# VST KO Morning Agenda

10.30-11.00	Coffee
11.00-11.30	VST ATLAS - science goals + current observing plan
	- science goals
	- modified survey footprint
	- exposure times
	<ul> <li>overheads with/without binning</li> </ul>
	– mosaicing+overlaps+gaps
	- order of observing ugr, iz
	<ul> <li>order of observing fields</li> </ul>
11.30-11.40	- Discussion
11.40-12.10	- CASU data reduction (Mike Irwin)

- pipline description
- data products
- astrometric accuracy
- UKIDSS/VISTA experience
- 12.10-12.20 Discussion
  - Global Calibration (Steve Maddox)
    - method
    - target accuracy for photometry
- 12.50-13.00 Discussion
- 13.00-14.00 Lunch

12.20 - 12.50

#### VST ATLAS Survey

- \* VST ATLAS (+VHS)  $\rightarrow$  Southern SDSS in ugriz(+YJHK)!
- \* 60s exposures in ugriz better throughput in u, z
- \* Footprint ~2600deg<sup>2</sup> in SGC and ~1450deg<sup>2</sup> in NGC
- \* 1"-1."4 seeing better than SDSS median at 1."4
- \* Best seeing used for KiDS
- \* ~45 nights per year for 2 years accelerated!

#### VST ATLAS Core Team

- \* CASU (Mike Irwin et al, Cambridge) does the basic reduction using the VST Data Flow pipeline
- Steve Maddox (Nottingham) leads the overall global calibration process
- Nigel Metcalfe, helped by Peter Draper
   (Durham) will do QC on the ATLAS products
- WFAU (Bob Mann et al, Edinburgh) to provide archiving facilities, additional to the ESO archive

# VST Mirror in Cell, 13 November



# VST Commissioning Schedule

ID	Task Name	Duration	Start	Finish	Jan 11         Feb 11         Mar 11         Apr 11         May 11         Jun 11         Jul 11           1         8         15/22/29/5         12         19/26/5         12         18/25/1         8         15/22/29/6         13/20/27/3         10         17/24/1         8         15/22/29
1	VST Commissioning	100d	06.01.11	30.06.11	
2	VST1A: Tracking/Guiding, AO S	52	06.01.11	27.02.11	VST
3	MS: Performance reached	0d	27.02.11	27.02.11	27.02
4	ESO Acceptance Tests 1	7	27.02.11	06.03.11	ESO ESO
5	OmegaCam Installation	11	06.03.11	17.03.11	OCAM
6	VST1B: Tracking Tuning with O	4	17.03.11	21.03.11	🖕 VST
7	MS: Ready for OCAM1	0d	21.03.11	21.03.11	▲ 21.03
8	OCM1A	9	21.03.11	30.03.11	OCAM
9	VST1C: Realignment	10	30.03.11	09.04.11	VST
10	VST2A: ADC Tuning	12	09.04.11	21.04.11	VST
11	ESO Acceptance Test 2	7	21.04.11	28.04.11	ESO
12	Spacer Manufacturing	84	30.03.11	22.06.11	
13	Analysis Data of OCM1A	28	30.03.11	27.04.11	OCAM
14	OCM1B	10	28.04.11	08.05.11	OCAM
15	VST2B: Details not specified	28	08.05.11	05.06.11	, VST
16	MS: Ready for OCAM2	0d	05.06.11	05.06.11	▲ 05.06
17	OCM2	14	05.06.11	19.06.11	ОСАМ
18	ESO Acceptance Test 3	7	19.06.11	26.06.11	ESO ESO
19	MS: End of Acceptance	0d	26.06.11	26.06.11	26.06
20	Spacer Installation	2	26.06.11	28.06.11	OCAM
21	Verification	2	28.06.11	30.06.11	SO ESO
22	MS: End of Commissioning	0d	30.06.11	30.06.11	♦ 30.06

# VST ATLAS – previous footprint

#### VST ATLAS Survey



R.A. (2000.0)

## VST ATLAS – new footprint

#### VST ATLAS Survey



R.A. (2000.0)

### ATLAS Science Summary

- \* VST ATLAS offers high quality science at low cost
- \* Cosmology Package to rival WFIRST!
  - \* BAO at z~1.5 via QSO clustering
  - \* BAO at z~3 via QSO Lyman  $\alpha$  forest
  - \* Gravitational Growth rate at z=1-3 via QSOs
  - \* ISW via LRGs
  - \* QSO Lensing vs galaxy ugrizYJHK photo-z
- \* Other Science
  - \* Stellar Streams + Galactic Archaeology
  - \* Z~7 QSOs via ATLAS+VHS z dropouts
  - \* Beyond the Great Attractor....

## WHDF vs SDSS



### VST vs SDSS Bands



VST/OmegaCAM throughput cf unobscured 2.6m aperture

# OmegaCAM



#### Overheads

- \* Telescope preset 10secs 5mins????
  \* Filter exchange 50-110secs
  \* Acquisition of guide stars?
  \* CCD Read out 29secs
  \* Write to file 12secs
  \* Ordered into fits file 20secs
- \* Although in SMP we use 38s for r/o+write

## Observing Sequence (1)

- \* ugr sequentially in dark time, iz in bright time
- \* Guide star acqn (if needed) during first exposure
- CR split assumed unnecessary
- Filter change/read out dominates at 50s/40s
- × ~100% overhead
- Offset by 54' between pointings
- \* 10% overlap in RA and Dec for photometric calibration
- 80" and 20" gaps left 43' offsets needed in RA to eliminate 80" gaps 25% increase in survey time

## Observing Sequence (2)

2x2 binning reduces r/o+write time to ~10s
2x30s exposures CR split – 4' offset in RA, 7' in Dec
Observe sequentially in one filter
Preset/acquisition main overhead 10s.
33% overhead - can double exposure to 60s?
Or just double u band exposure?

#### Other Issues

\* Does ATLAS image the KiDS South area first for calibration?

- \* Does VHS do KiDS area before VIKING?
- Coordination between VISTA Hemisphere Survey and ATLAS

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## VST KO Afternoon Agenda

14.00-14.30	- Quality Control (Nigel Metcalfe/Peter Draper) - DQC for VST ATLAS - PanSTARRS experience
14.30-14.40 14.40-15.10	<ul> <li>Discussion</li> <li>Data Archiving (Eckhard Sutorius)</li> <li>ESO Archive</li> <li>WFAU Archive</li> <li>UKIDSS/VISTA experience</li> </ul>
15.10-15.20	- Discussion
15.20-15.40	- Tea
15.40-17.00	<ul> <li>General Discussion</li> <li>Coordination with VST KIDS+VISTA VHS/VIKING (Will Sutherland)</li> <li>Data release policy</li> <li>communications - wiki etc</li> <li>Science Goals y pipelinetarchive check</li> </ul>

17.00 End of meeting

