

Just good friends: The non-causal origin of black hole–galaxy scaling relations

Knud Jahnke

Andrea Macciò



Max-Planck-Institut für Astronomie, Heidelberg



www.mpia.de/coevolution

What drives the growth of black holes?
Durham, 29th July 2010

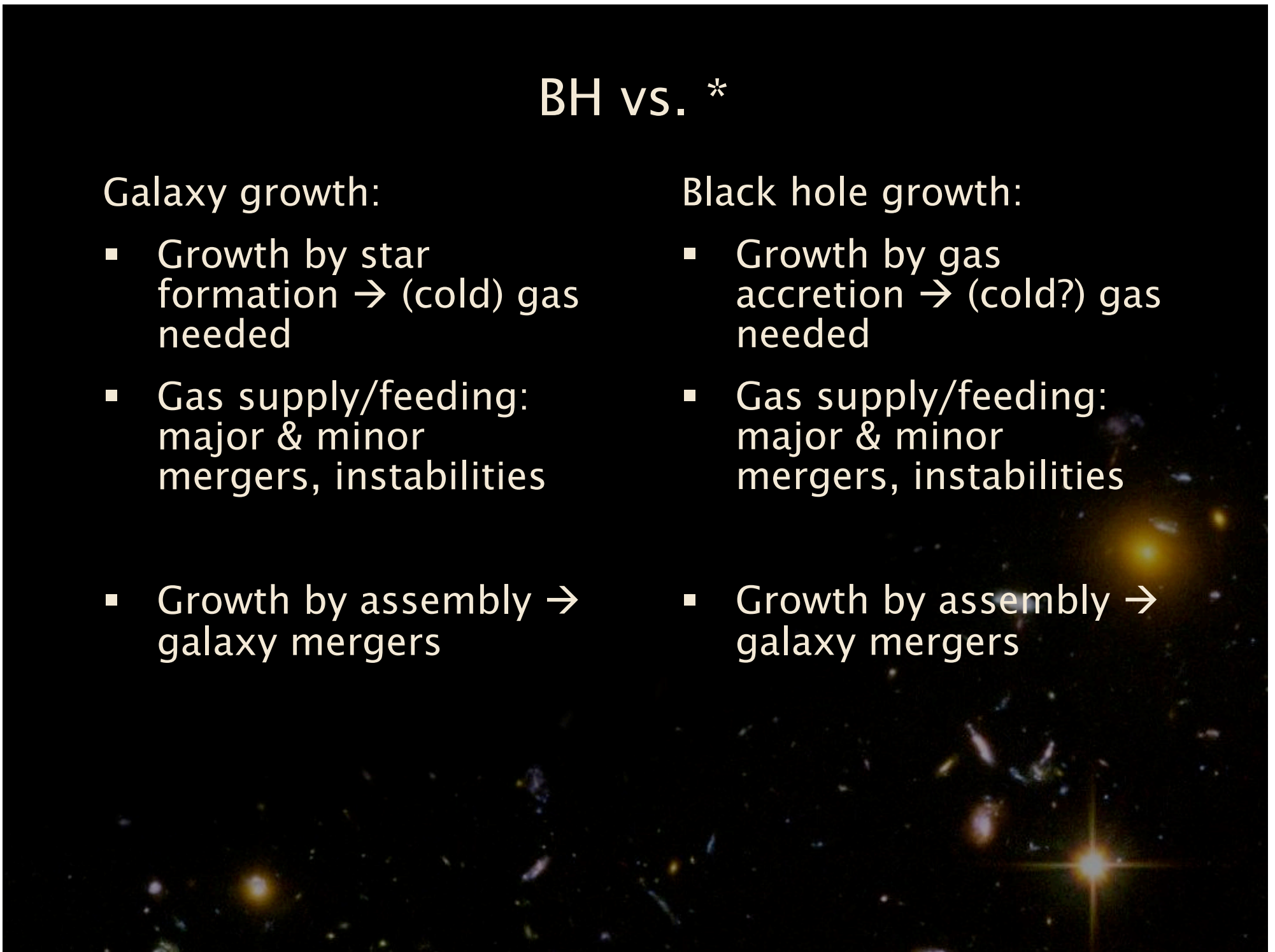
BH vs. *

Galaxy growth:

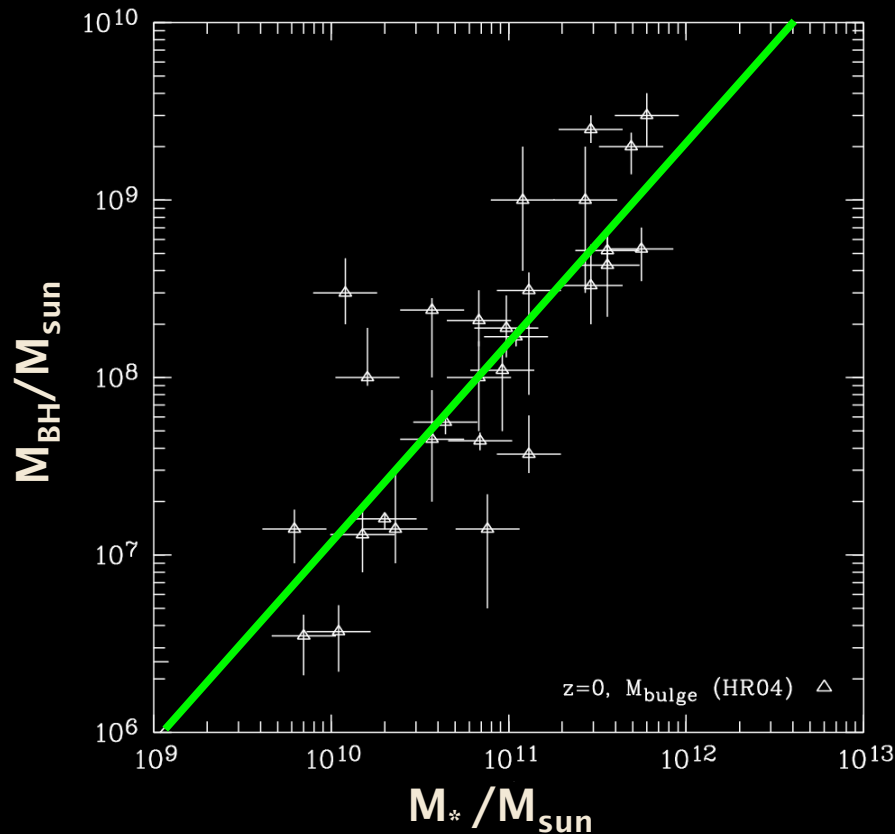
- Growth by star formation → (cold) gas needed
- Gas supply/feeding: major & minor mergers, instabilities
- Growth by assembly → galaxy mergers

Black hole growth:

- Growth by gas accretion → (cold?) gas needed
- Gas supply/feeding: major & minor mergers, instabilities
- Growth by assembly → galaxy mergers



BH vs. *



$z=0$: Häring&Rix 2004

- Tight correlation: BH-bulge, 0.3 dex scatter (measurement or intrinsic?)
- Evolution with z :
 - $z < 1.5$: ~ low
 - $z = 3$: x2-6
 - $z = 6$: x30

Dave A. intro: „somehow the black hole knows about the spheroid it is located in“

I USED TO THINK
CORRELATION IMPLIED
CAUSATION.



THEN I TOOK A
STATISTICS CLASS.
NOW I DON'T.

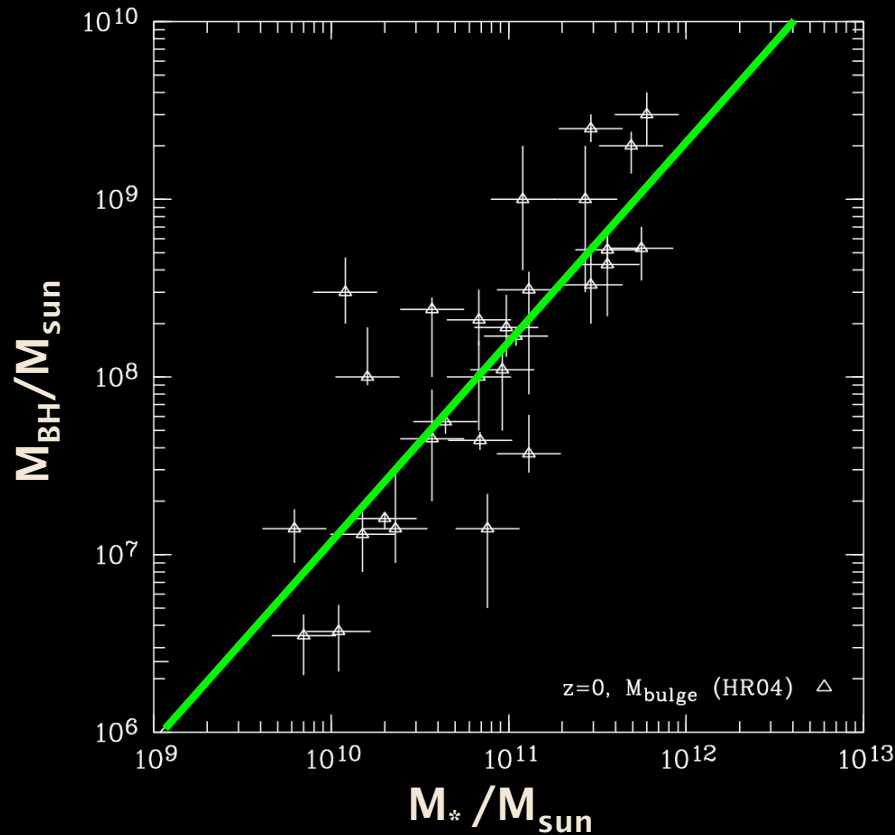


SOUNDS LIKE THE
CLASS HELPED.

WELL, MAYBE.



BH vs. $*$: Correlation or causation?



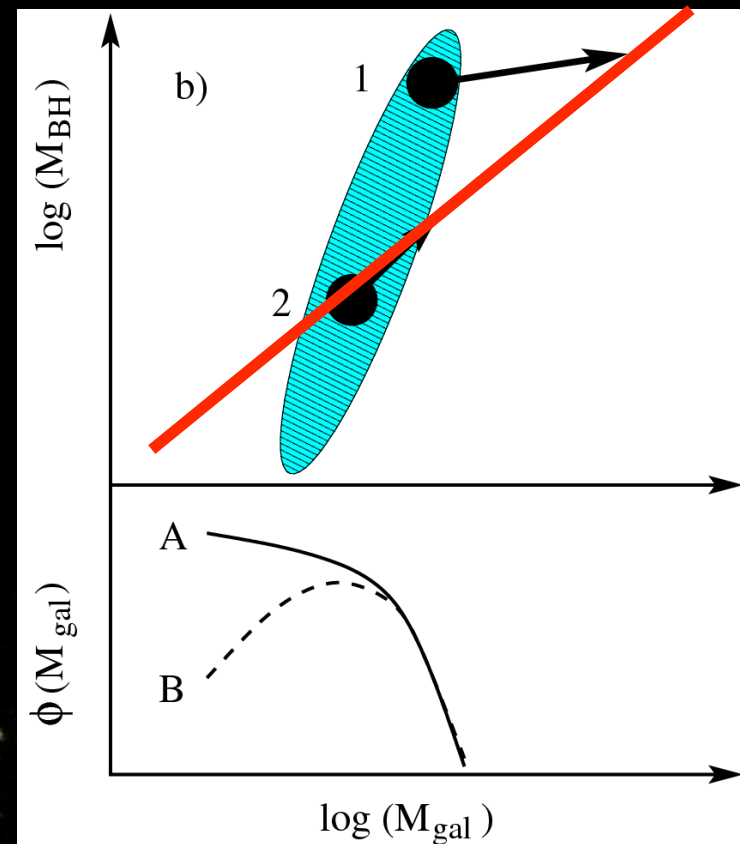
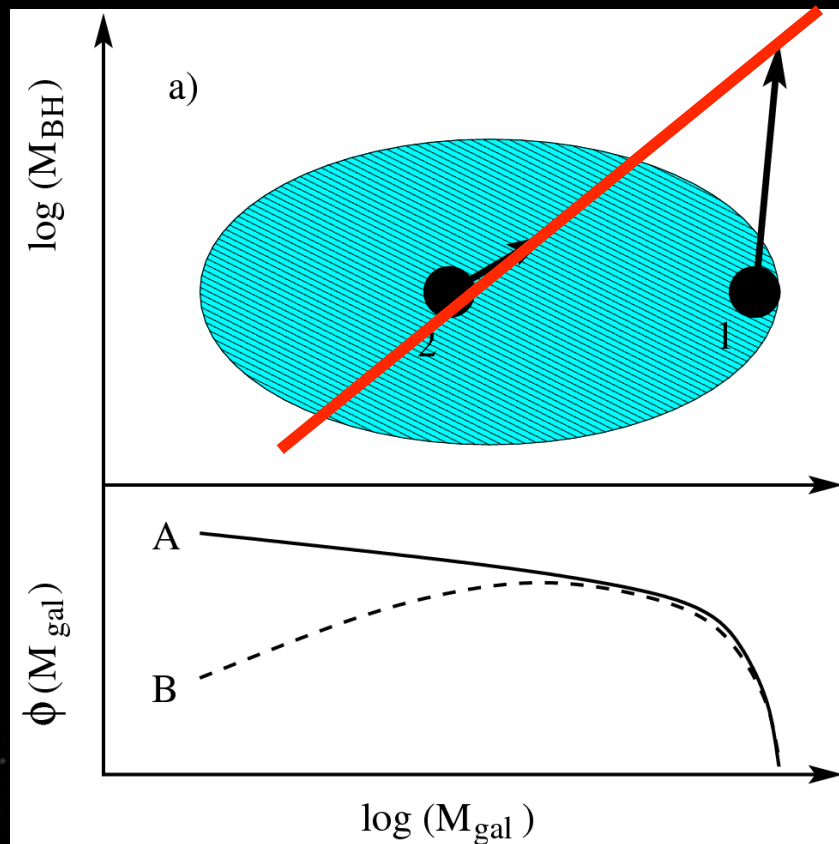
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→ Coupled evolution?

BH vs. $*$: Correlation or causation?

Chien Y. Peng (2007): Galaxy merging averages properties; is $M_{\text{BH}}-M_*$ relation due to „central limit theorem“?



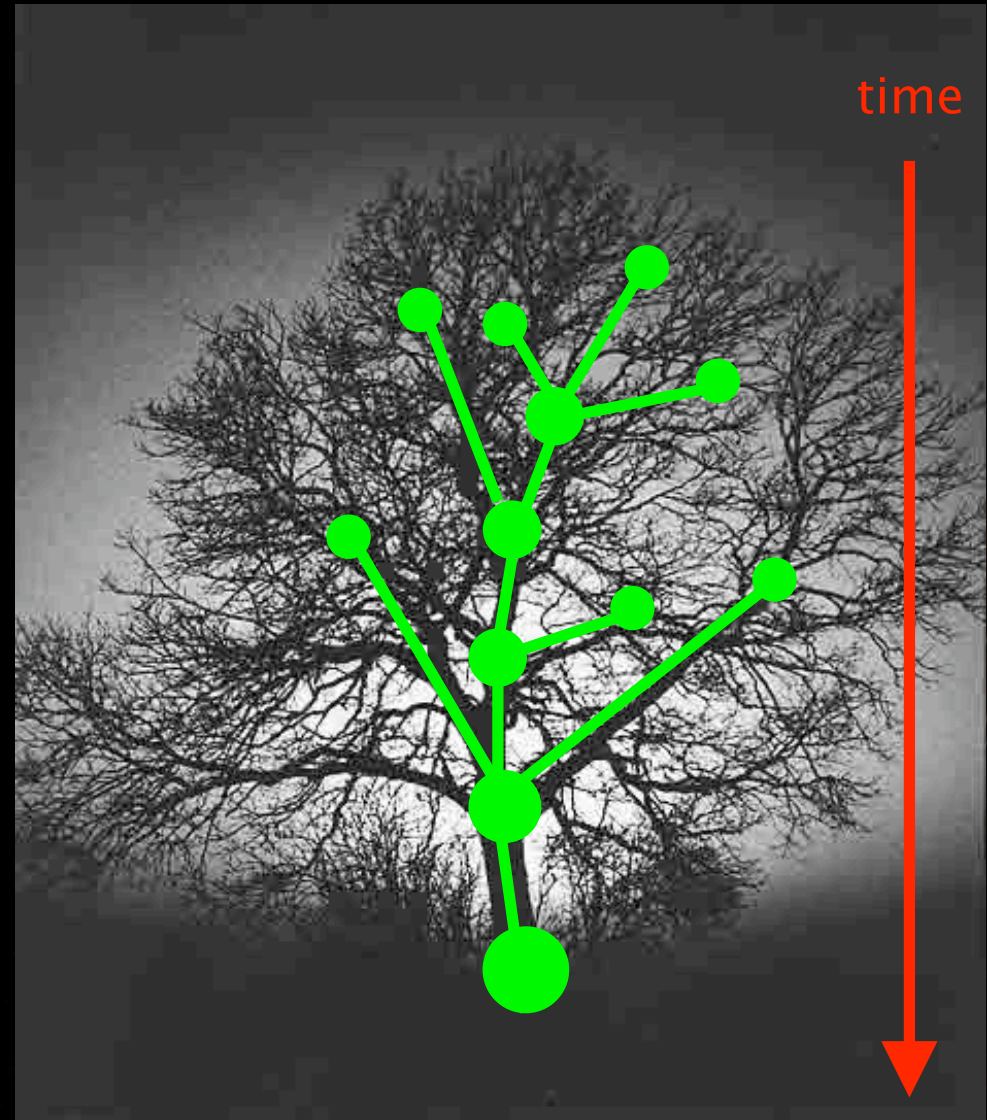
Using a realistic Universe

What about the real Universe?

Use simulated set of DM halos and it's assembly merger tree (*Pinocchio* code, Monaco et al.)

→ BH seeds? M_* seeds?

KJ & Macciò, *subm. to ApJL*
arXiv:1006.0482



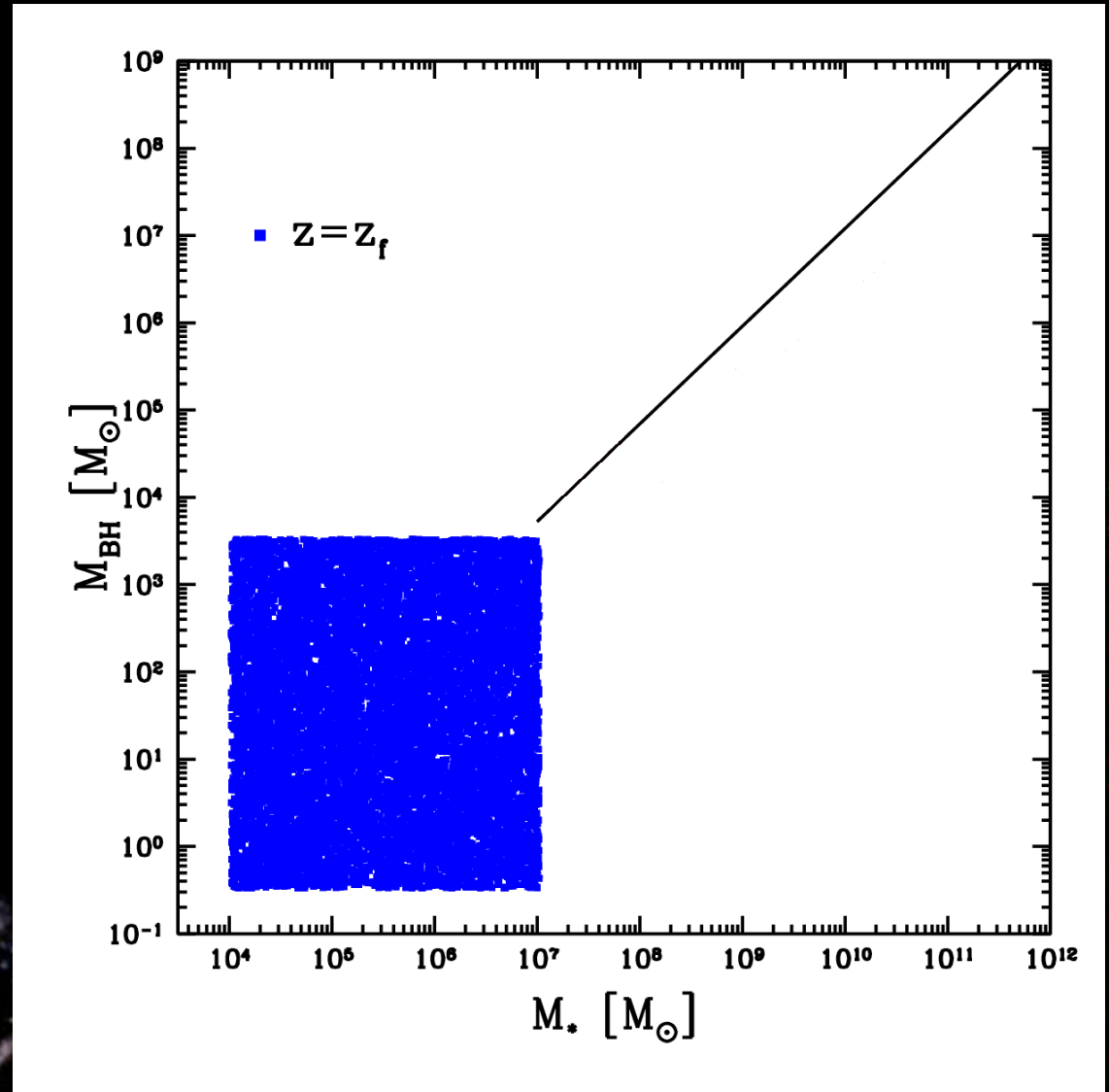
Using a realistic Universe

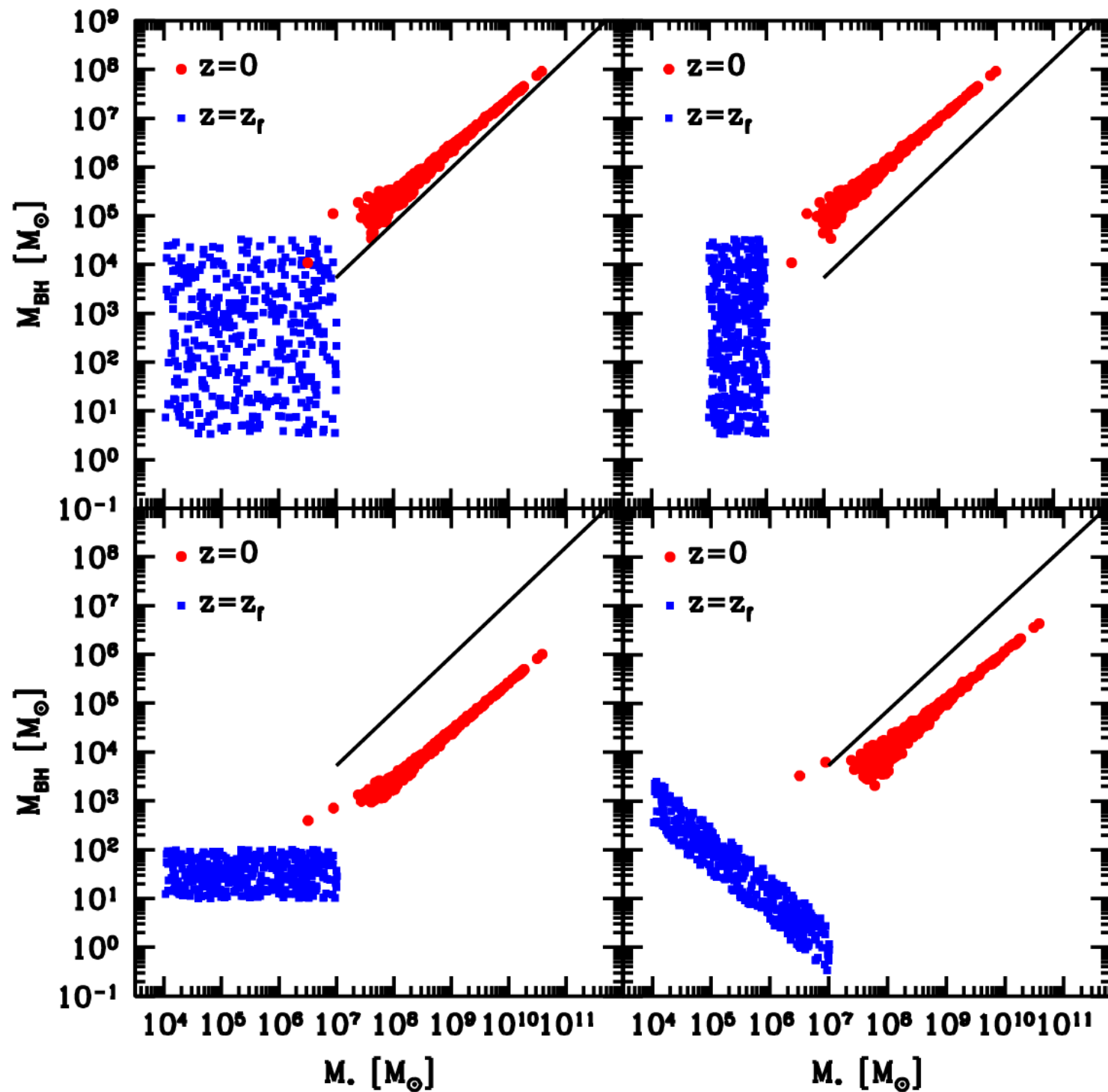
w/ **Andrea Macciò** (MPIA):

- dark matter merger tree ($z=20\dots 0$)
- seeded with M_* , M_{BH}
- uncorrelated at large z

→ Produces very tight relation at $z=0$

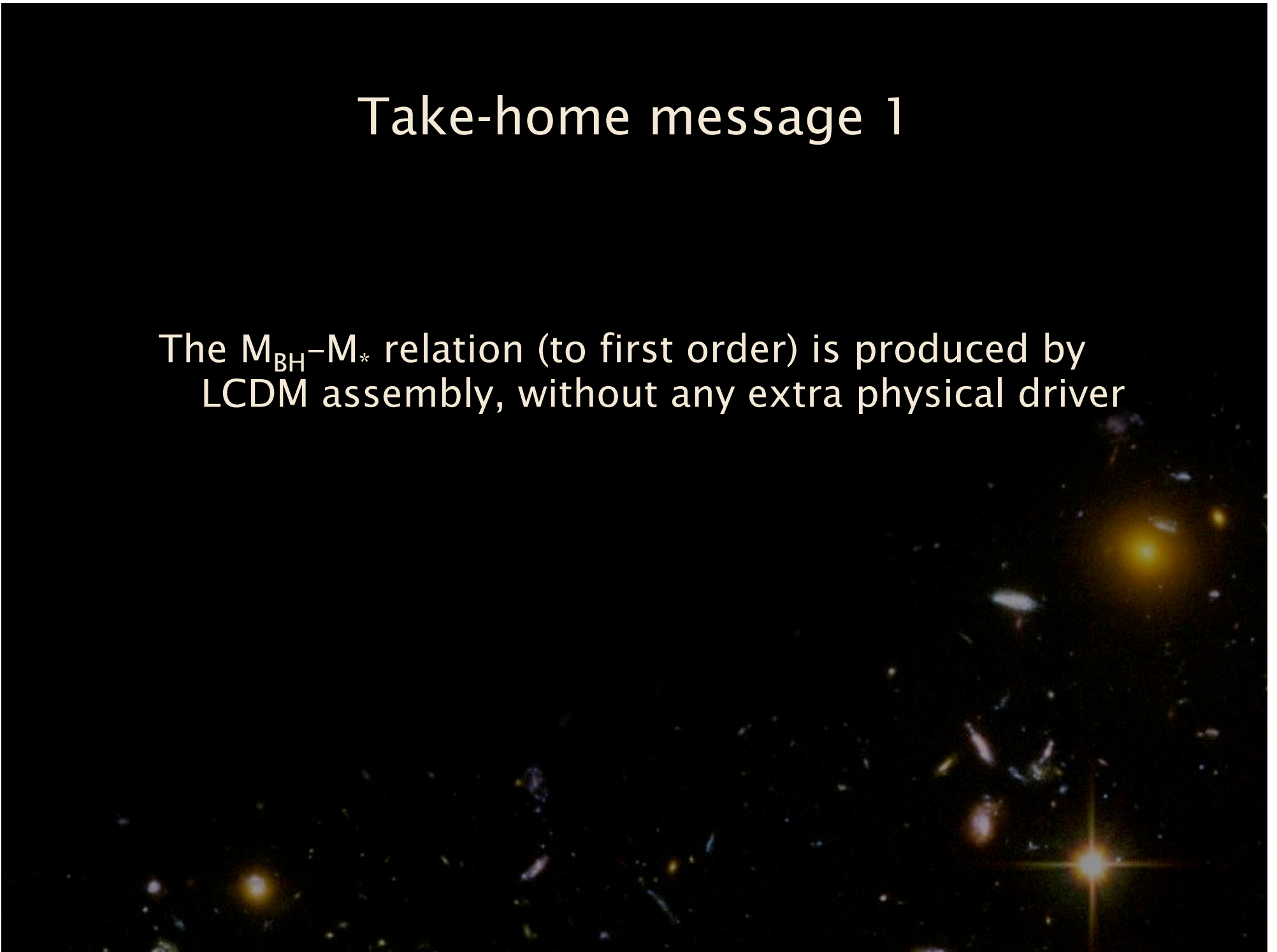
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Take-home message 1

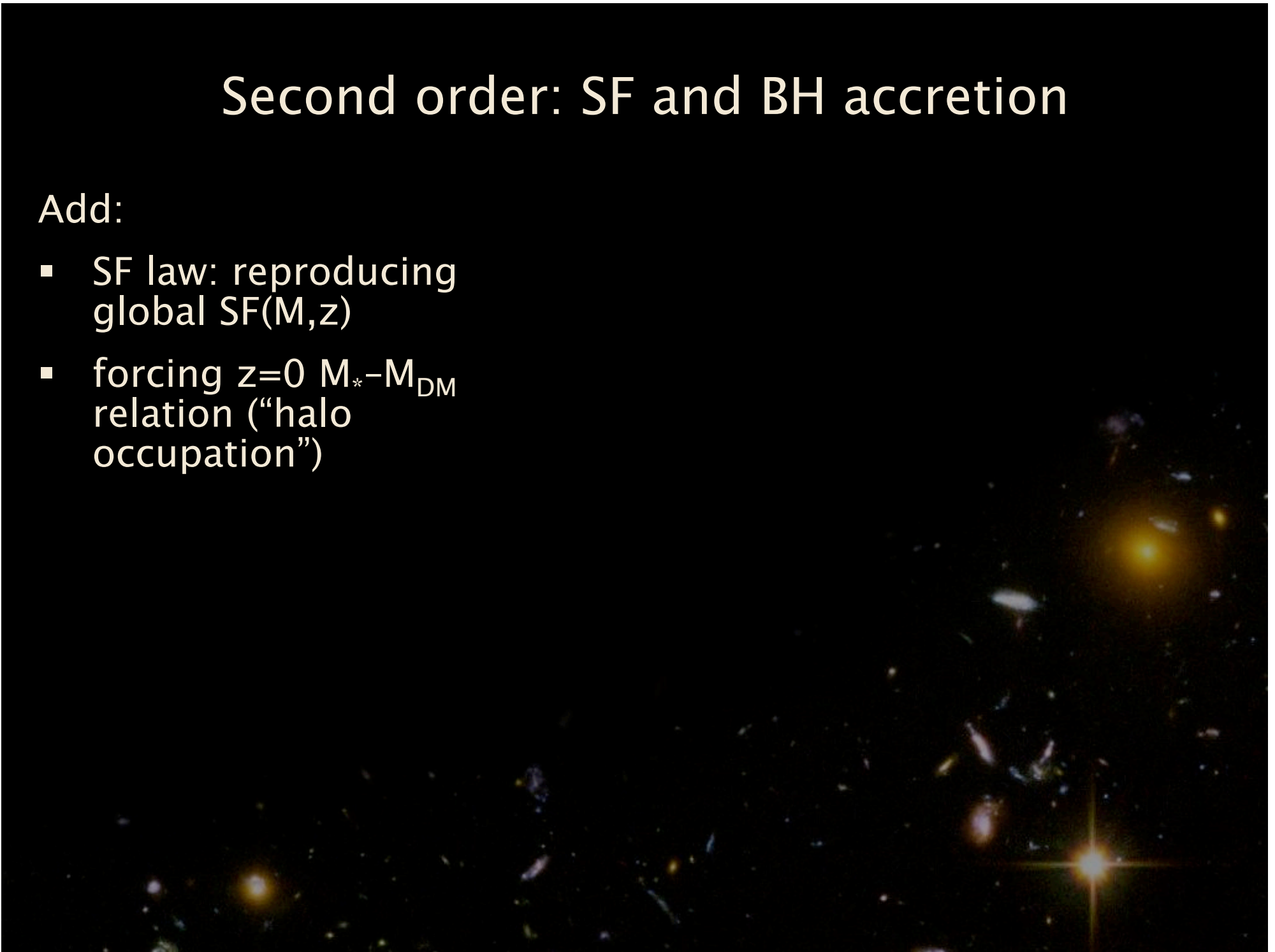
The $M_{\text{BH}}-M_*$ relation (to first order) is produced by LCDM assembly, without any extra physical driver



Second order: SF and BH accretion

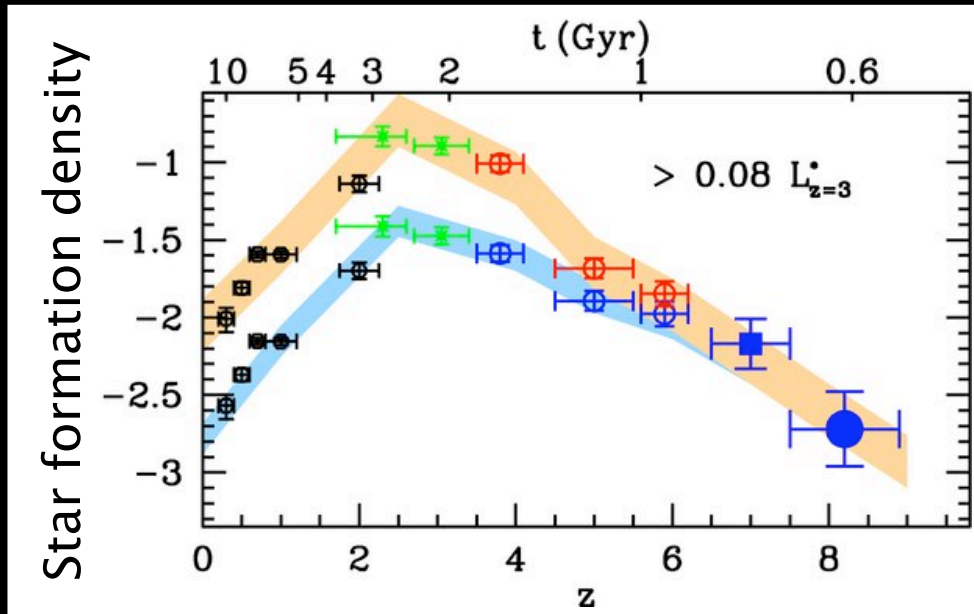
Add:

- SF law: reproducing global SF(M,z)
- forcing $z=0$ M_*-M_{DM} relation (“halo occupation”)



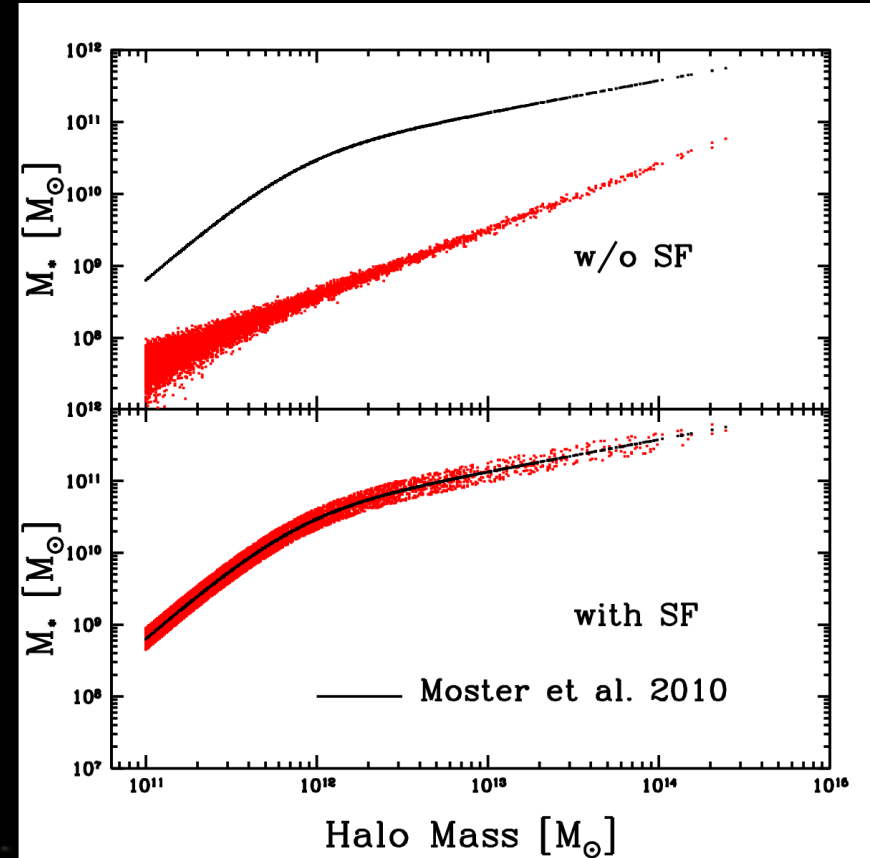
Second order: SF and BH accretion

Add:



$dM_*(z)$ from star formation

Bouwens+ 2010b



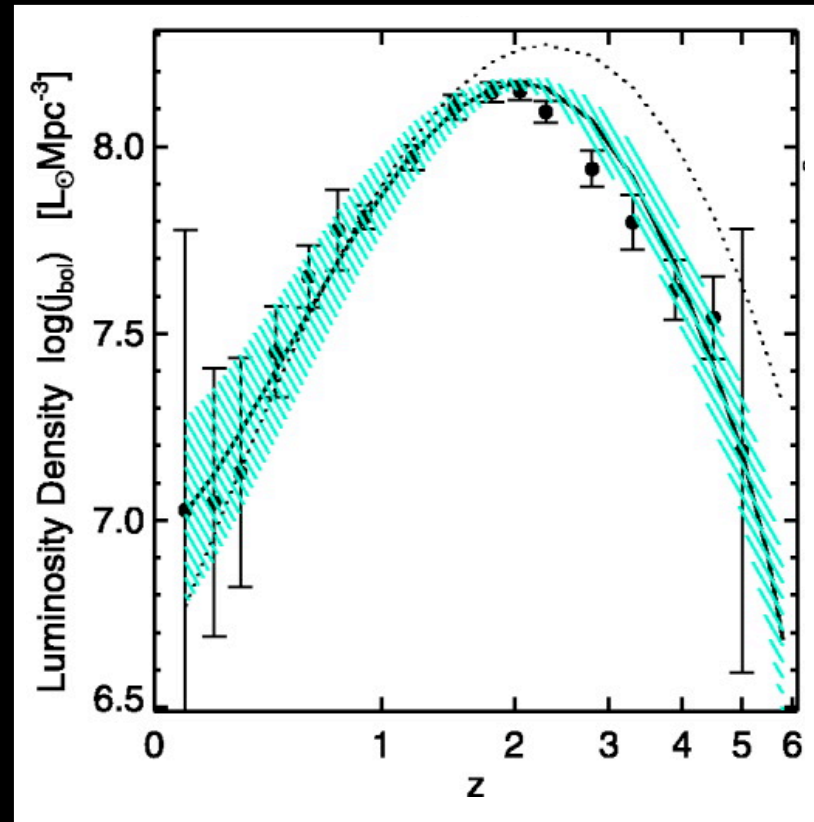
Halo occupation distribution

Moster+ 2010

Second order: SF and BH accretion

Add:

- SF law: reproducing global SF(M, z)
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- BHA law: global BHA(z) + random doubling events (matching $z=0$ normalization)



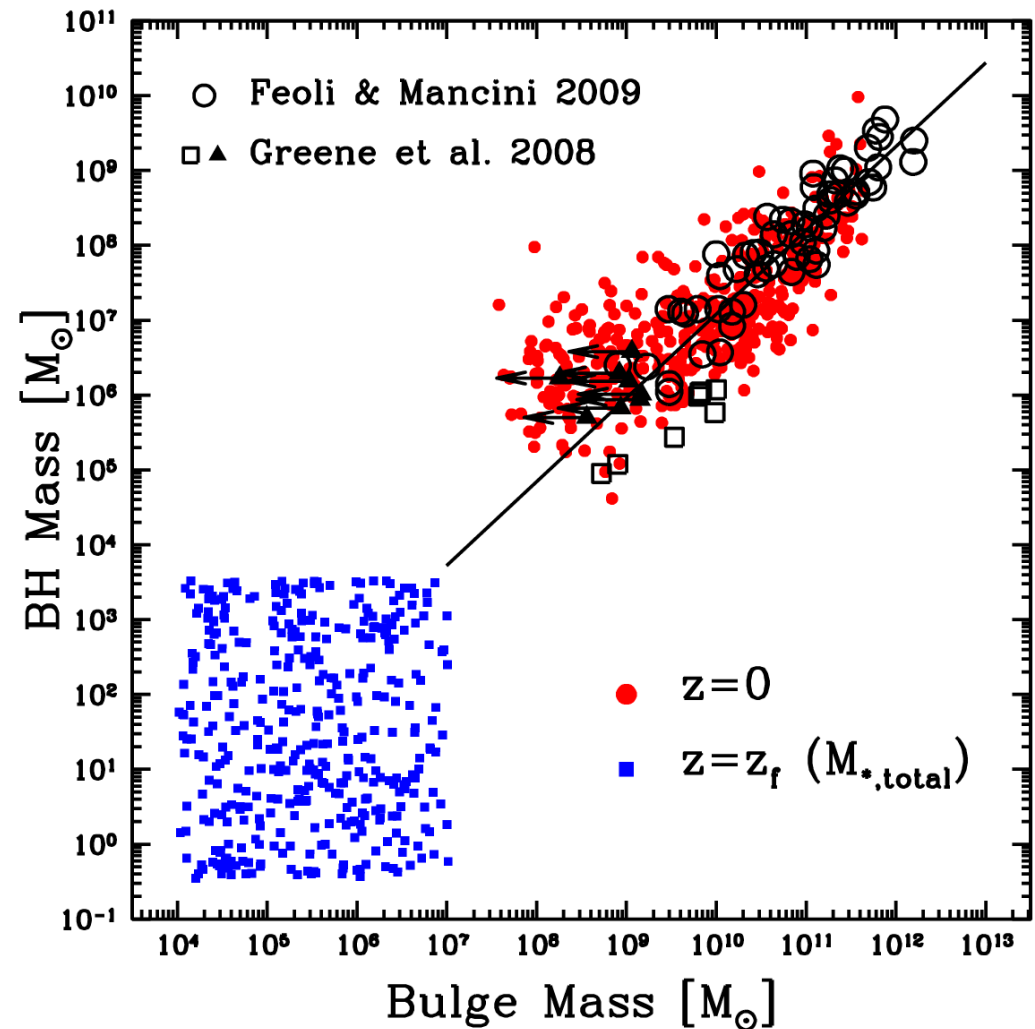
$dM_{\text{BH}}(z)$ from BH accretion

Hopkins+ 2007

The origin of the BH–galaxy scaling relations

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- SF law: reproducing global SF(M, z)
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- disk \rightarrow bulge mass conversion when merging

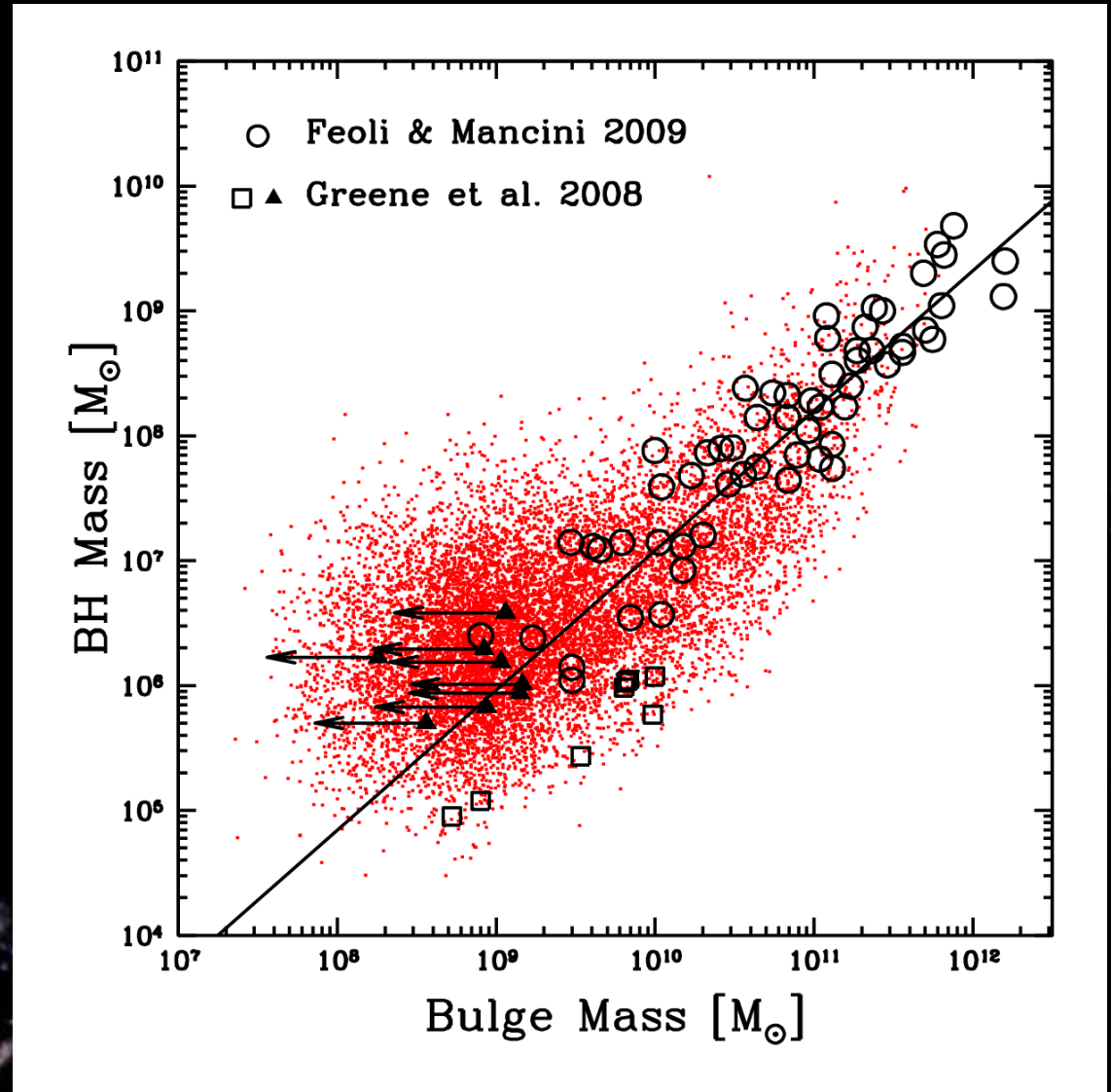


The origin of the BH–galaxy scaling relations

Add:

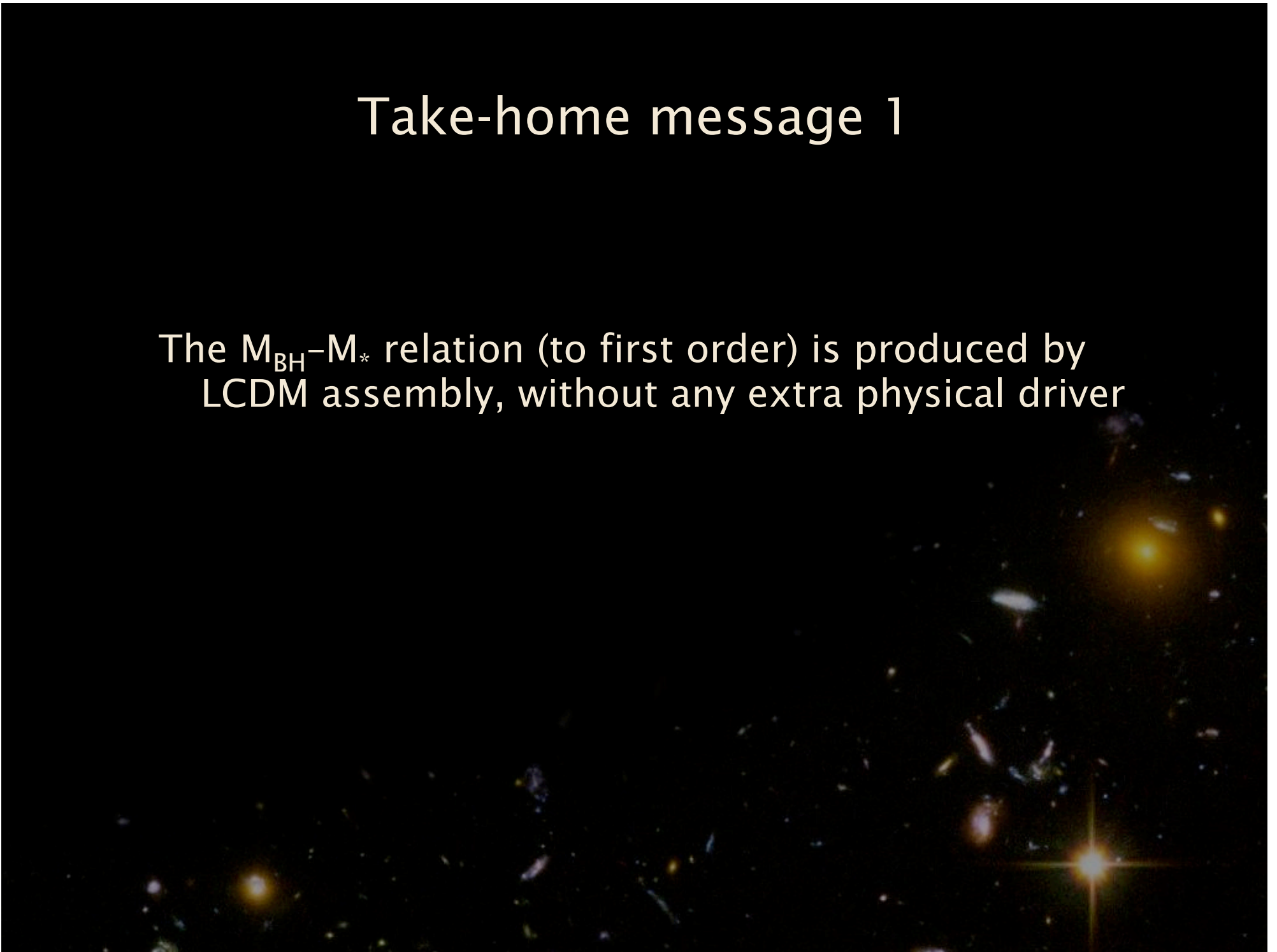
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($6.5 \times 10^6 \rightarrow \sim 10,000$ halos)



Take-home message 1

The $M_{\text{BH}}-M_*$ relation (to first order) is produced by LCDM assembly, without any extra physical driver



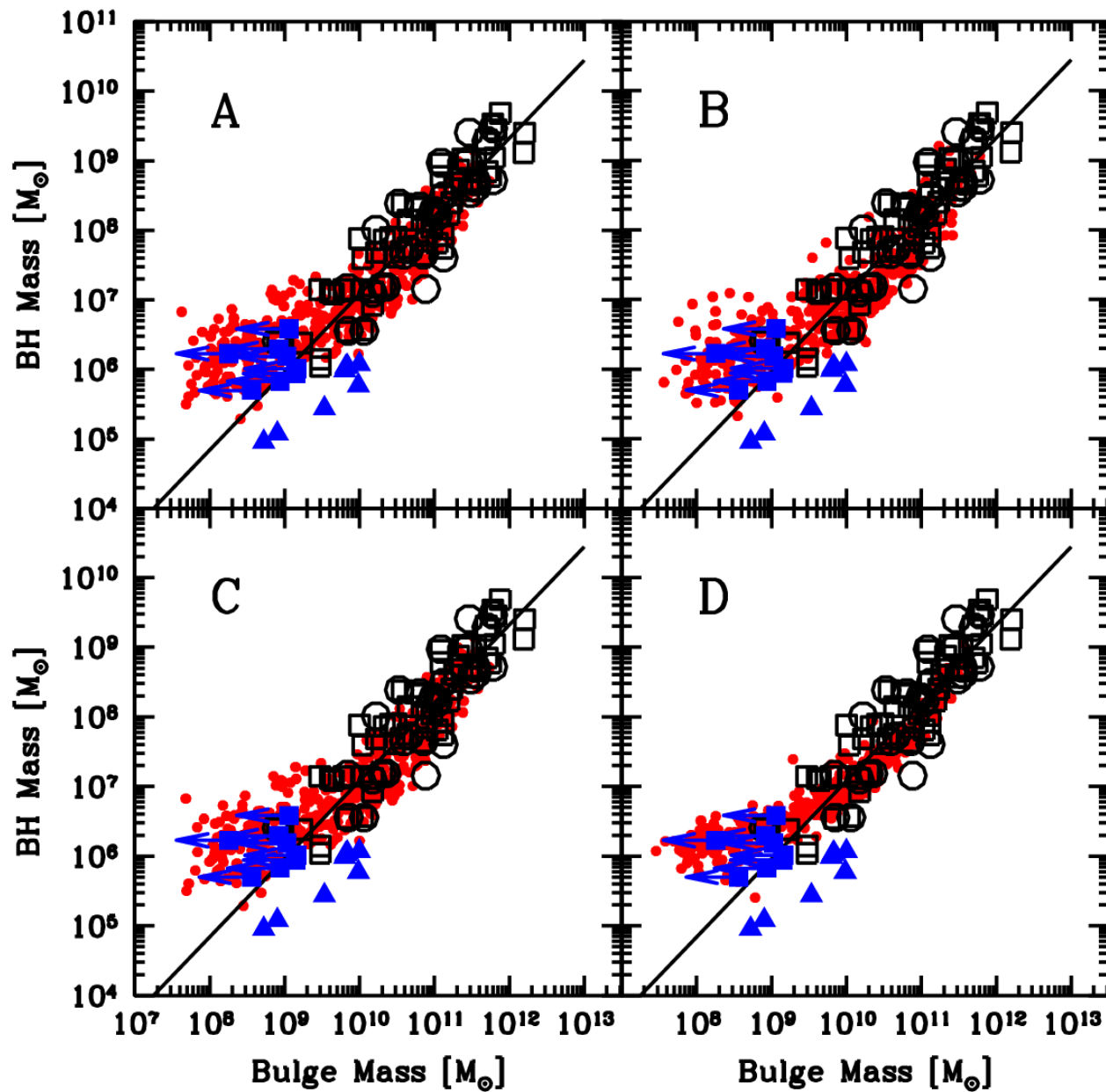
Take-home message(s) 2

- Exact shape/deviation from slope=1 due to 2nd order effects (SF cutoff at massive end \rightarrow halo occupation)
- AGN feedback not needed (for scaling relations!), but possibly for 2nd order (on par to grav. heating, modified SN feedback)
- Evolution in $M_{\text{BH}}/M_{\text{bulge}}$ at $z>1$: yes \rightarrow early growth of BHs \rightarrow so SF and BHA not strictly parallel
- Scatter evolution interesting diagnostics for seed BHs



~the end~

- A: reference
- B: $SF(z) = \text{const.}$
- C: $SF(z) = \text{const.}$ & $BHA(z) = \text{const.}$
- D: no random component in BHA



Finer consequences

Merger assembly/averaging path:

- Correlation with BH applies to all components taking part in merger assembly (bulge, halo,...)
- Hopkins: Correlations of BH with bulge but not central stellar/gas density → not necessarily taking part in this assembly process
- Automatic: more massive BHs → more luminous AGN live in more massive halos (Alison Coil and others)

Batcheldor 2010: „M-sigma is only limiting case, upper limit“ → only if mergers not taken into account

Ric Davies: 40% of gals without mergers since $z=2$ → different mode of bulge formation, so not properly represented in our sims

Hidden parameters: short/mid-term merger history not modelled, has influence on morphology and extra parameters (radius, compactness, binding energy, etc.)