

KiDS Survey Status

Jelte de Jong, Leiden Observatory

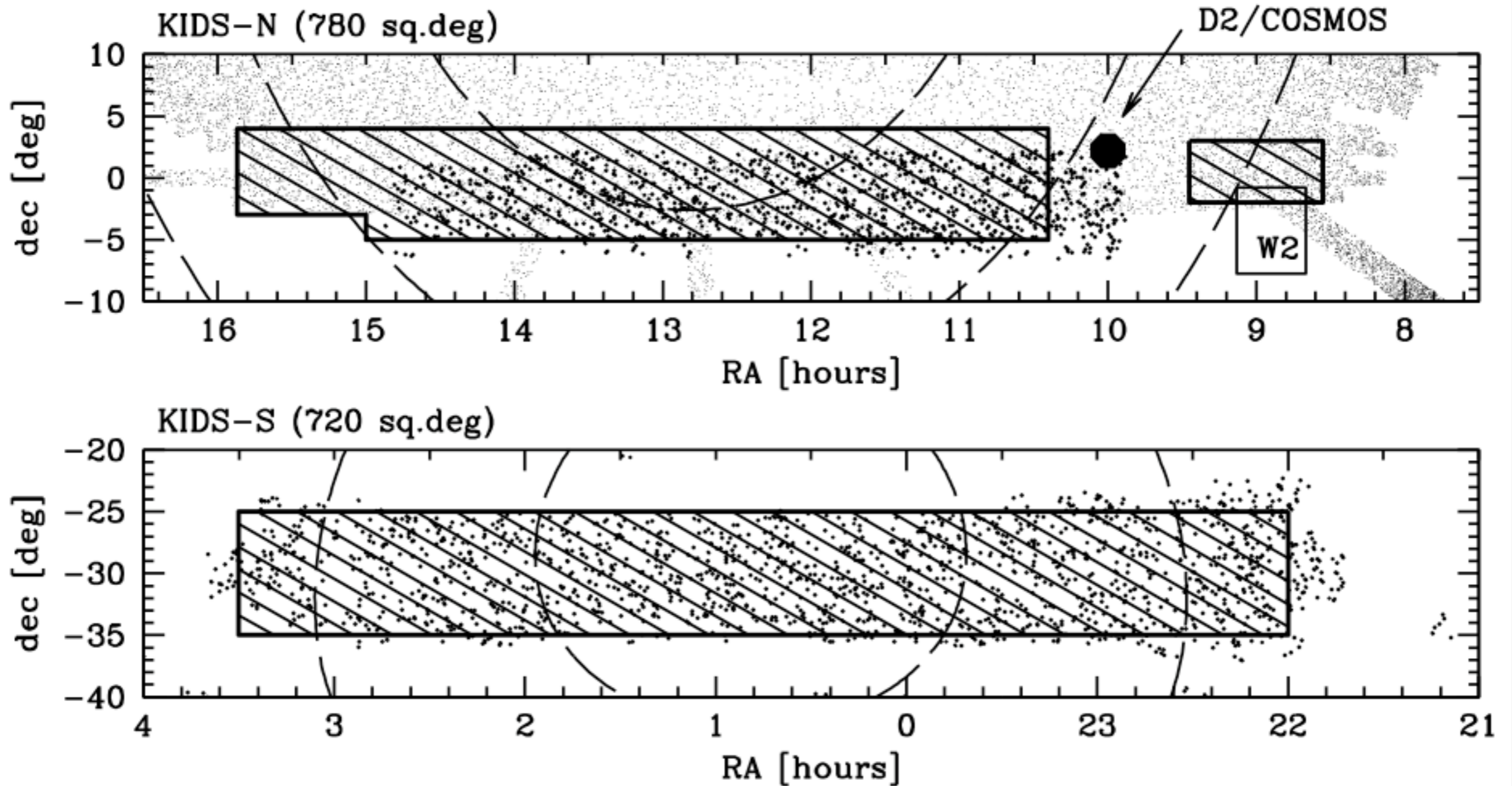
KiDS Survey Status

- Survey overview
- Observational planning
- Data in hand
- Data processing
- VST calibration plan

Survey overview

- deep survey: 1500 sq.deg. in u, g, r, i
- two fields: KiDS-S ($-35^\circ < \delta < -25^\circ$) and KiDS-N ($-5^\circ < \delta < +5^\circ$)
- overlapping with VIKING
- ~~shallow photometric pass for phot. cal.~~
- repeat pass in g, providing 2-year baseline

Survey overview



Survey overview

Fields chosen such that:

- observations can be done year-round
- 2dF and SDSS (KiDS-N) spectroscopy available for several 100,000 galaxies

Survey overview

Observing conditions and exposure times such that:

- balanced over astro-climate conditions
- median galaxy redshift ~ 0.8
- best seeing for r-band, the weak lensing band

Filter	Exp. time	Median seeing	Moon	10σ AB limit
u	1000 s	1.0''	Dark	24.8
g	900 s	0.8''	Dark	25.4
r	1800 s	0.6''	Dark	25.2
i	1200 s	0.7'' (broad)	Bright	24.2

Survey overview

- Survey 'tiles' follow the Astro-WISE platesystem
- 1 OB per filter; per tile 1 group container
- OBs set-up as follows:

Filter	Exp. time	Dither	Seeing	Airmass	Moon
u	4x250 s	diag	<1.1''	<1.2	FLI<0.4; dist>90°
g	5x180 s	diag	<0.9''	<1.6	FLI<0.4; dist>90°
r	5x360 s	diag	<0.7''	<1.2	FLI<0.4; dist>90°
i	5x240 s	diag	<1.1''	<2.0	any FLI ; dist>60°

Observational planning

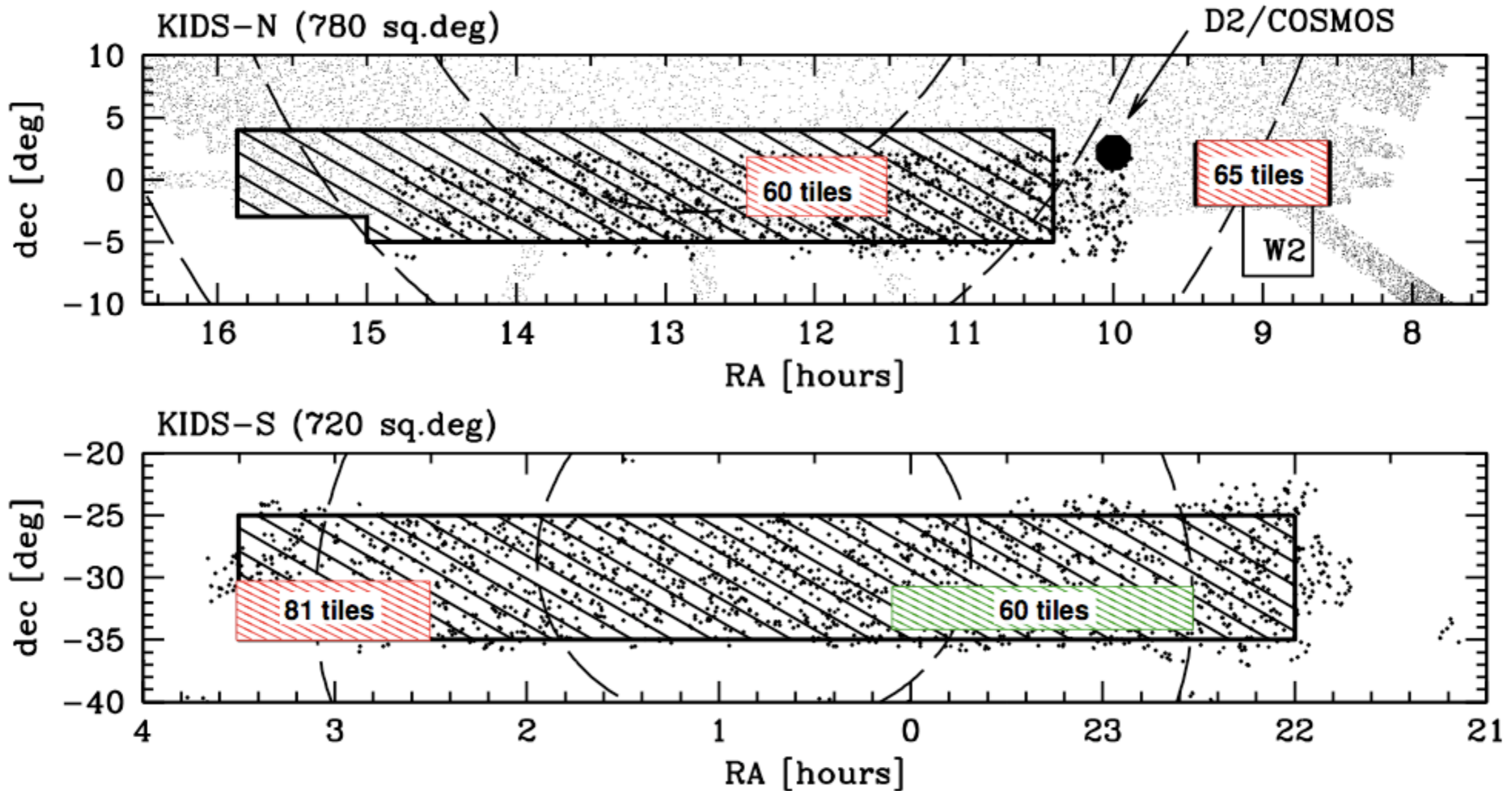
Goal 1: provide full 9-band photometry a.s.a.p.:

- Highest priority to areas already covered by VIKING
- Grouping of OB's per tile: first observe tiles that have been partially observed already

Goal 2: complete survey (excl. repeat) in 3 years:

- Aim for 250 sq.degrees per semester
- Ask time for 400 sq.degrees per semester

Observational planning



KiDS observing plan for P88

Time awarded: 420 hours
Tiles submitted: 210 (exec. time 419.7 hr)
Not completed EST OBs are also carried over to P88

Targeted areas coincide with existing VIKING data. Distribution over RA guarantees observable tiles during complete semester.

Data in hand

Early Science Time (Sept - 15 Oct):

- 240 OBs (60 tiles) submitted
- 45 OBs “completed” and ingested into Astro-WISE: 8 u, 8 g, 29 i, 0(!) r
- 23 OK; 16 NOK (11 due to dead CCD);
6 TBD

Also handful of OBs pre-EST and 73 OBs P88

Data processing

KiDS data processed with Astro-WISE

Astro-WISE:

- was designed for VST data
- data-centric survey handling system, based on a fixed data model, in which all data objects are instantiations of a data class
- uses a federated database, allowing people at different locations to work together on data processing and quality control
- stores data lineage of all objects

Data processing

[Astro-WISE](#)
[Astro-WISE](#)
[Homepage](#)

Target Processor

Contact
[Willem-Jan Vriend](#)

DB User
[awjdejong](#)

Help
[Getting Started](#)

Project

KIDS

Instrument

OMEGACAM

State

1. Preselect Target
2. Specify Target
3. Select Target(s)
4. **Process or Query**

Options

- [Preferences](#)
- [Process Parameters *](#)
- [Upload Code](#)
- [Job overview](#)

Jobs for awjdejong

DPU	ID	Actions	User	Status	Submitted	Details
dpu.hpc.rug.astro-wise.org	9130	view stop	awjdejong	SENDDATA	Fri Dec 2 14:58:09 2011	Q/R/F/E/A/S/U 98/0/0/0/0/0

[Refresh](#) [Cancel](#)

Browse previous processing logs of [awjdejong](#)

page generated 2011-12-02 14:58:14.129548
generation time 0:00:08.777072
For optimal experience use [firefox](#) browser

empowered by



1 2011-08-26 06:22:50 [Sci-EHELMICH-OMEGACAM-----OCAM_u_SDSS---Coadd---Sci-55862.5417669-deff3fbeb3eft](#)

Quality of RawScienceFrame:

OMEGACAM.2011-09-06T08:19:11.265_30.fits

[AstroWISE DBView](#) [CaITS Process](#)

[see 2 previous comments](#)

DBname: [awidejong](#) project: KIDS

is_valid =

Processing Details

creation_date	2011-09-06 15:46:13
is_valid	1
quality_flags	0
Privileges	3

Image Statistics Details

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median	+8.670e+02
stdev	+1.276e+03
min	+2.580e+02
max	+6.554e+04

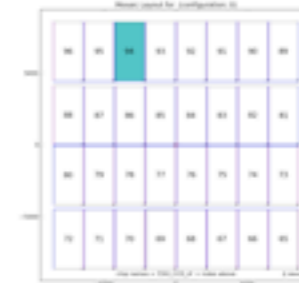
Photometry Details

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zp_error	0.106
zp_origin	derived
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extinction	0.046
ext_error	0.003

Observational Details

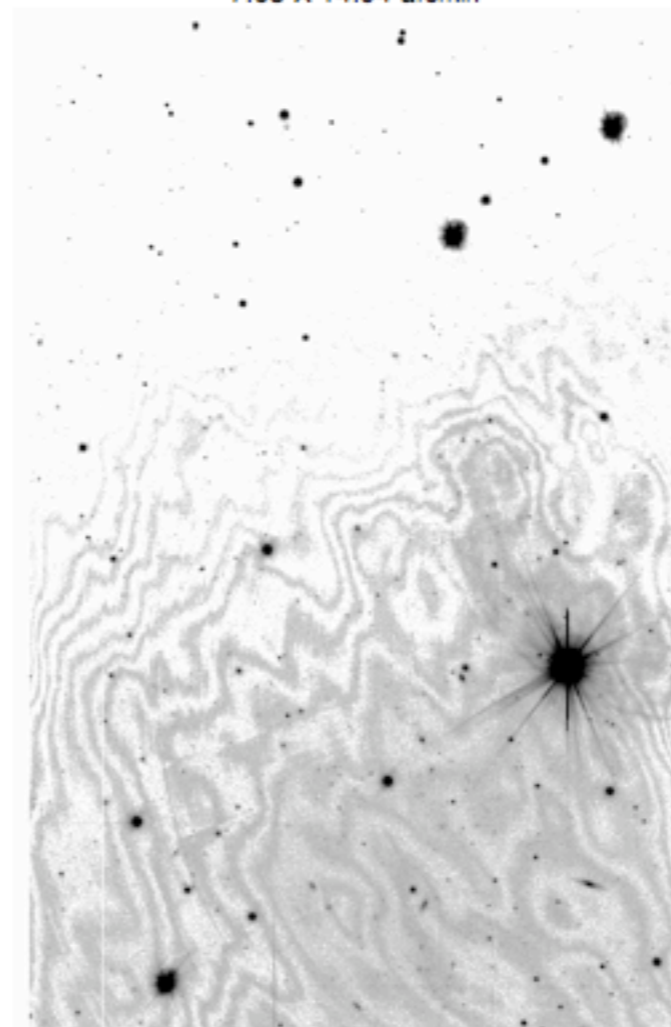
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R.A.	22:34:25.3322	AIRMEND	1.696
Dec.	-33:09:39.6360	Filter	OCAM_i_SDSS
		mag_id	SloanI

Chip ESO_CCD_#94 of Instrument
OMEGACAM



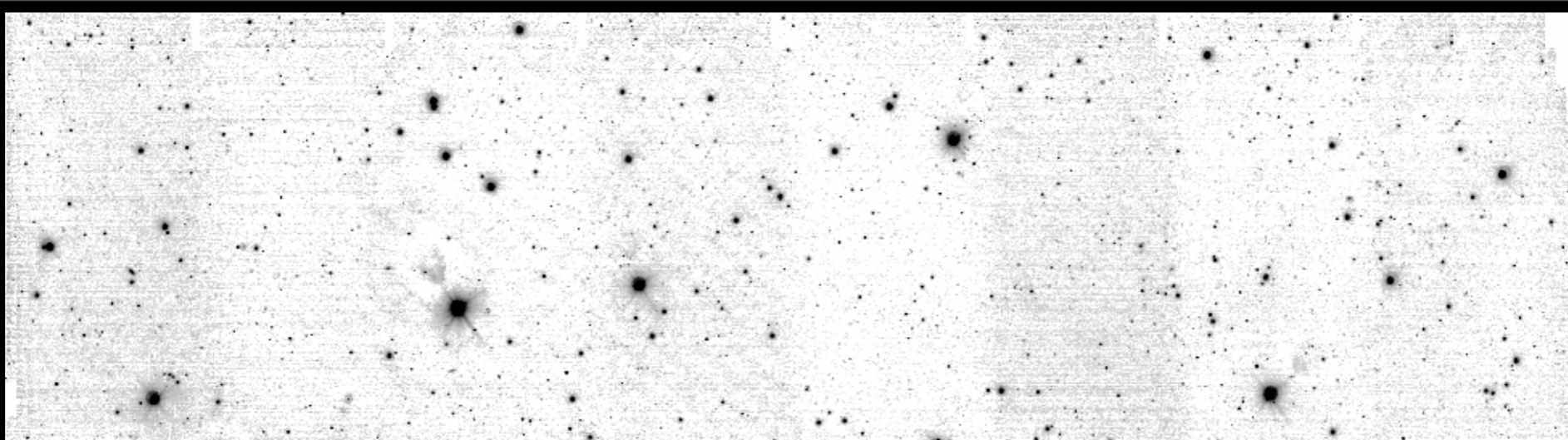
RawScienceFrame

2144 X 4200 pixel
7.63 X 14.94 arcmin

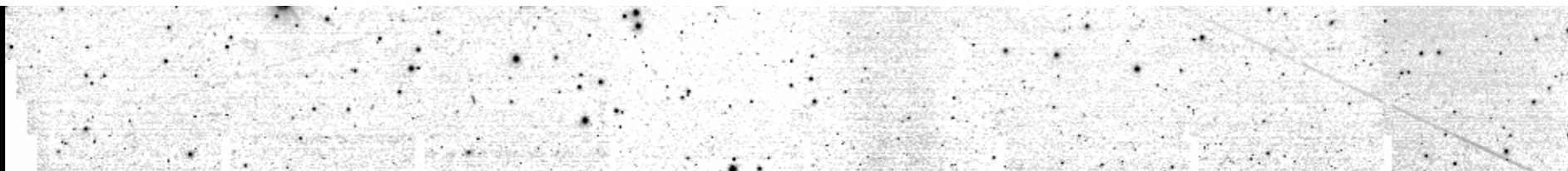
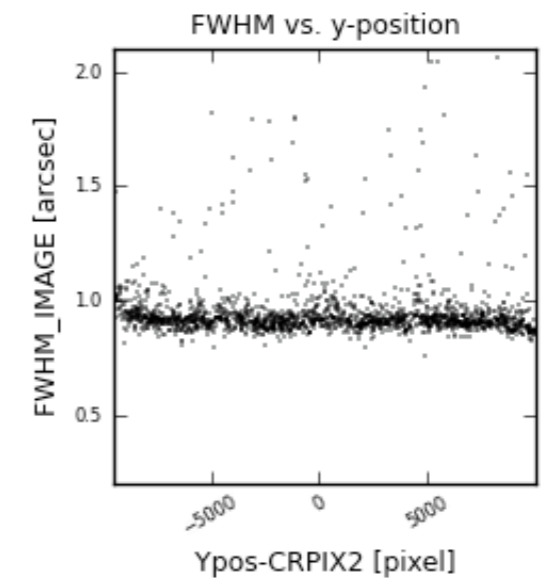
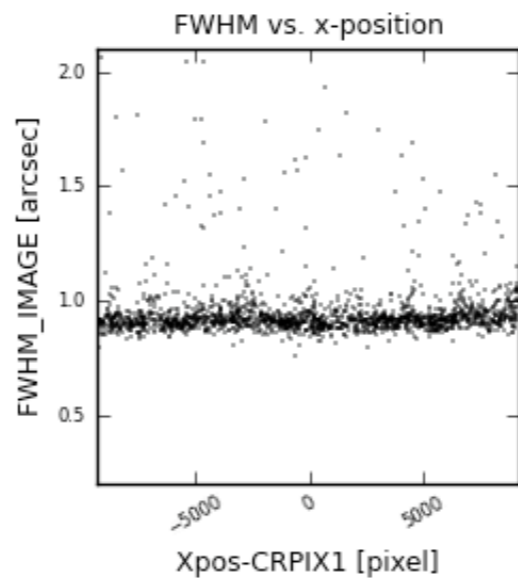
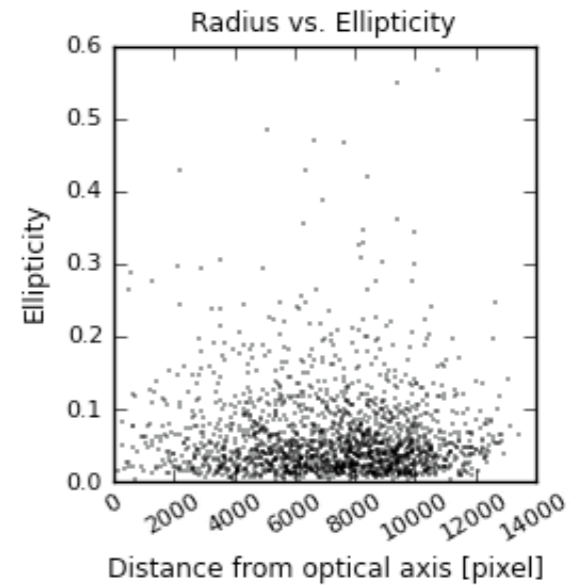
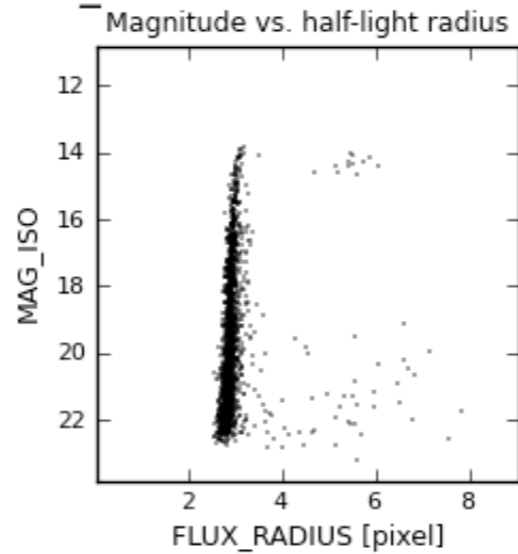
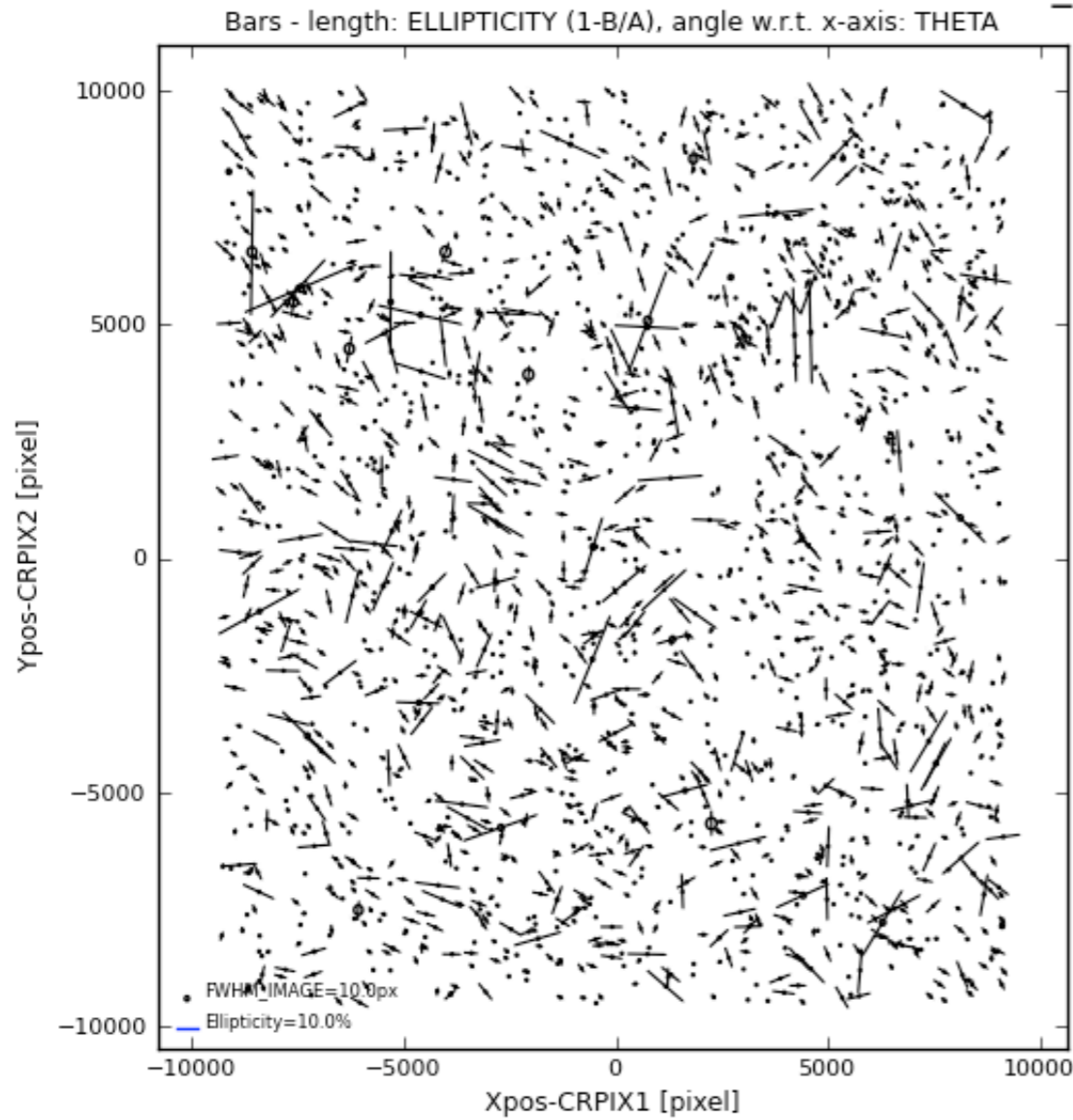


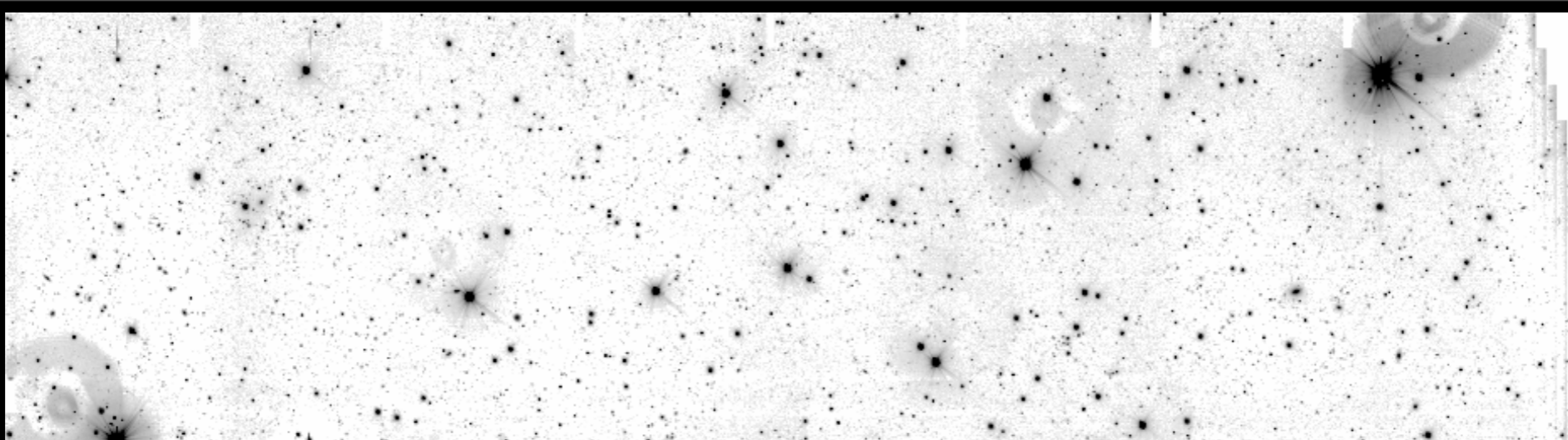
Data processing

- Per chip calibration
- Master flat = dome (small scale) + twilight (large scale)
- Illumination correction tied to a master flat
- Background subtraction upon coadding

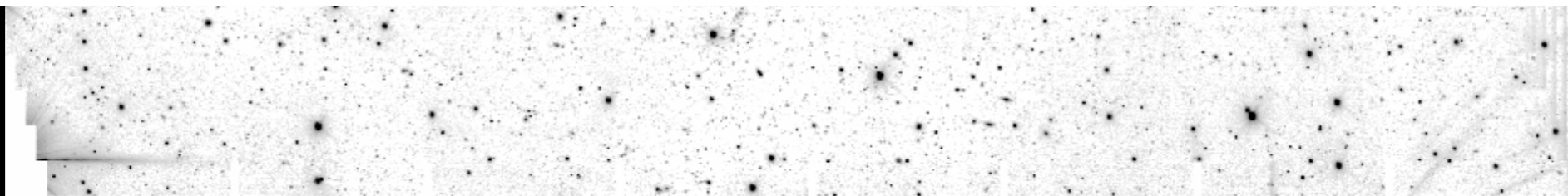
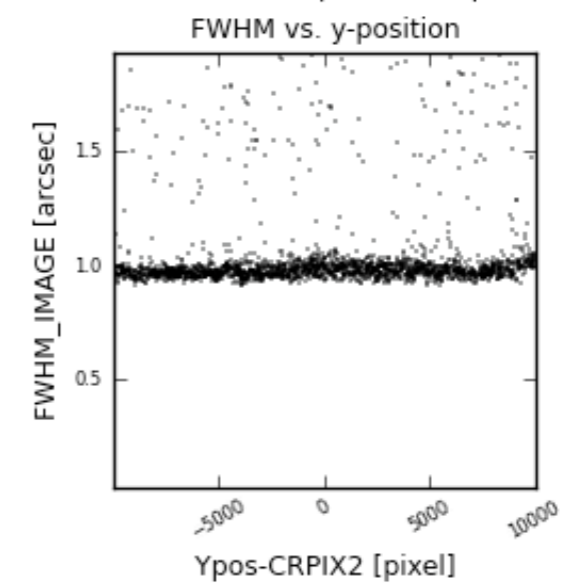
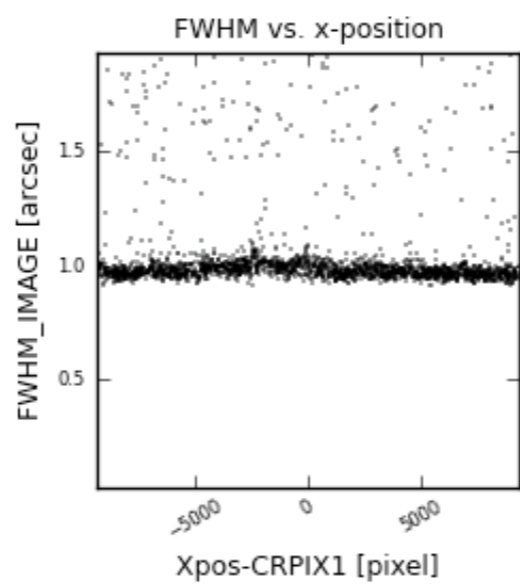
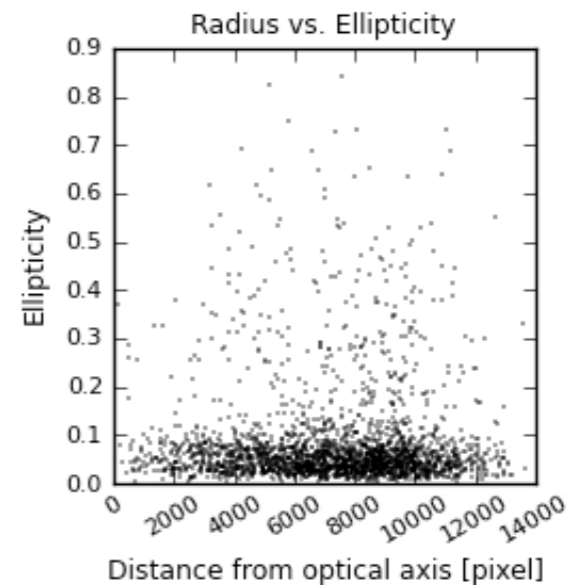
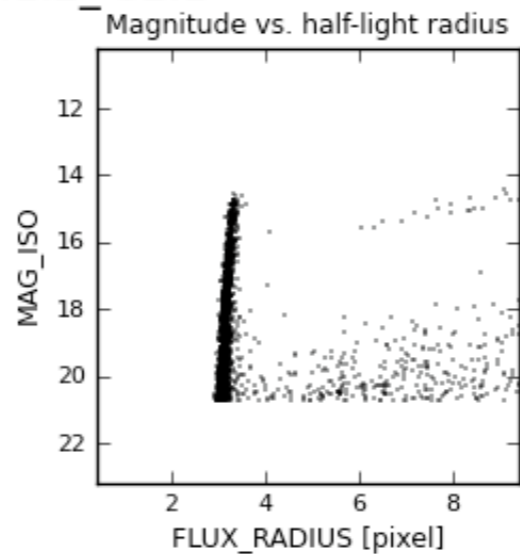
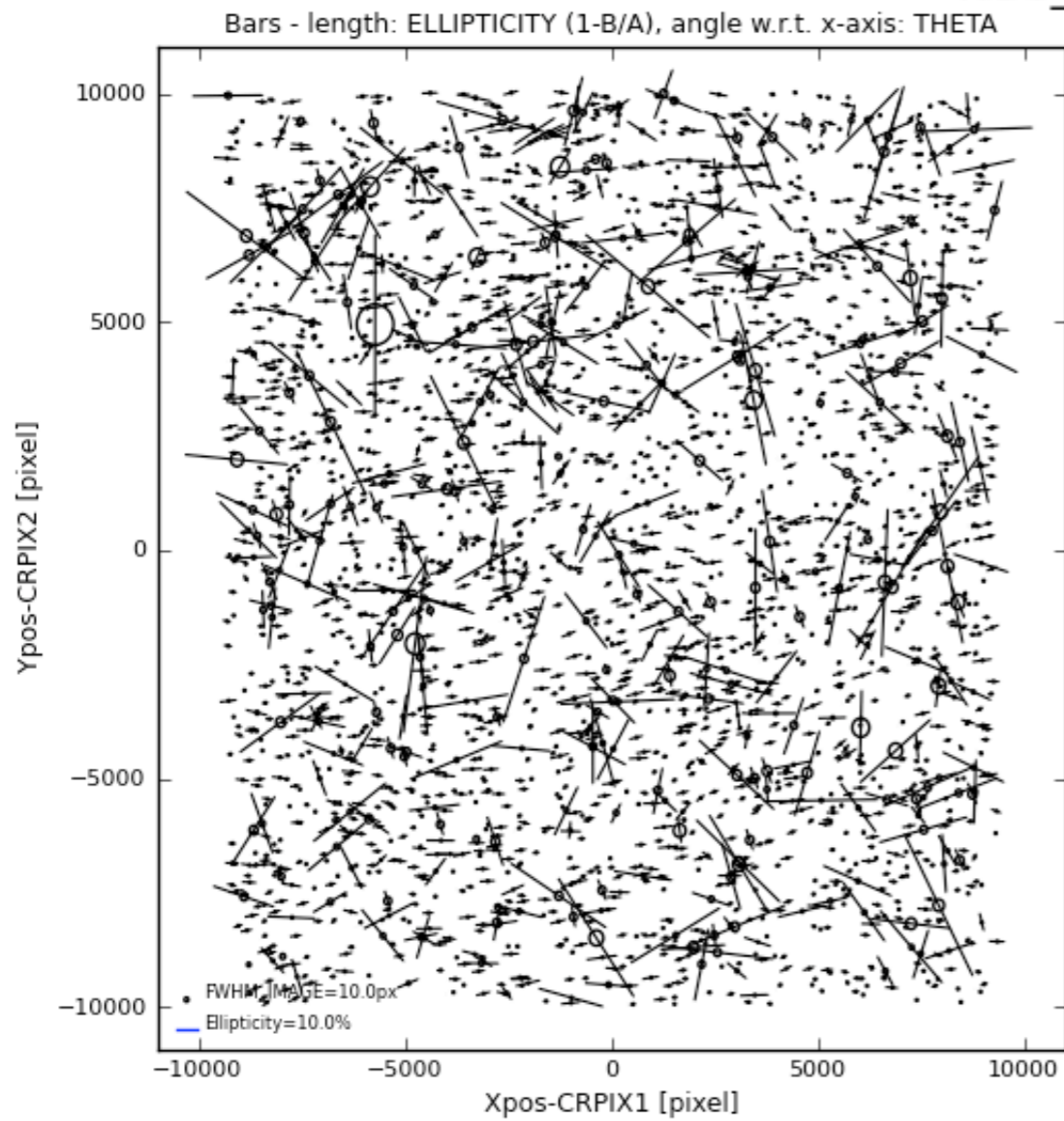


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KIDS_338.6_-33.1

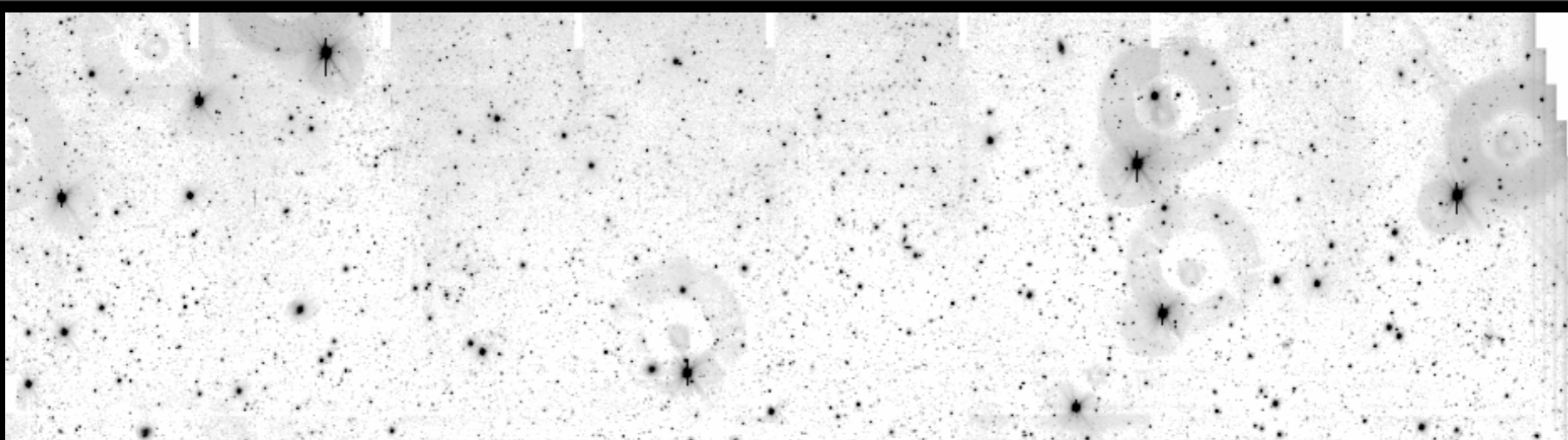




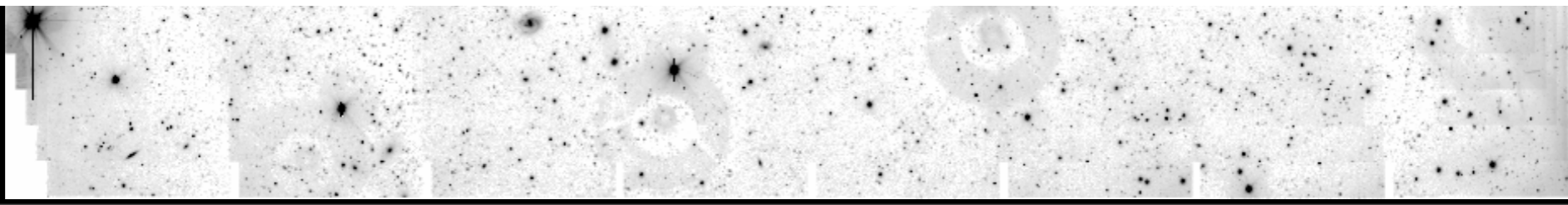
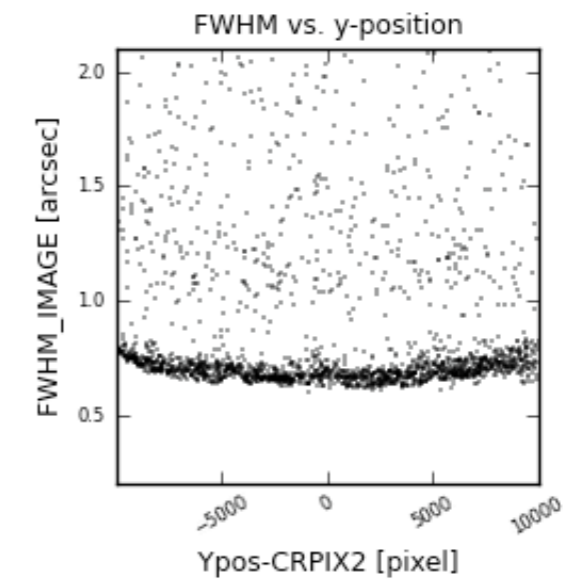
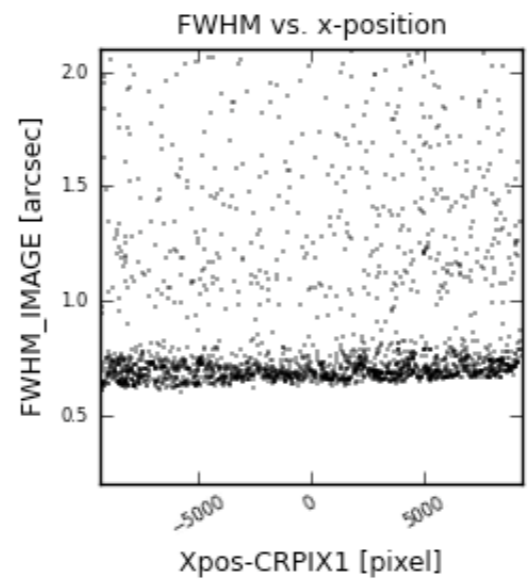
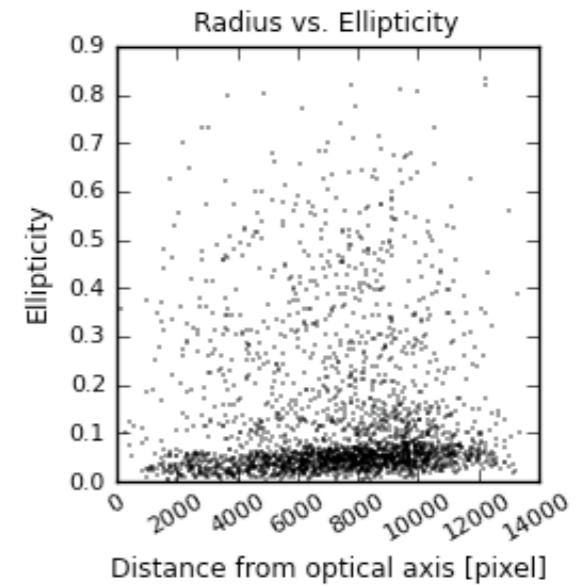
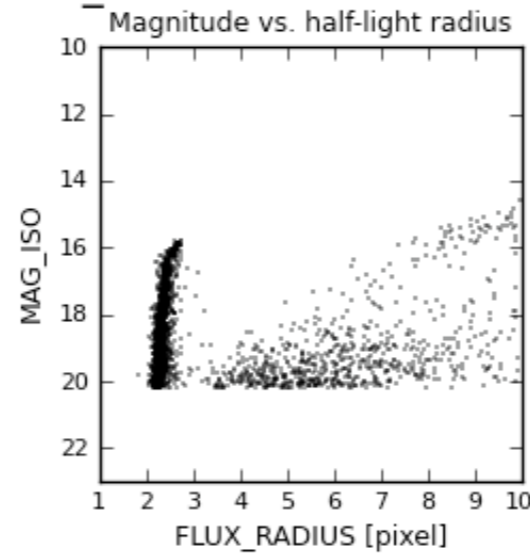
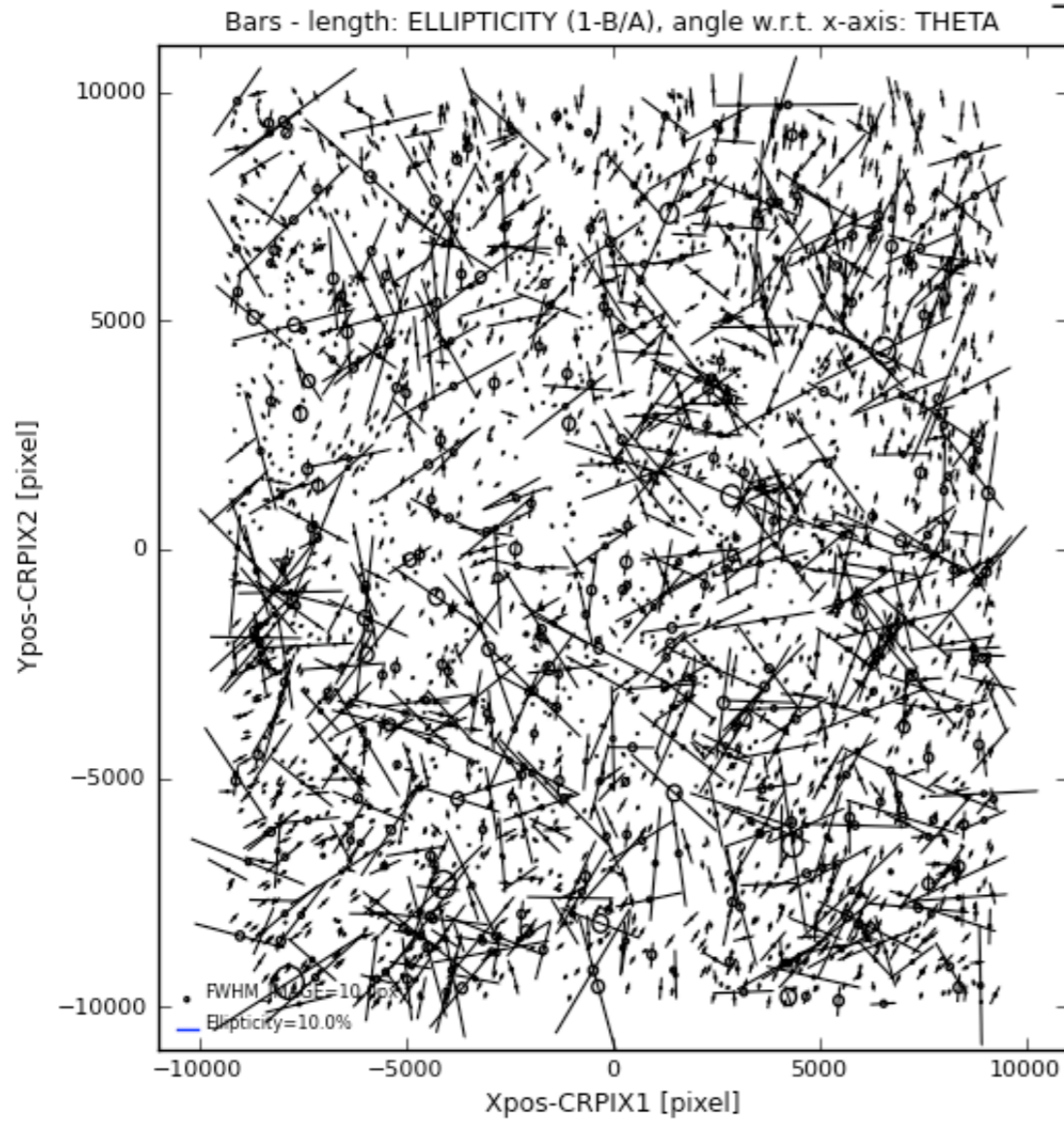
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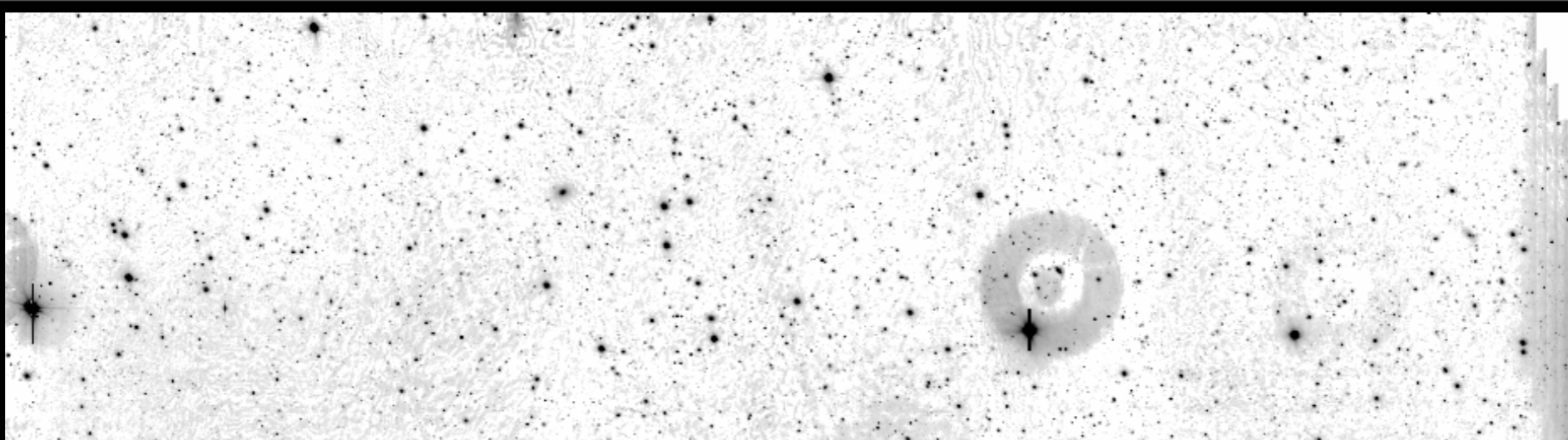


r

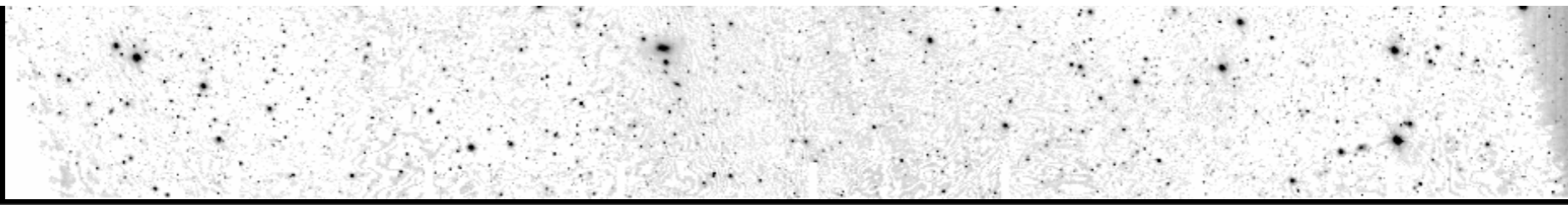
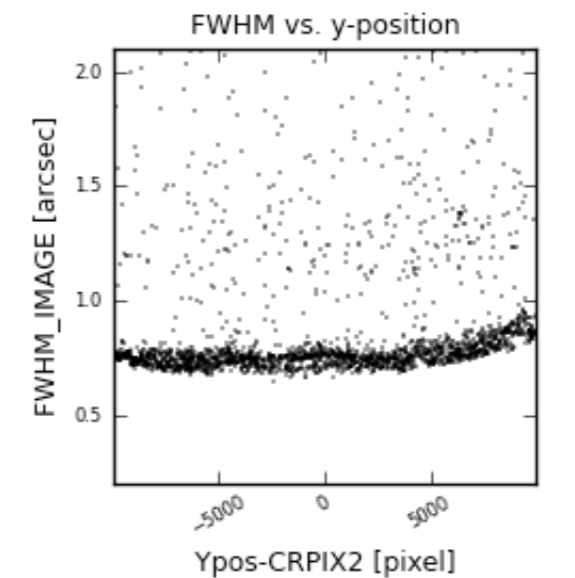
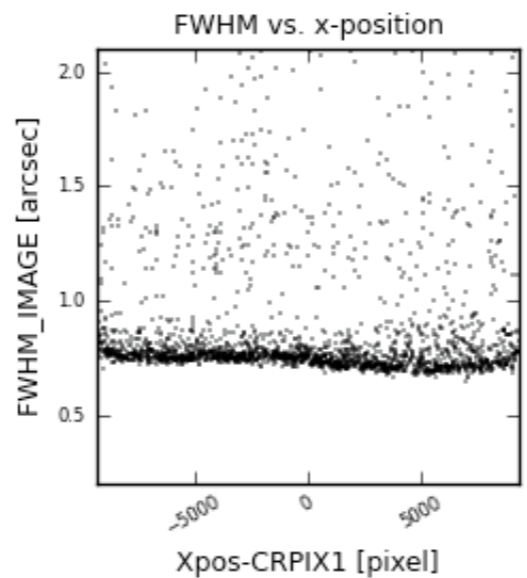
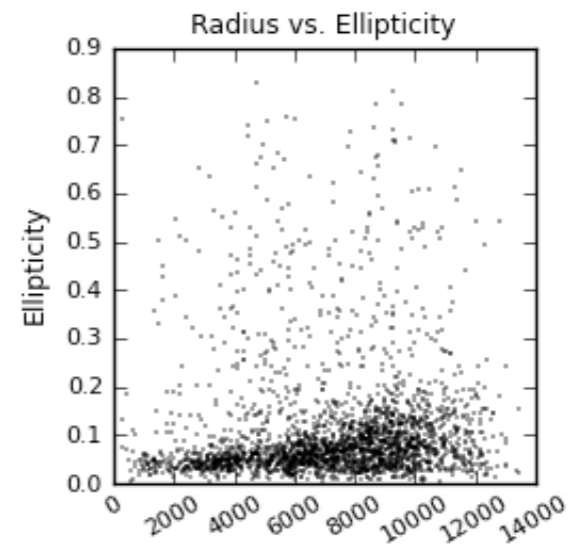
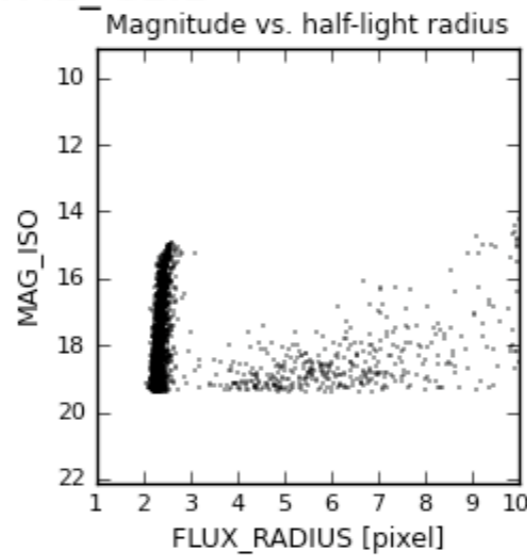
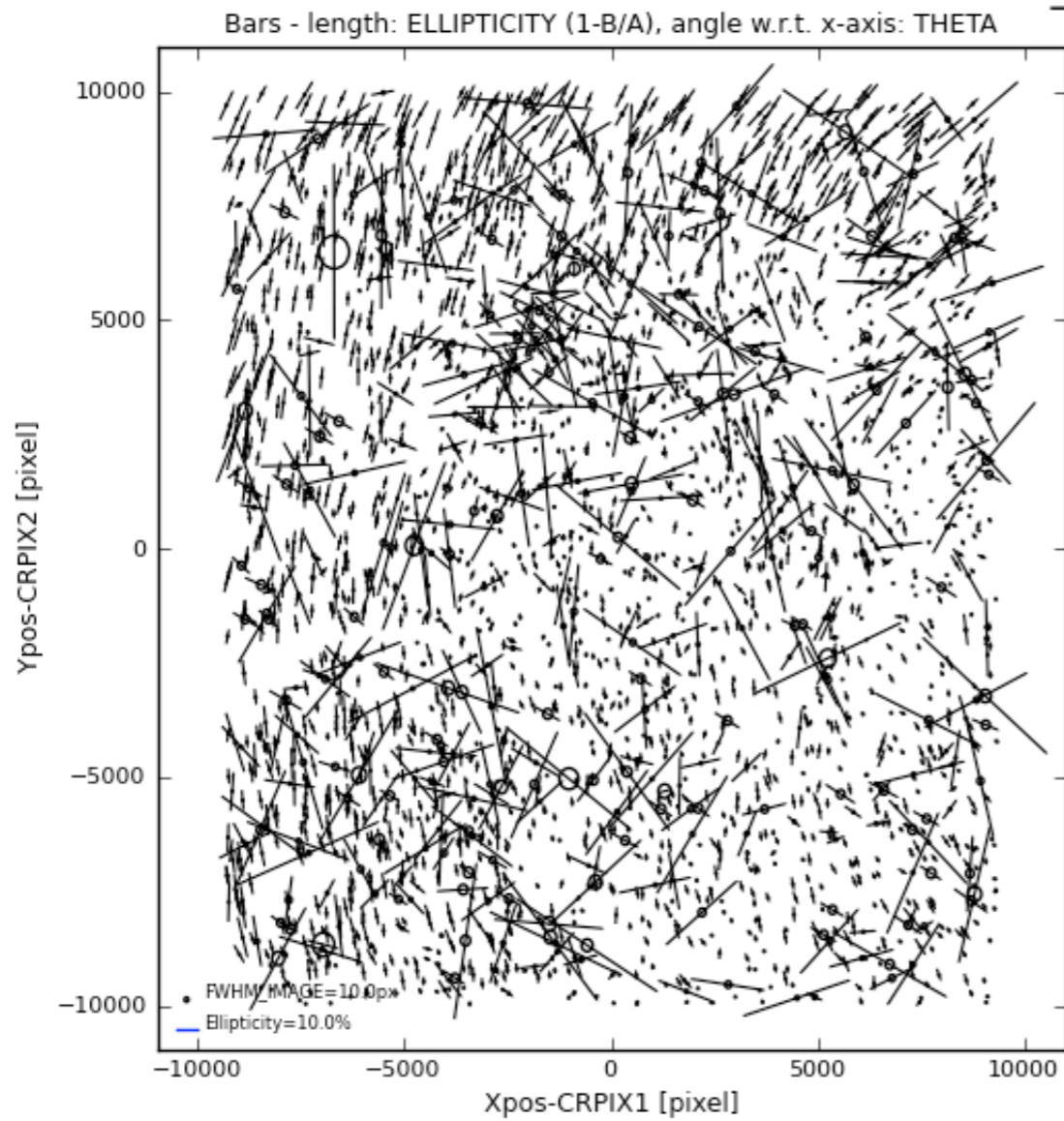


PSF Anisotropy for 1 frame(s) with DATE-OBS ??, filter OCAM_r_SDSS
KIDS_355.3_-31.2





PSF Anisotropy for 1 frame(s) with DATE-OBS ??, filter OCAM_i_SDSS
KIDS_347.1_-32.1



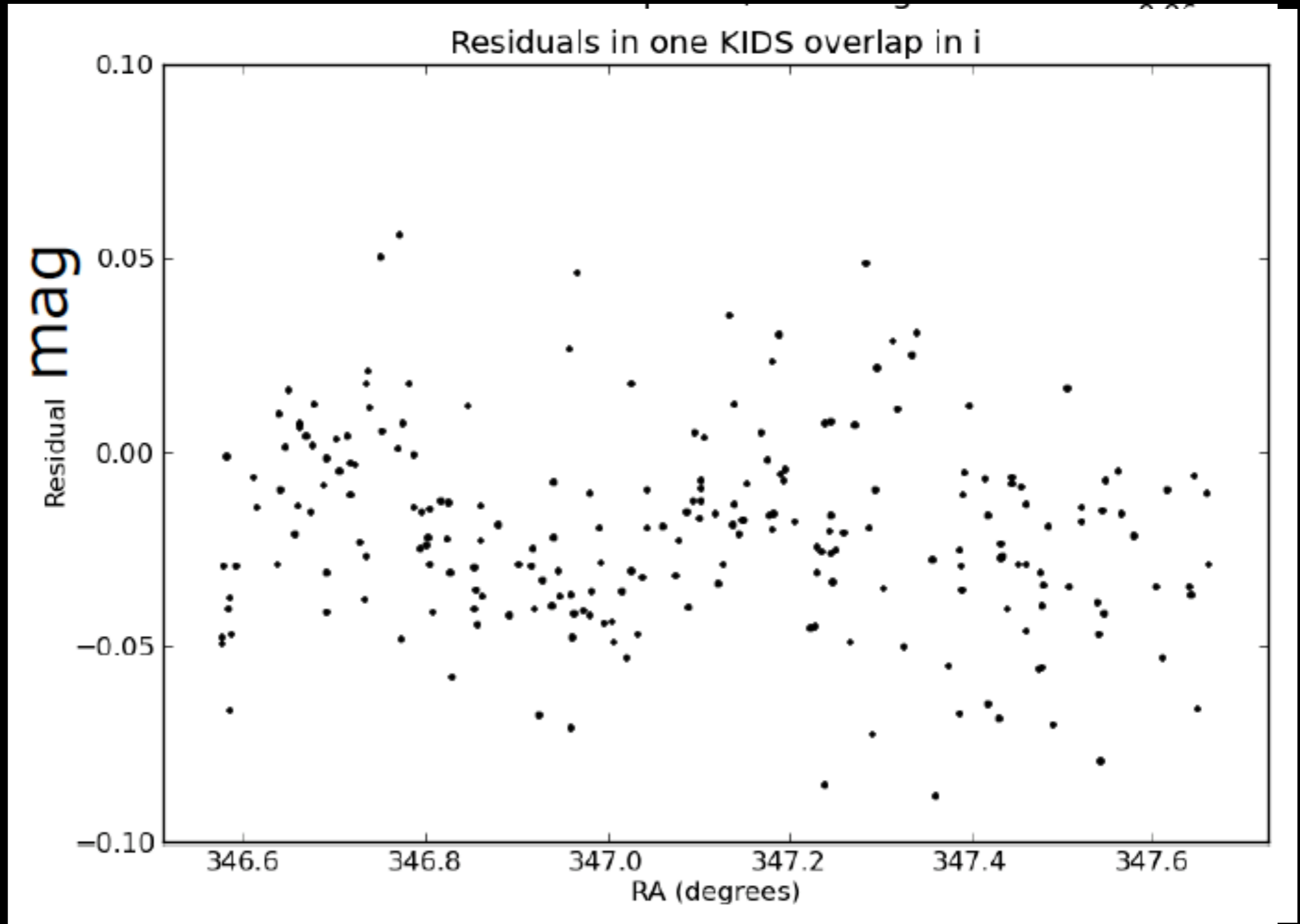
Data processing

Photometric accuracy in overlap regions

1 night
(03 Oct 2011)

seeing: 0.6",
0.7", 0.8"

exptime = 1200s



Photometry from dithers on SA fields ...work in progress...

Photom errors	u	g	r	i
stdev internal end-to-end	0.019	0.012	0.011	0.014
Stdev in <ZPT> for mosaic	0.011	0.010	0.012	0.011
Stdev in <ZPT> per ccd	0.037	0.010	0.016	~0.015
D(ZPT) / CCD	0.015	0.005	0.005	0.007

Internal photometry error for single source within dither 1-2%
 → good hope to get to 1% syst. uncertainty in relative photom.

Also good hope to get to 1.5% absolute photometry per band

More work to be done on ZPT determination per CCD

VST calibration plan

Calibration data taken at specific intervals:

- biases daily
- dome flats once every three days
- darks once every week
- sky flats every clear night in 2-3 filters (key bands u, g, r, i once every three days)

Master flat fields are combination of dome flats (for small scale variations) and sky flats (for large scale variations)

VST calibration plan

Zeropoints derived based on:

- sky transparency and instrument efficiency measurement 3 times per night, by monitoring polar standard field in composite u,g,r,i filter
- at evening twilight and midnight standard star field is observed with composite u,g,r,i filter + individual key filters
- If 'user band' is used, standard field also observed in this filter

Due to North pointing restrictions, southern secondary standard fields being defined

VST calibration plan

Monitoring the Photometric Calibration

