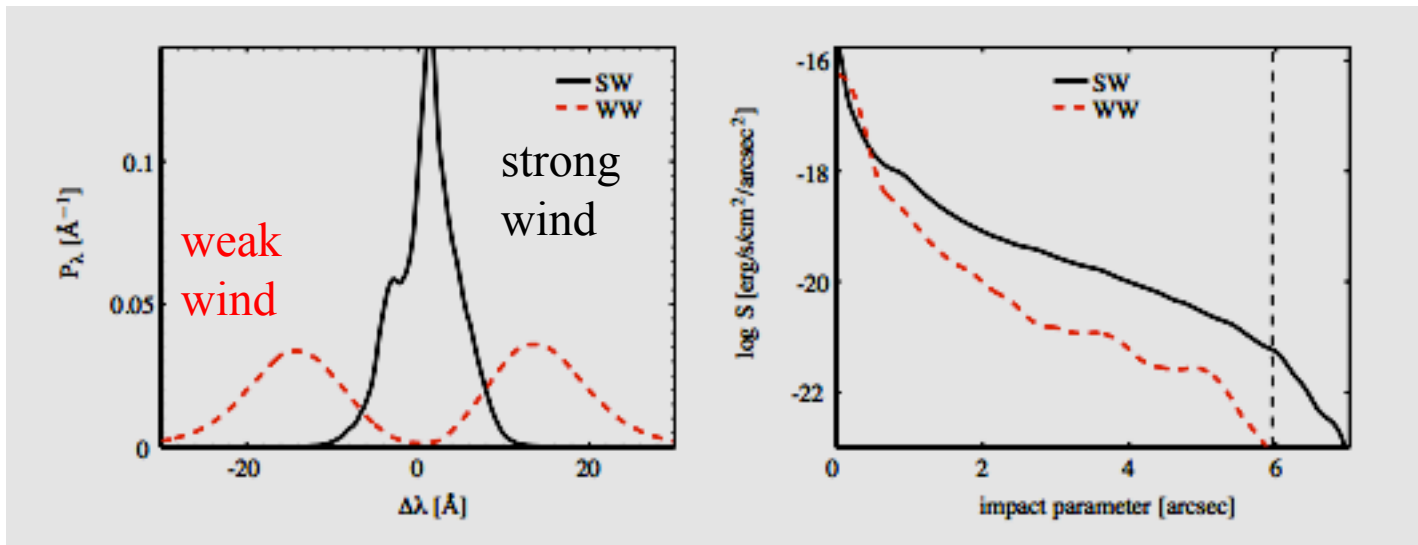


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Barnes, PhD thesis



# Probing outflows/inflows



Barnes et al. 2011

**The Lyman-alpha emission depends strongly on the strength of a galactic wind.**

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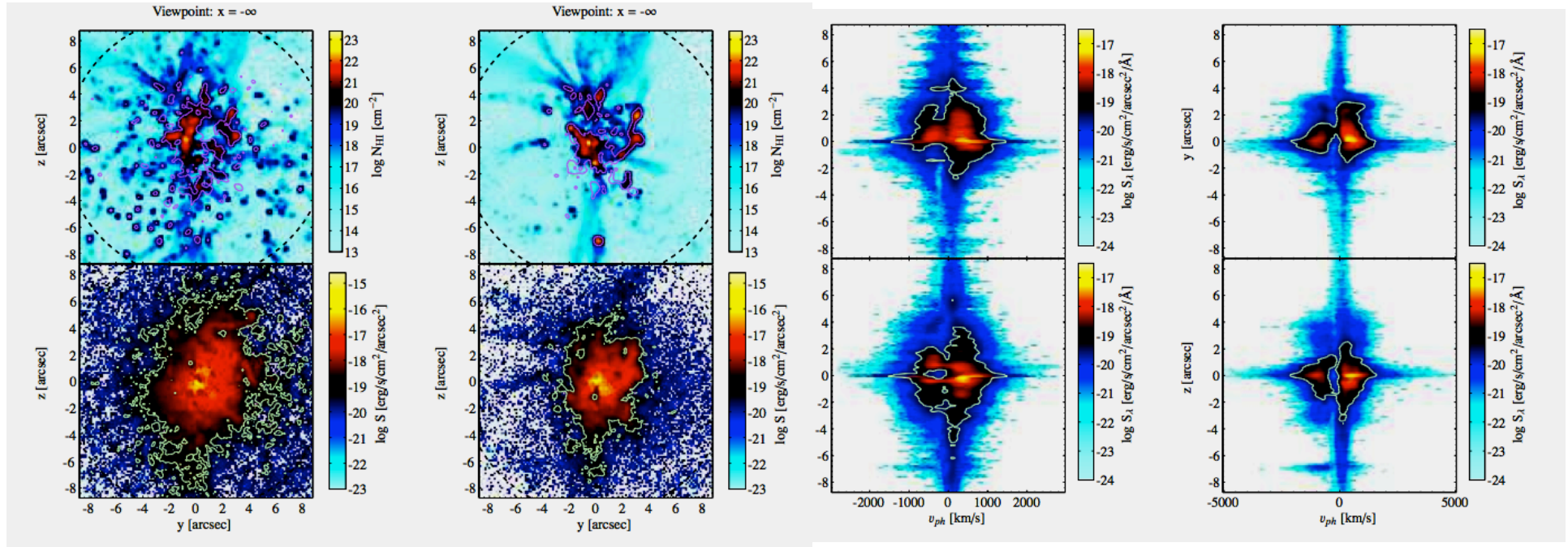


strong winds

weak winds

strong winds

weak winds



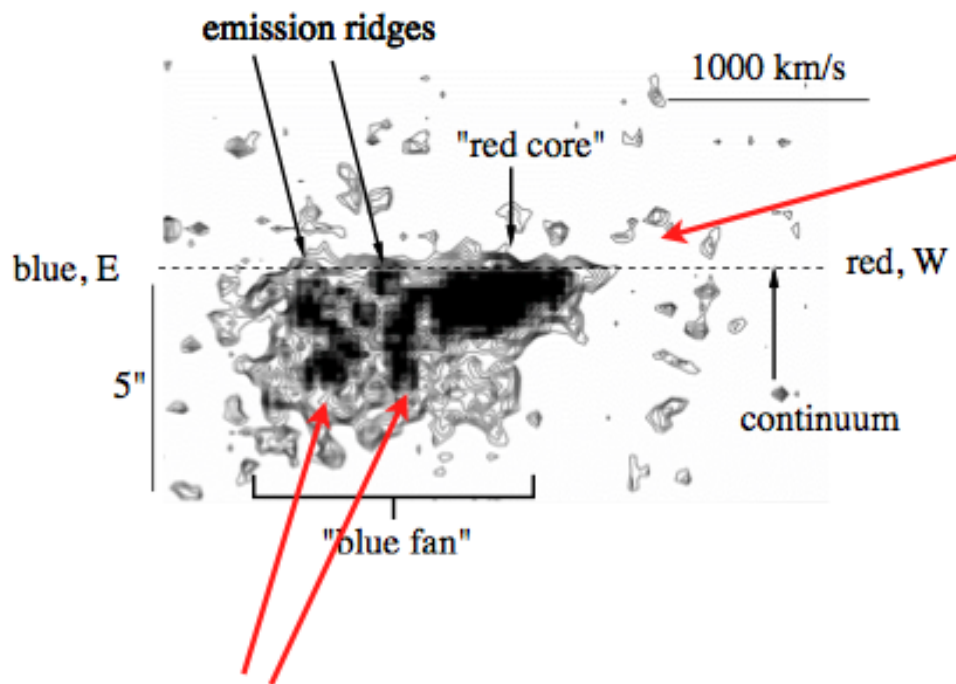
Barnes et al. 2011

IFU spectroscopy should become an excellent tool to discriminate between the rather crude implementations of galactic winds in numerical simulations.

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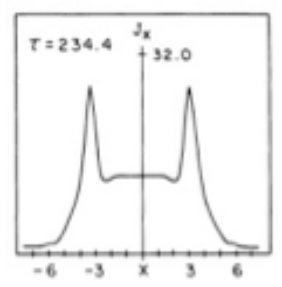




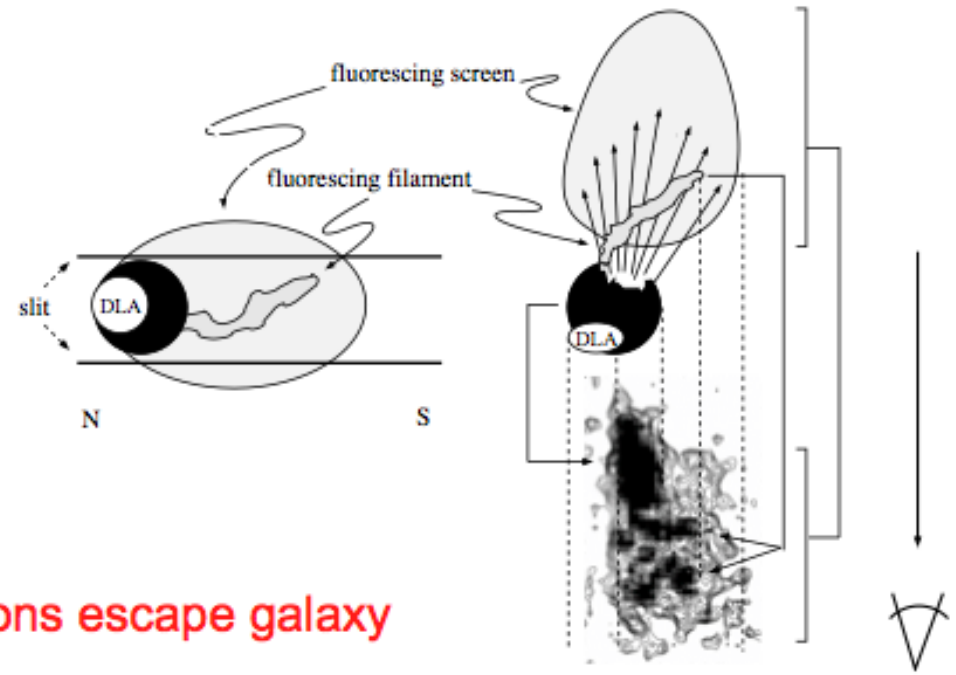


DLA sits in front of blueshifted, extragalactic gas --> Infall

infalling filament fluoresces in Ly $\alpha$  (double-humped profile!)



Urbaniak & Wolfe 81



~50 % of ionizing photons escape galaxy to hit blue infalling gas

## Summary

- **very faint mostly spatially extended Ly $\alpha$  emitters detected at  $2.67 < z < 3.75$  in 92h deep FORS2 and 40h Keck exposure**
  - **very steep faint end of Ly $\alpha$  luminosity function**
  - **inferred incidence rate corresponds to that of optically thick QSO absorption systems**
  - **probably host galaxy population of DLAS and LLS detected ( $M_{\text{tot}} \sim 10^{10} - 10^{11} M_{\odot}$ ,  $V_{\text{vir}} \sim 50 - 150 \text{ km/s}$ ); mainly powered by (spatially compact) star formation**
  - **building block of Milky-Way type galaxies**
- **Ly $\alpha$  most likely due to central star formation**
- **Ly $\alpha$  scatters in frequency and real space. Spatial extent and spectral shape is very sensitive to details of the kinematics of the gas.**