



QSOs with VST-ATLAS & VHS

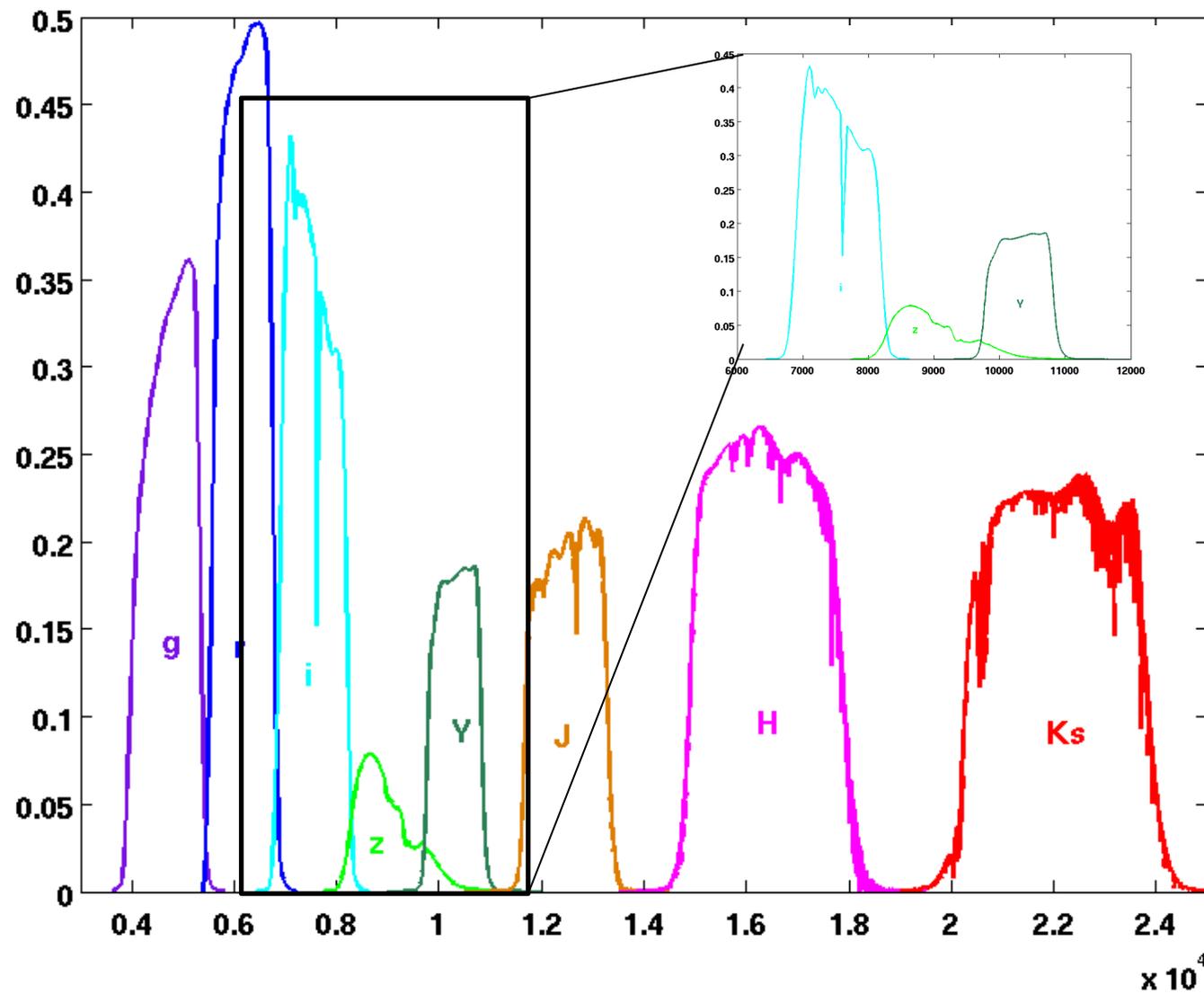
Manda Banerji (IoA, Cambridge)

Context

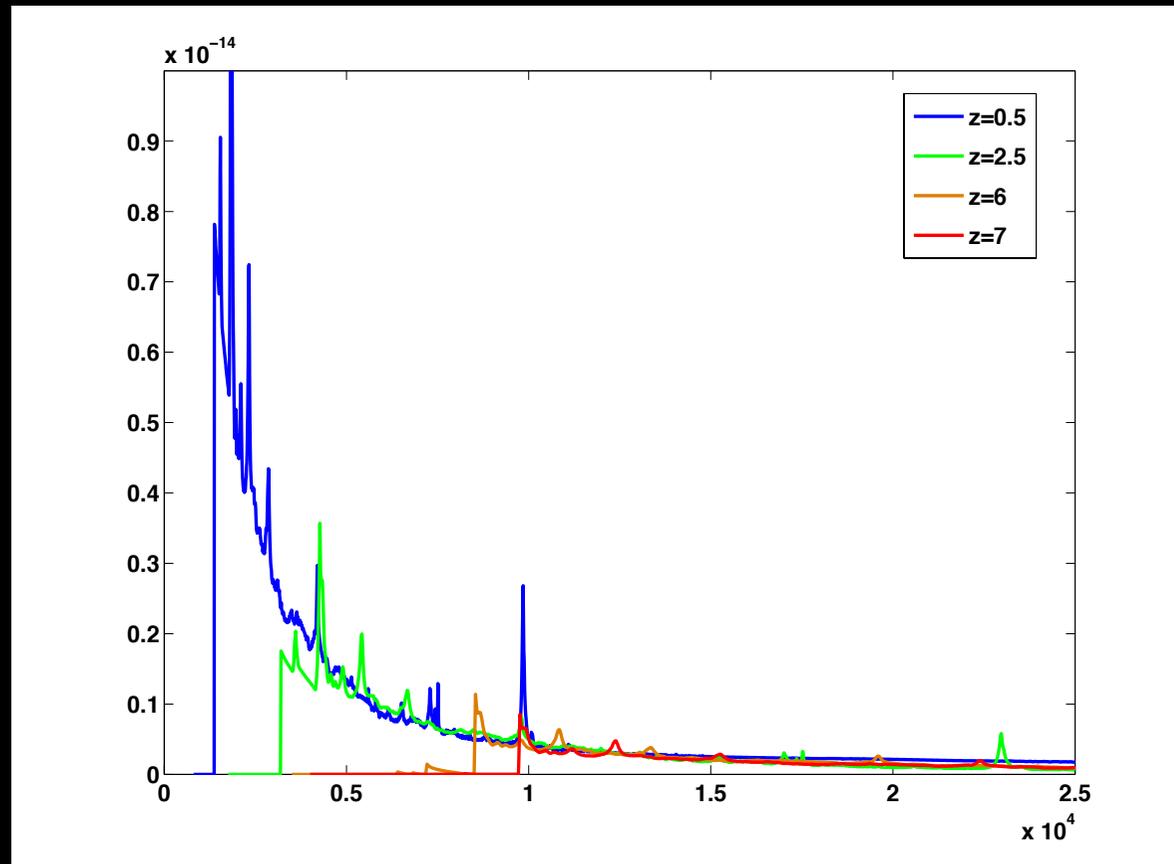
- VST-ATLAS+VHS is the southern hemisphere equivalent of SDSS+UKIDSS LAS which has already been very successful at finding high-z quasars (Bram Venemans talk)
- Bright high-z quasars suitable for detailed follow-up with next generation telescopes like E-ELT, ALMA and SKA
- VHS has already covered a significant fraction of VST-ATLAS area (Richard McMahon's talk this morning)

Simulating QSOs within
VST-ATLAS + VHS

Filters



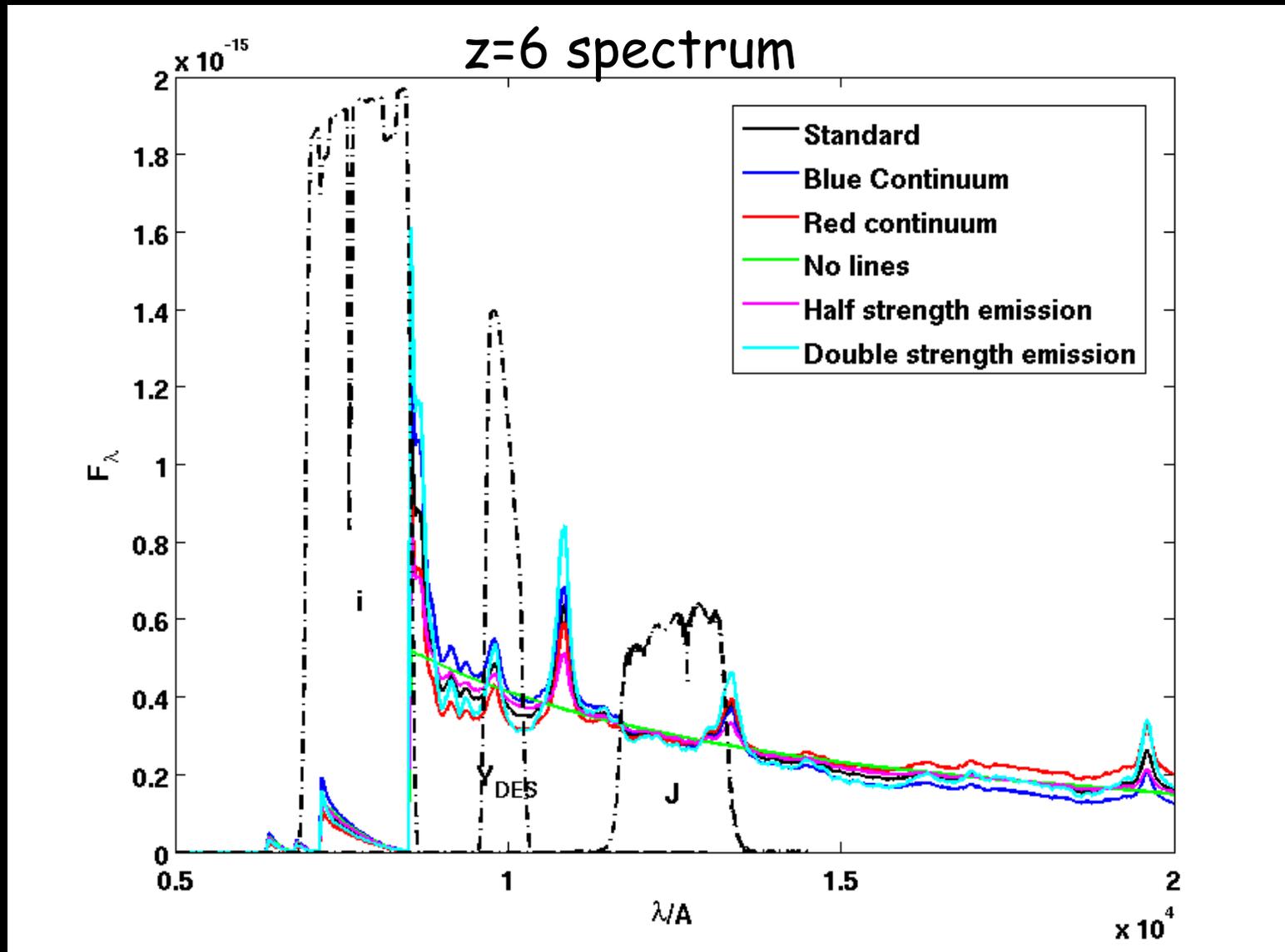
The QSO Spectrum



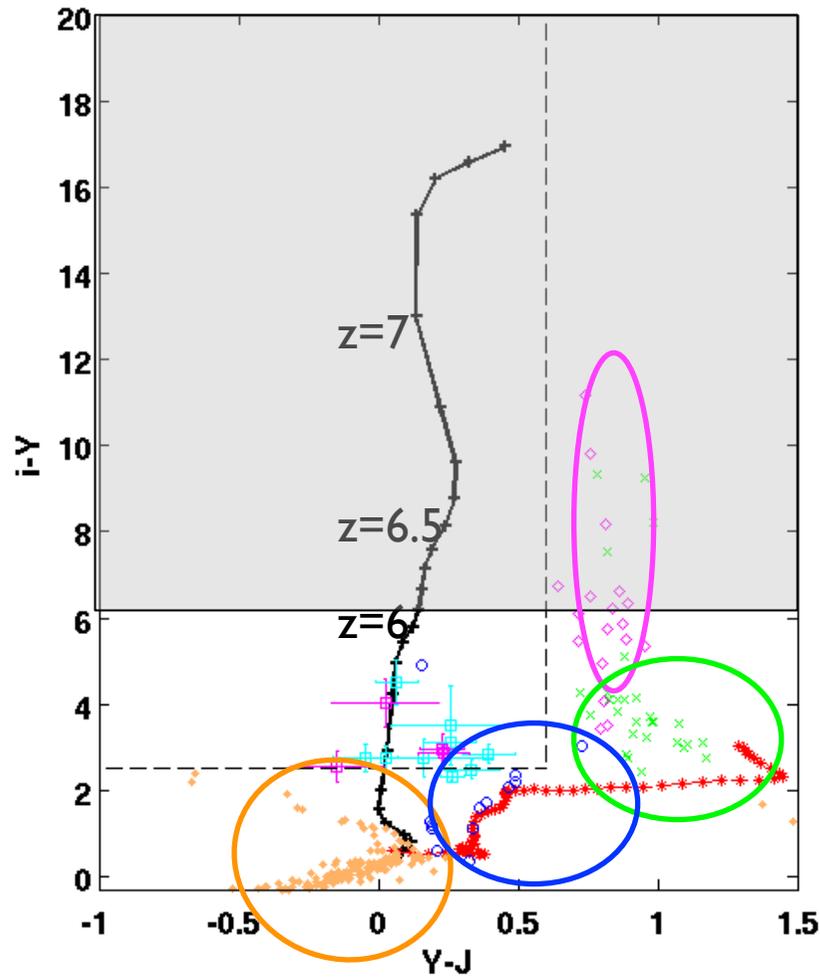
- Spectra taken from Maddox, Hewett, Warren & Croom (2008)

- Ly α absorption modeled according to Songaila 2004

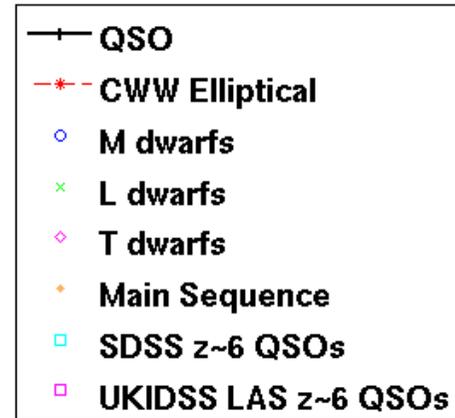
The QSO Spectrum



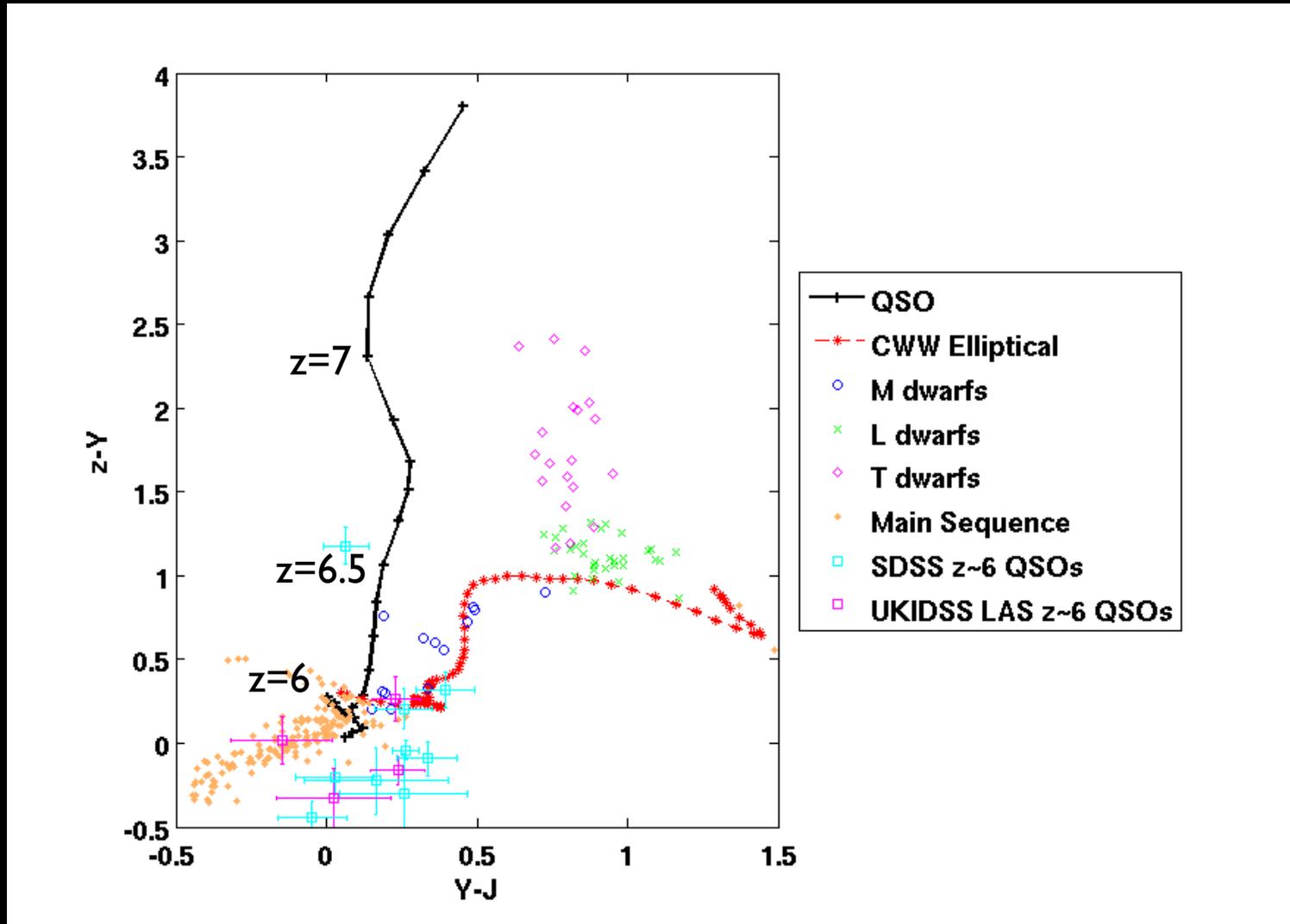
High-z QSO Selection



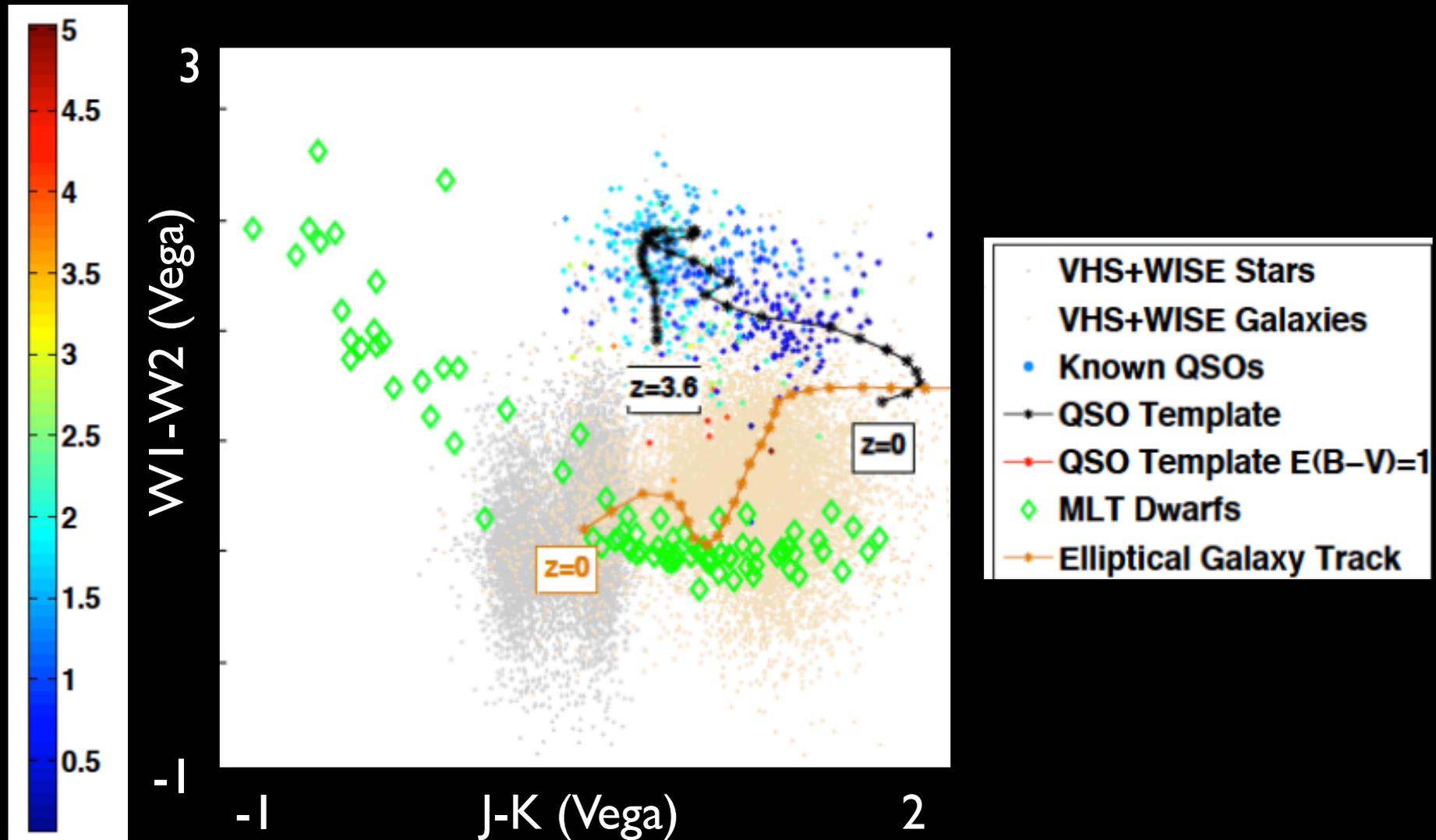
QSO too faint
to be detected
in the i-band



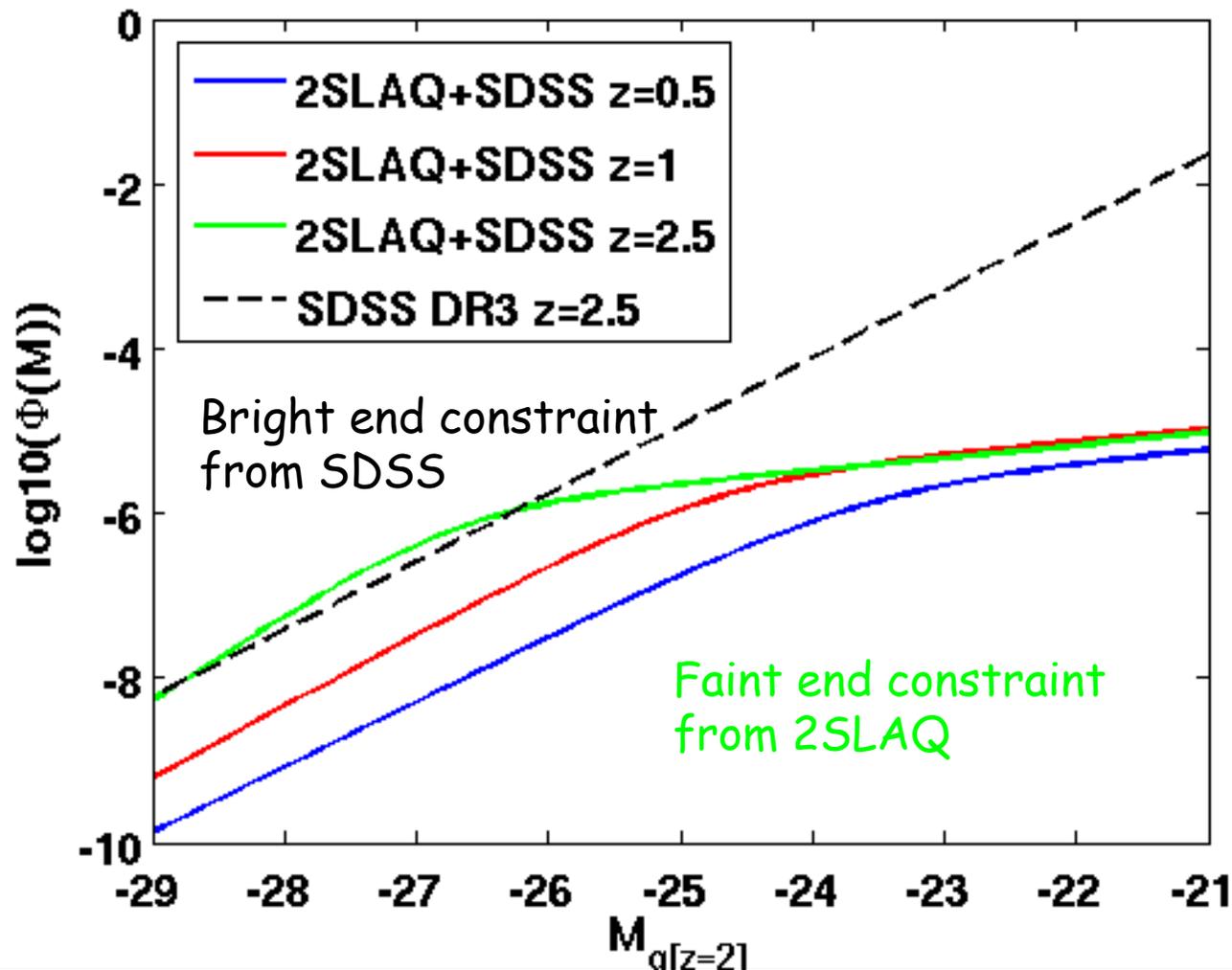
High- z QSO Selection



WISE Selection

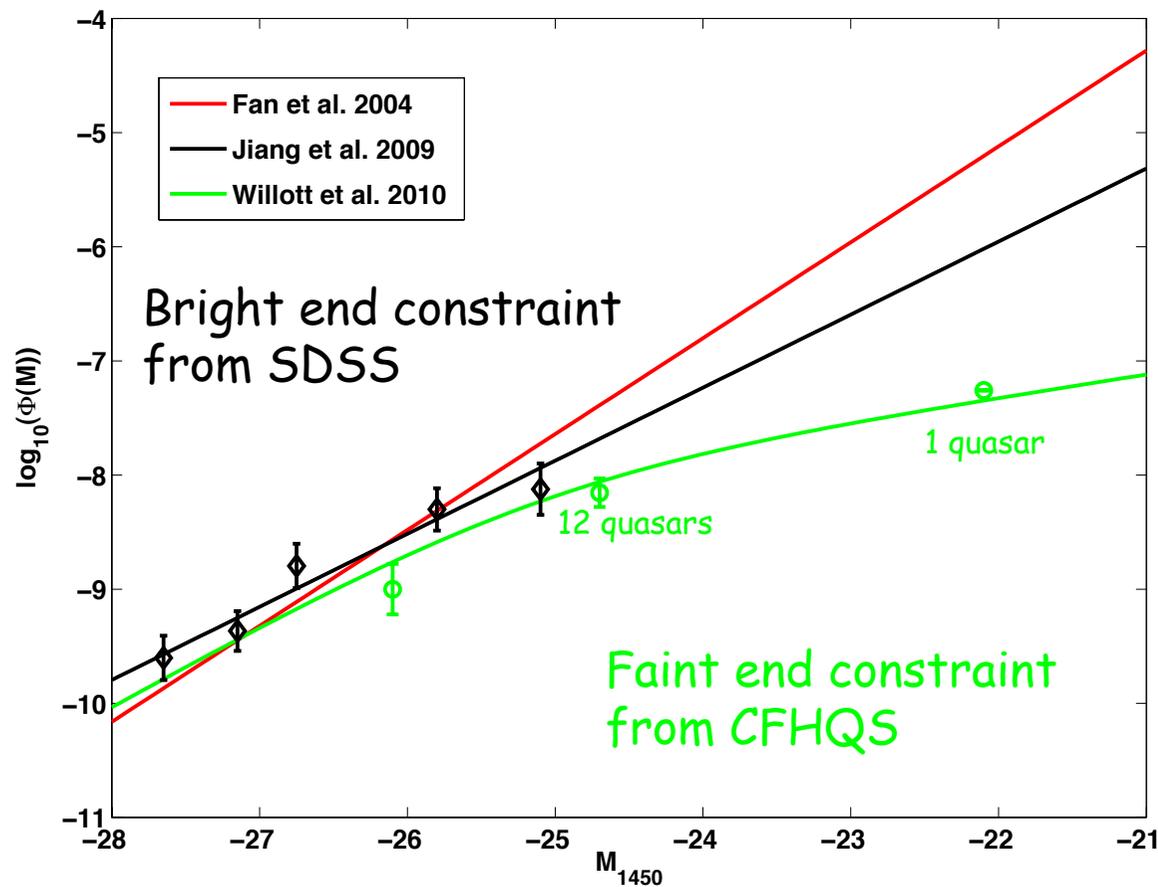


The QSO Luminosity Function at $0.5 < z < 2.5$



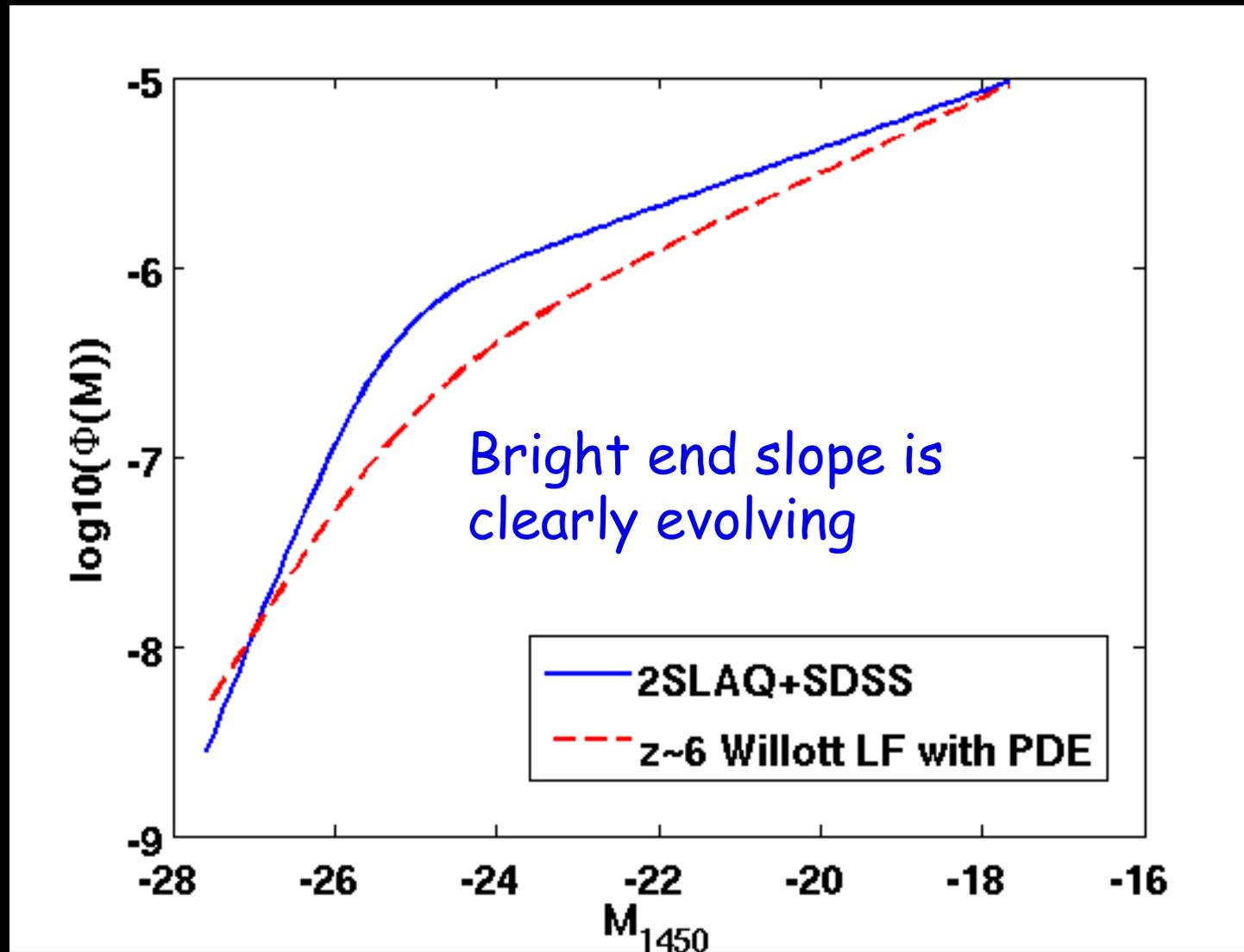
Croon et al. 2009

The QSO Luminosity Function at $z \sim 6$



Willott et al (2010) LF used to generate QSOs above redshift 5.5 assuming pure density evolution

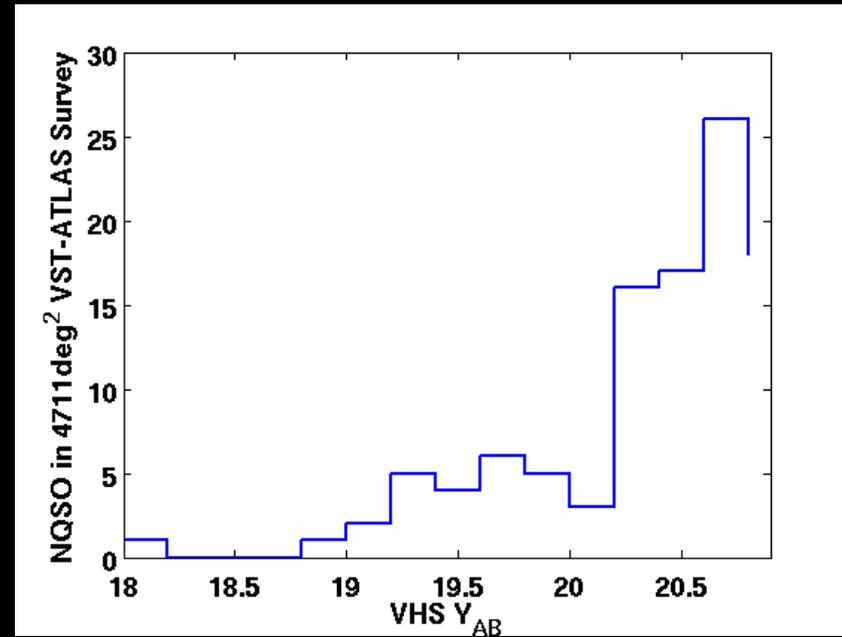
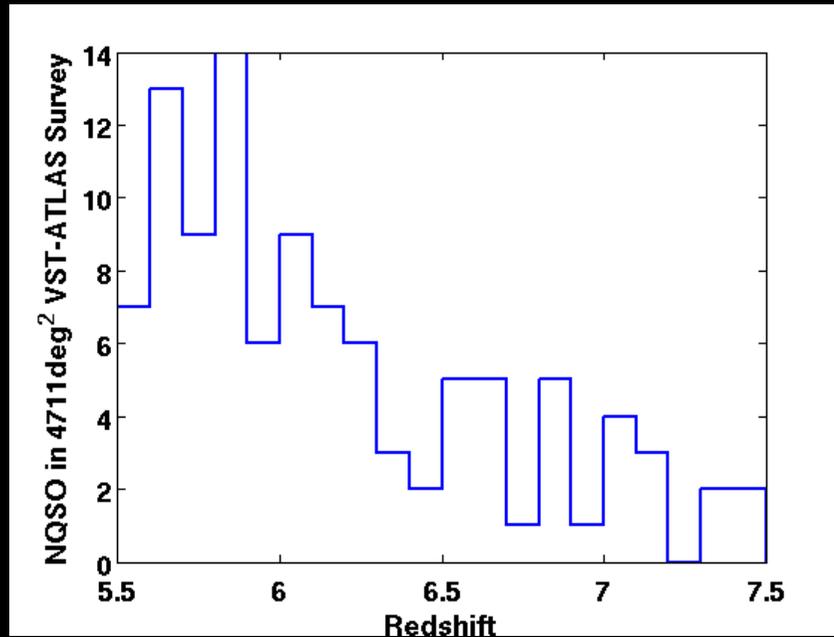
Intermediate redshifts e.g. $z=3$



High- z QSOs

- Assume Willott et al. LF with **pure density evolution** for $5.5 < z < 7.5$ in 20 bins of width 0.1
- Simulate over **4711 deg²**
- Apply flux limit in VHS Y-band (20.9 AB)

High- z QSOs

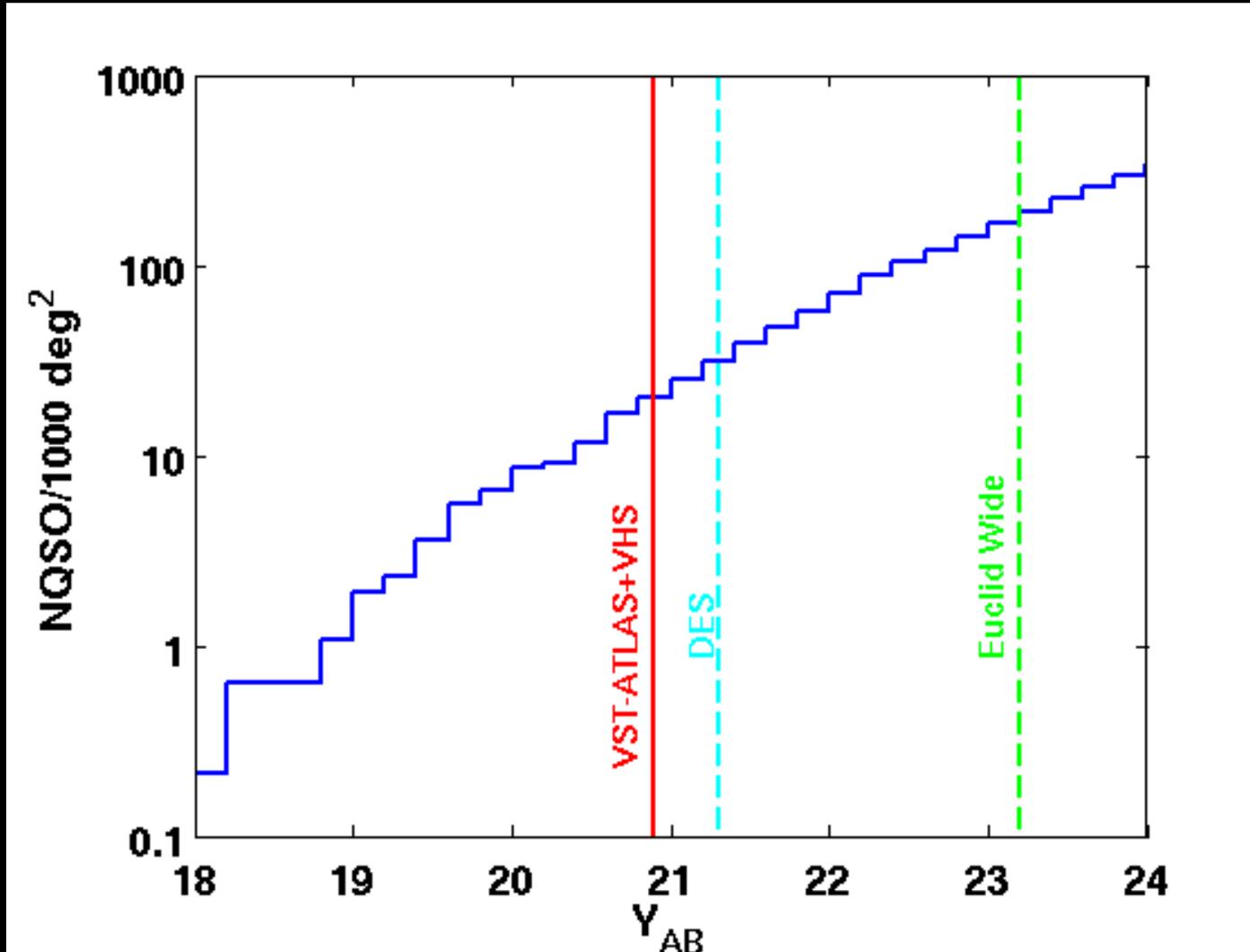


Assuming VHS Y-band limit of 20.9 (AB):

- 104 high- z QSOs with $5.5 < z < 7.5$ over 4711 sq deg
 - 20 at $z > 6.5$
 - 8 at $z > 7$

Predicted Surface Densities

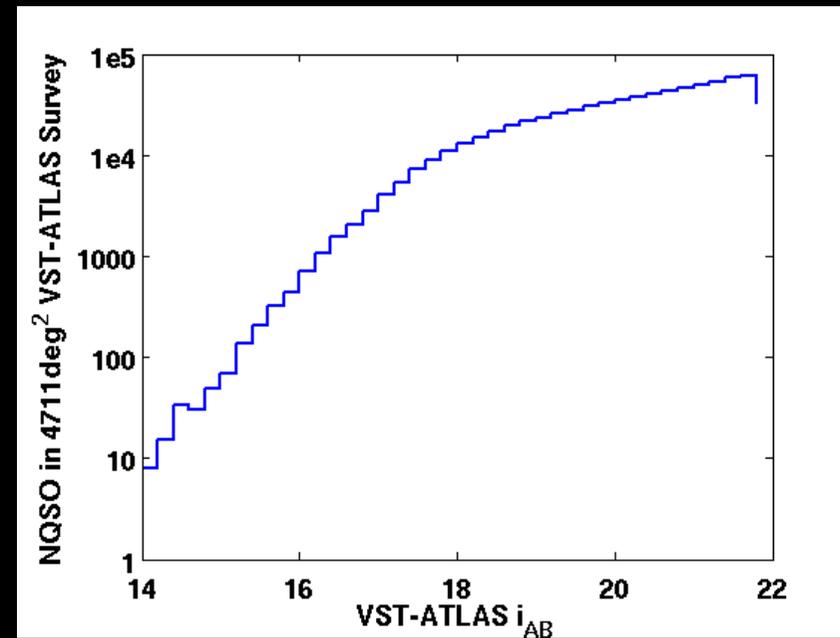
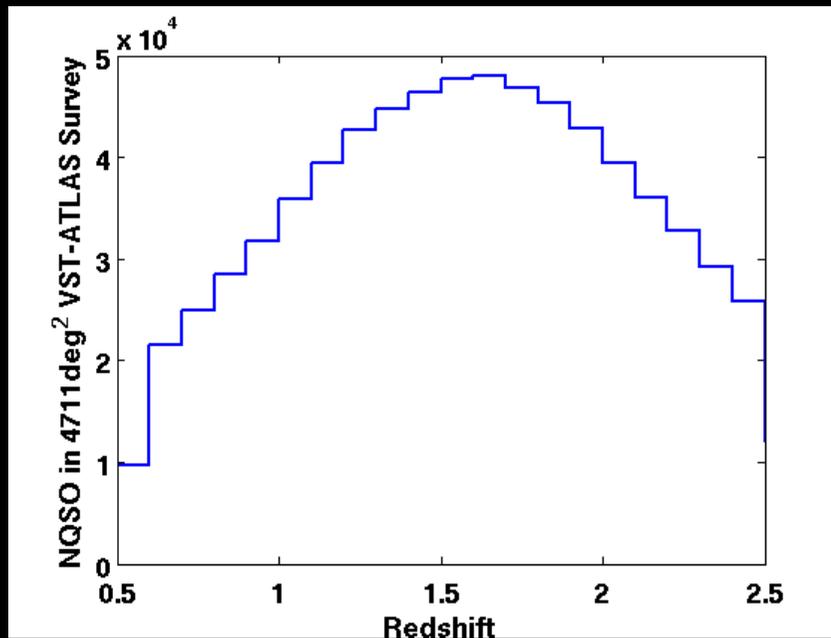
$5.5 < z < 7.5$



Low-z QSOs

- Assume Croom et al. LF with **luminosity dependent density evolution** for $0.5 < z < 2.5$ in 20 bins of width 0.1
- Simulate over **4711 deg²**
- Apply flux limit in **VST-ATLAS i-band** (**21.8, 10 σ AB**)

Low-z QSOs

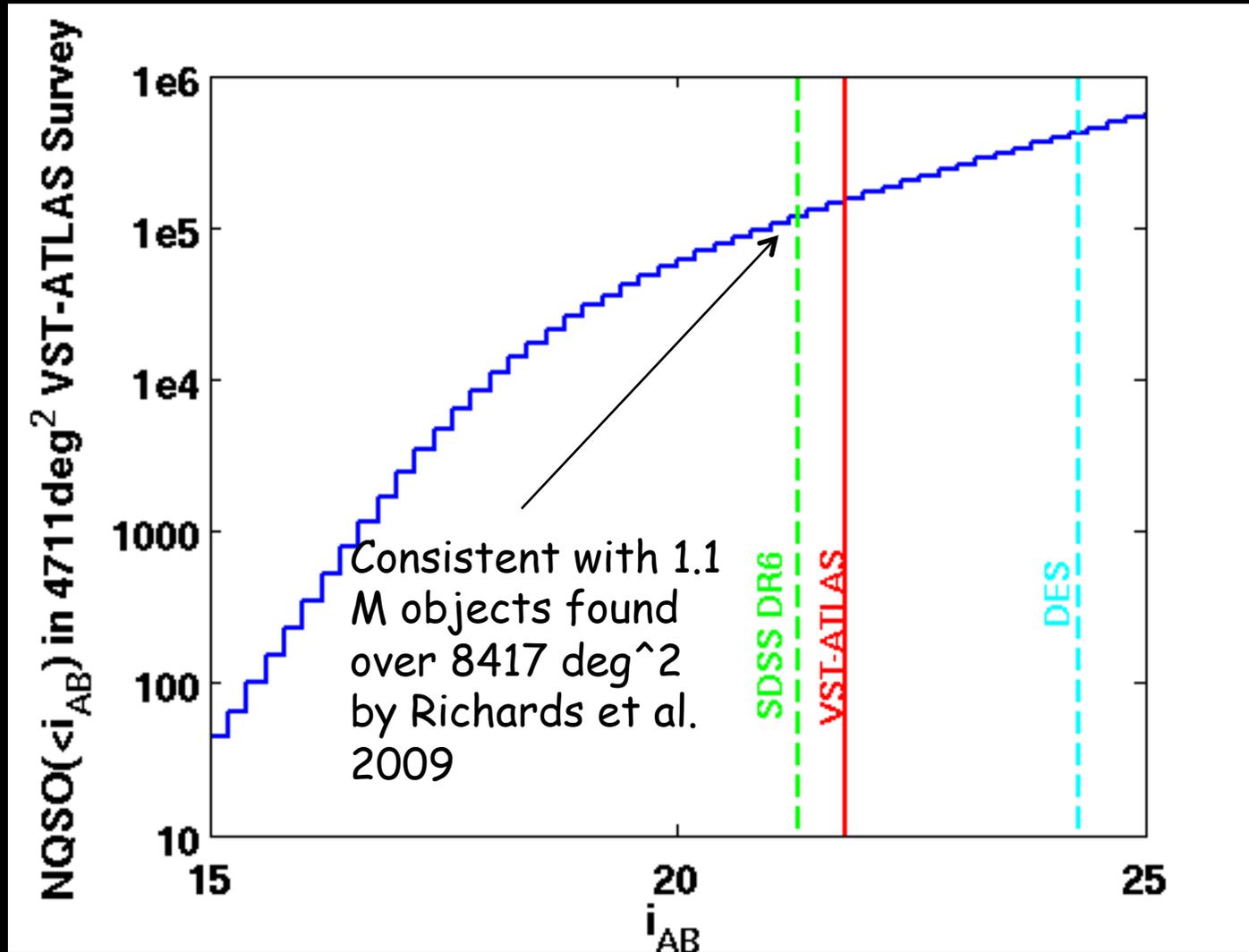


Assuming a VST-ATLAS i-band limit of 21.8 (AB)

- **731, 808 quasars** with $0.5 < z < 2.5$ over 4711 sq deg
 - Redshift distribution **peaks at $z \sim 1-2$** - interesting epoch for both galaxy formation and cosmology

Predicted Surface Densities

$0.5 < z < 2.5$



Conclusions

- VST-ATLAS + VHS will be a rich dataset for QSO science:
 - Low- z QSOs for cosmology
 - Constraints on the QSO LF at intermediate redshifts
 - The highest redshift QSOs suitable for detailed follow-up, studies of the IGM and epoch of reionization