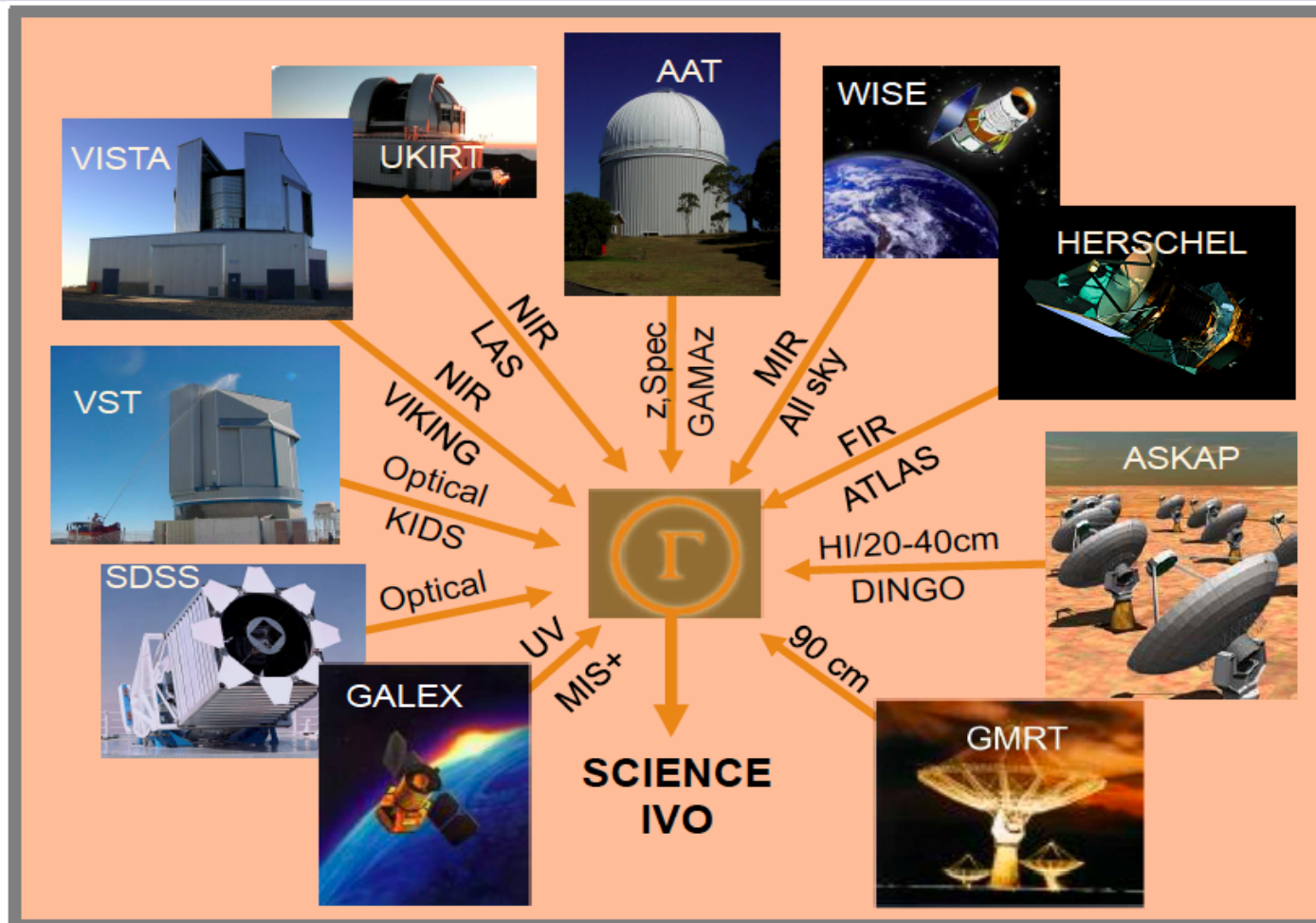


Galaxy And Mass Assembly survey and VST ATLAS/KIDS



Peder Norberg

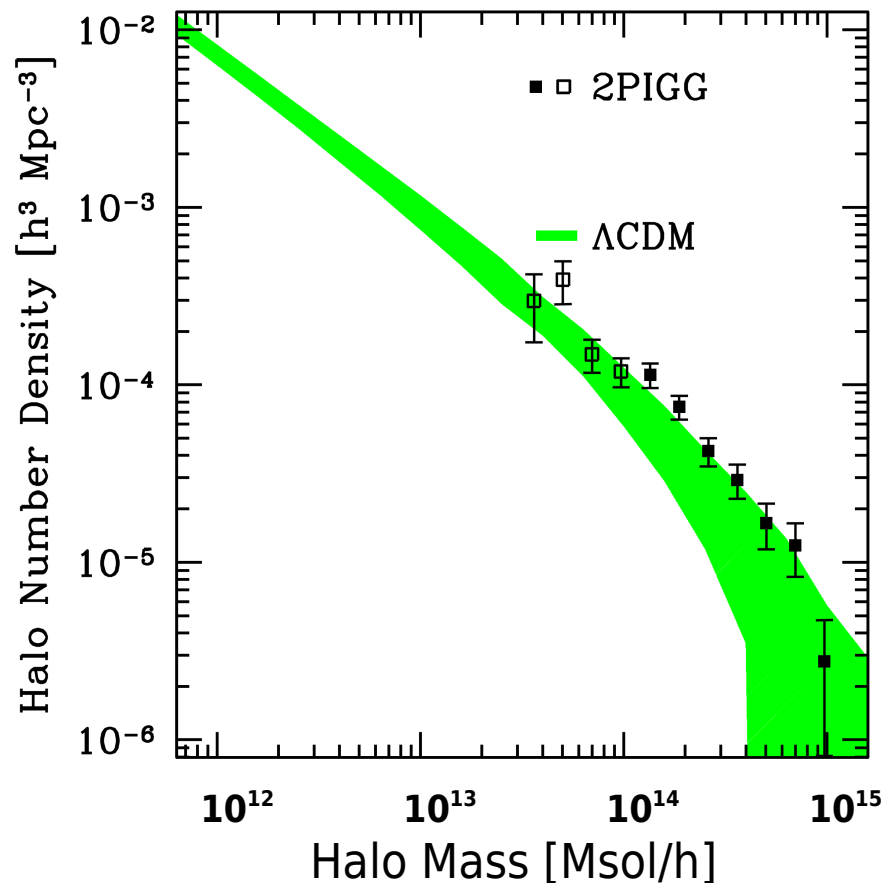
Institute of Computational Cosmology, Durham University

Many thanks to The Royal Society & the ERC for their financial support

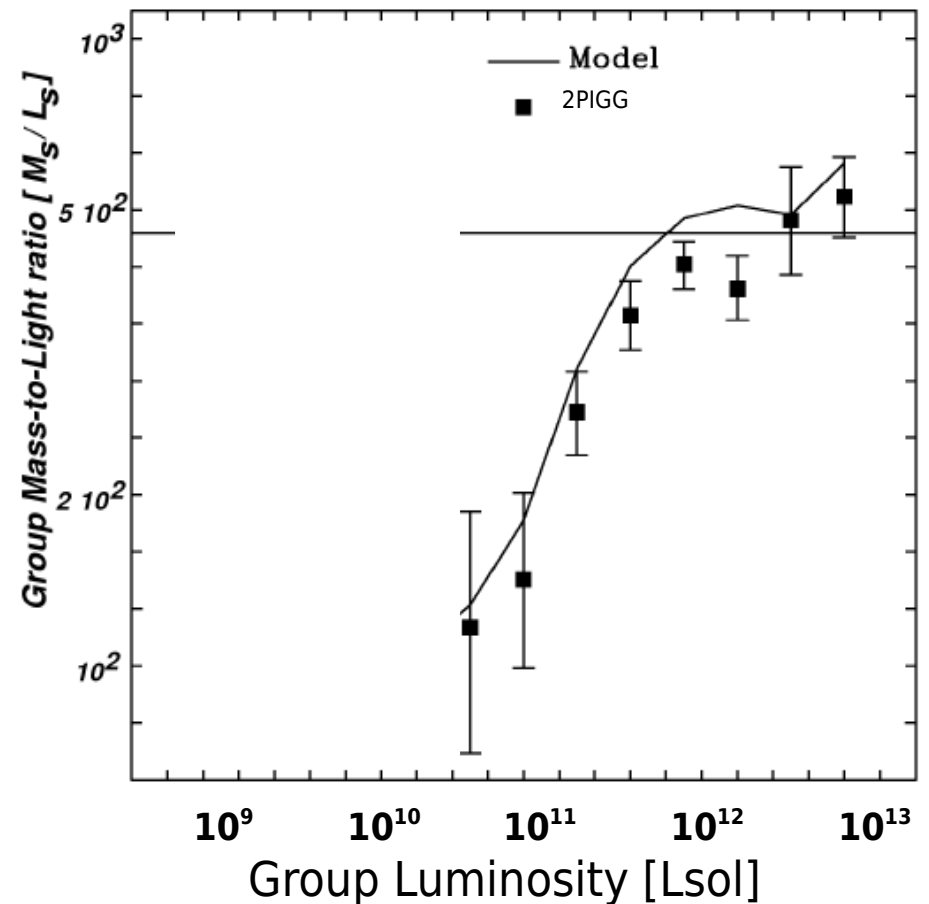
GAMA: la raison d'être

(with predictions from semi-analytic galaxy formation models)

Dark Matter Halo Mass Function



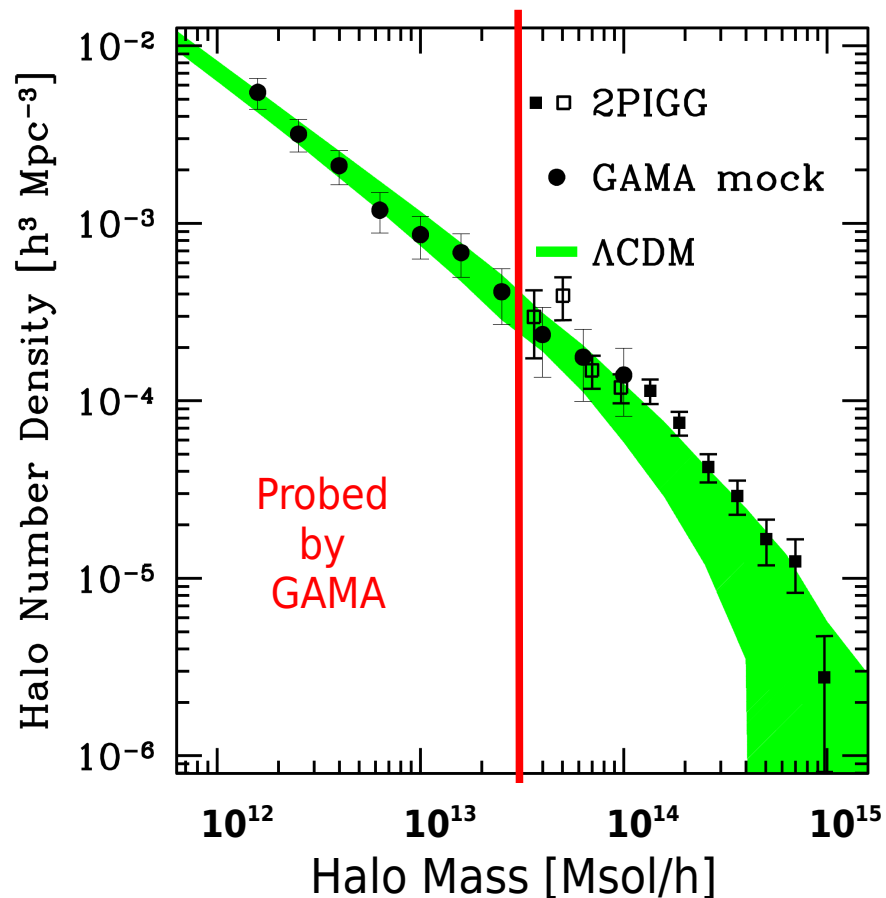
Galaxy Formation Efficiency



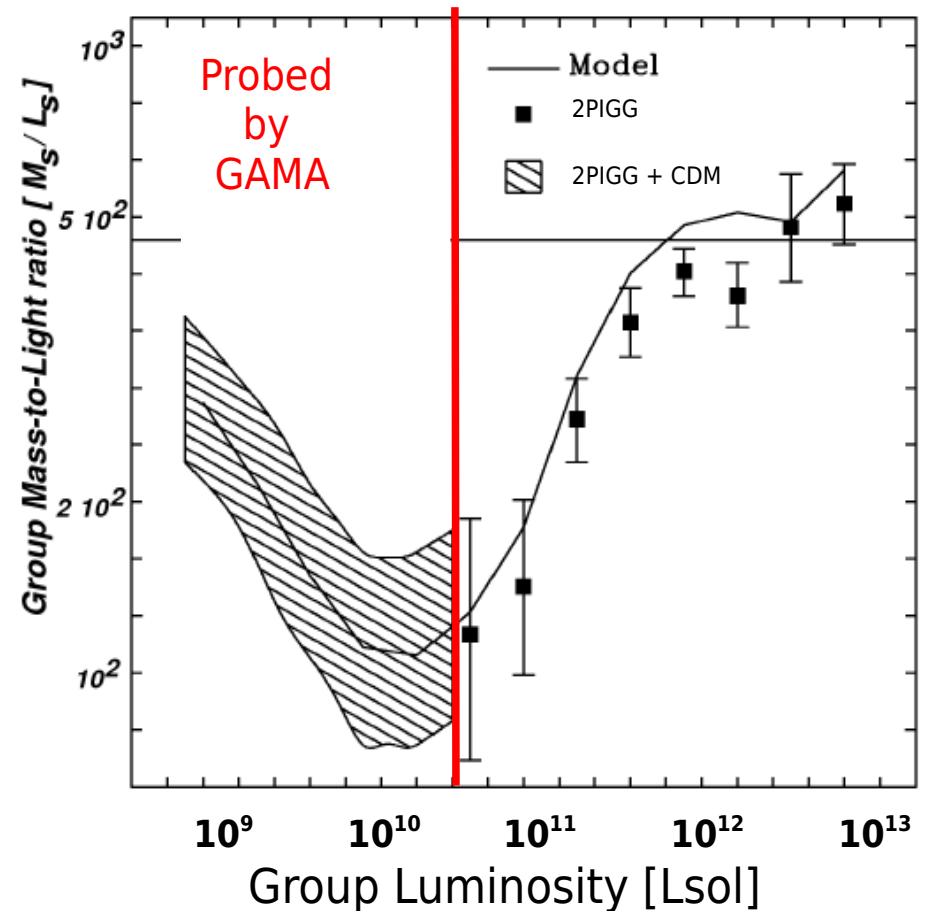
GAMA: la raison d'être

(with predictions from semi-analytic galaxy formation models)

Dark Matter Halo Mass Function



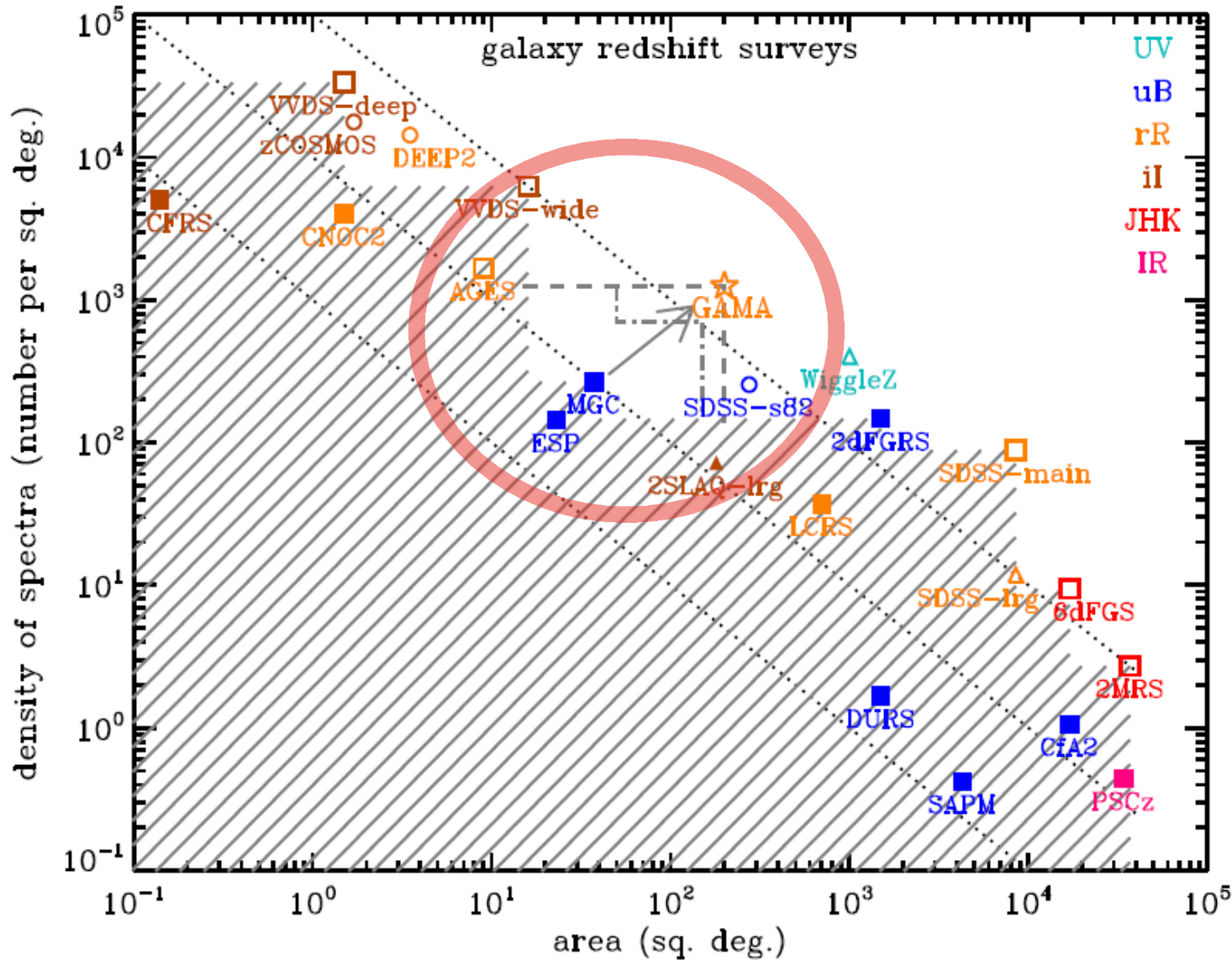
Galaxy Formation Efficiency



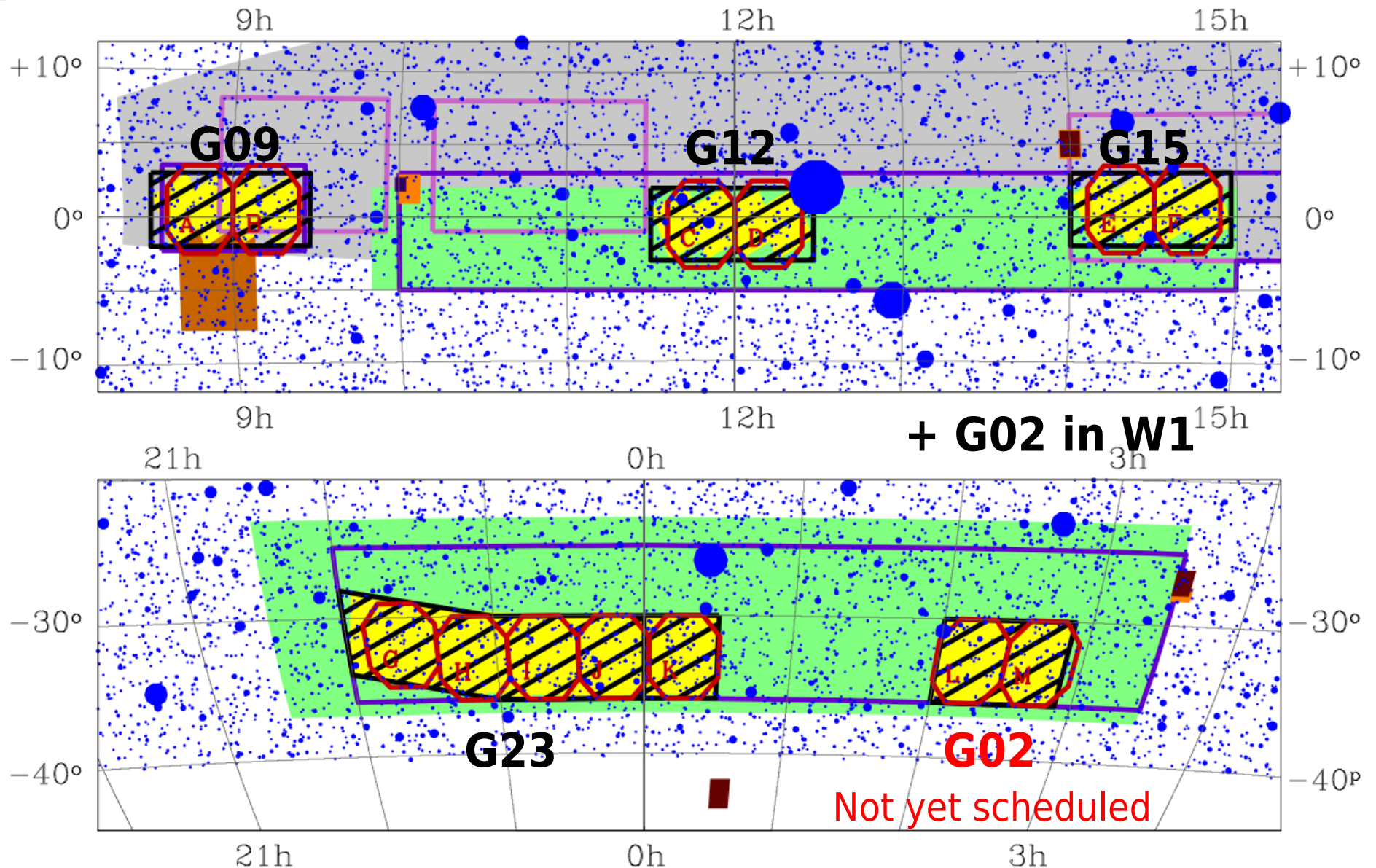
Galaxy And Mass Assembly Survey:

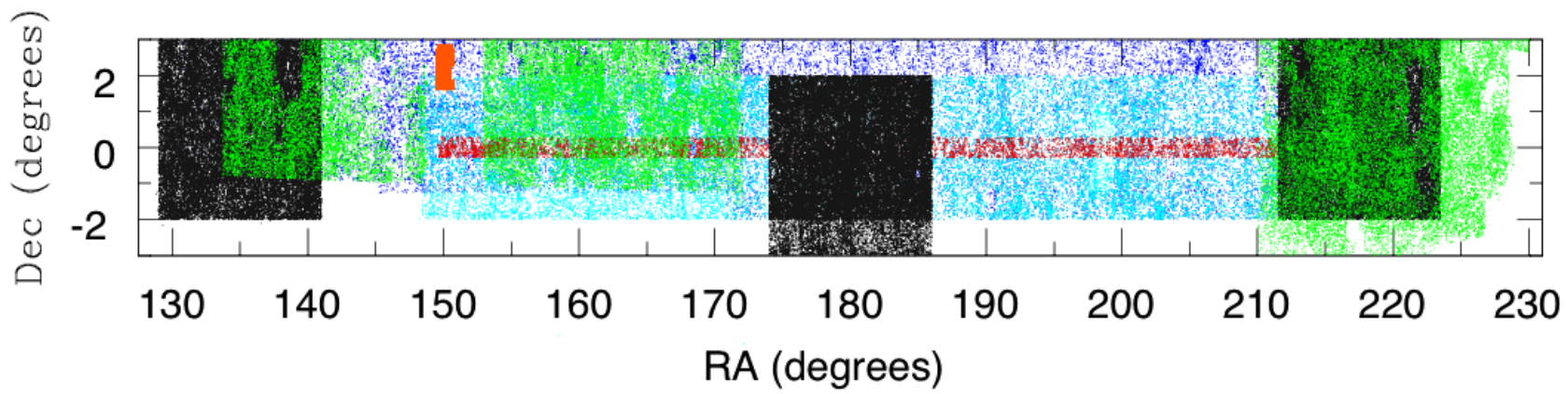
- Next generation galaxy redshift survey:
 - ~400,000 galaxy spectra to $r_{AB} \sim 19.8$:
 - 2 mag. fainter than SDSS => L^* at $z \sim 0.30$ [~ 4 Gyr]
 - ~360 sq. deg. wide, overlapping with SDSS and 2dFGRS
 - >150 nights on AAT with AAOmega (2008-2012)
 - GAMA is also K-band limited, with $K_{AB} < 17.6$
- GAMA is a unique survey and fills an essential gap in the current generation of redshift surveys, between the very wide low-z and very narrow high-z.
- Already in the GAMA bag (GAMA-I):
 - >150k spectra with uniform coverage over 144 sq.deg.

Galaxy And Mass Assembly Survey: germane connection between shallow-wide & deep-narrow



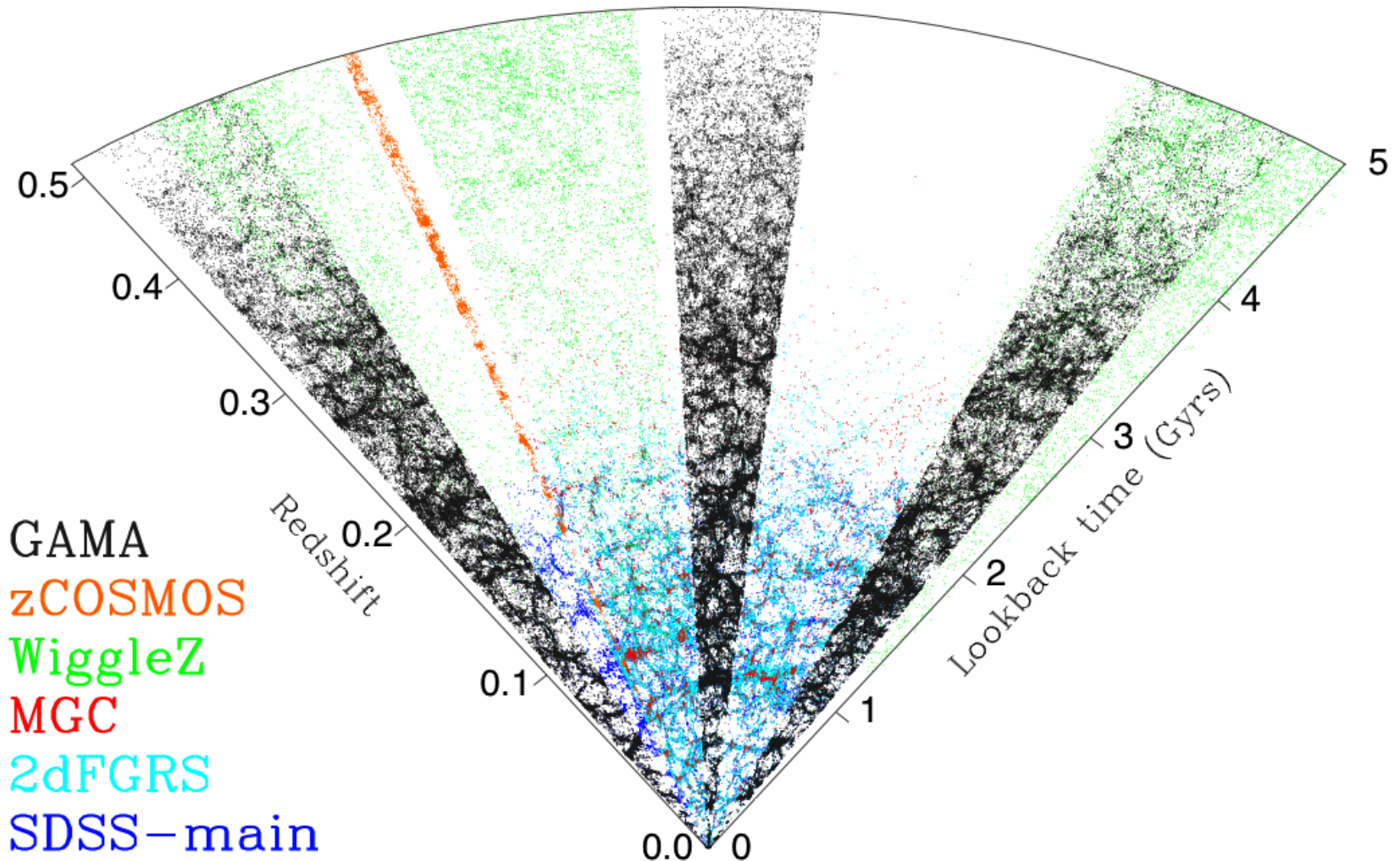
Location of the 5 (6) GAMA fields, each of ~70 sq.deg.





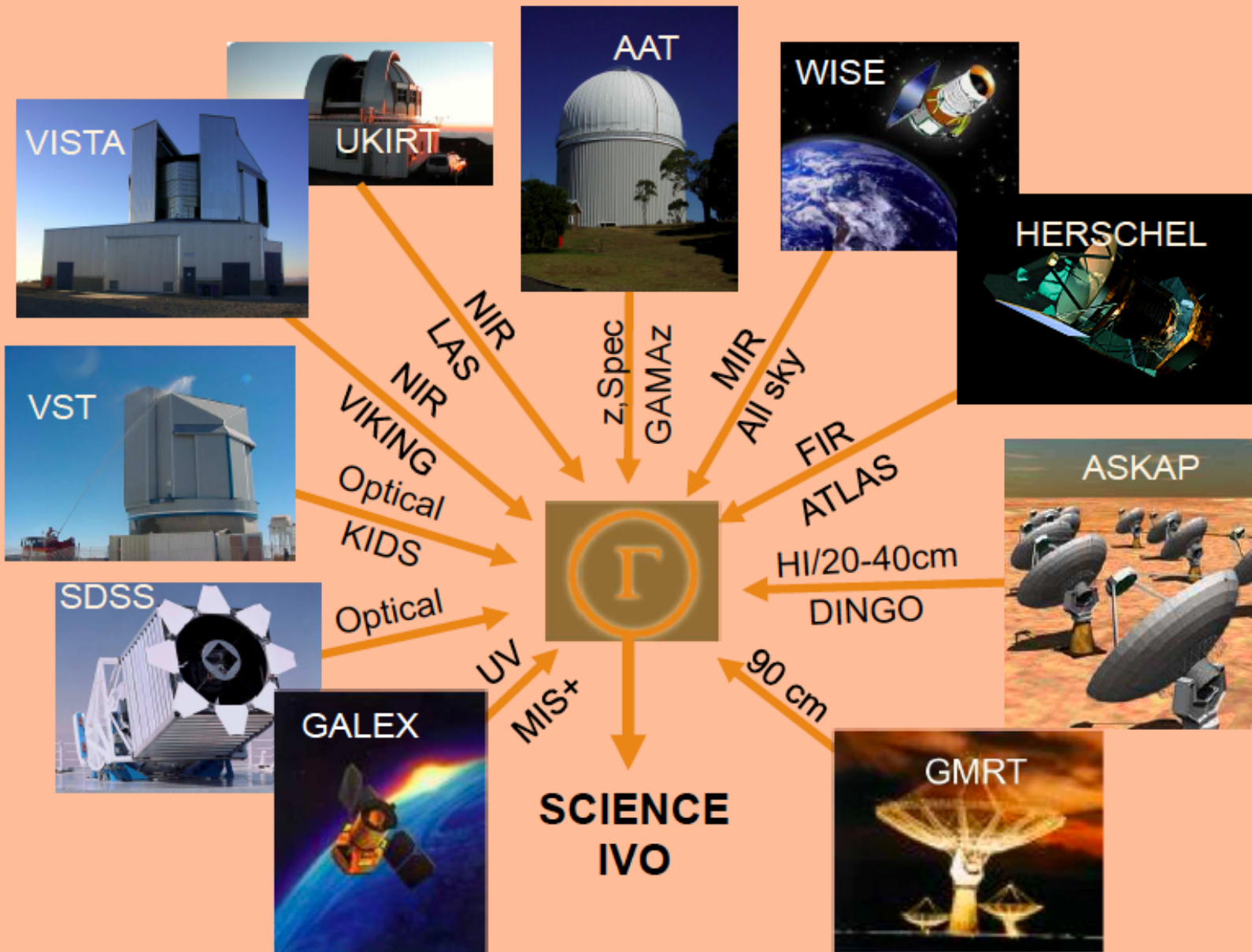
G
A
M
A
I

G09
G12
G15



GAMA
zCOSMOS
WiggleZ
MGC
2dFGRS
SDSS-main

GAMA: Contributing Facilities

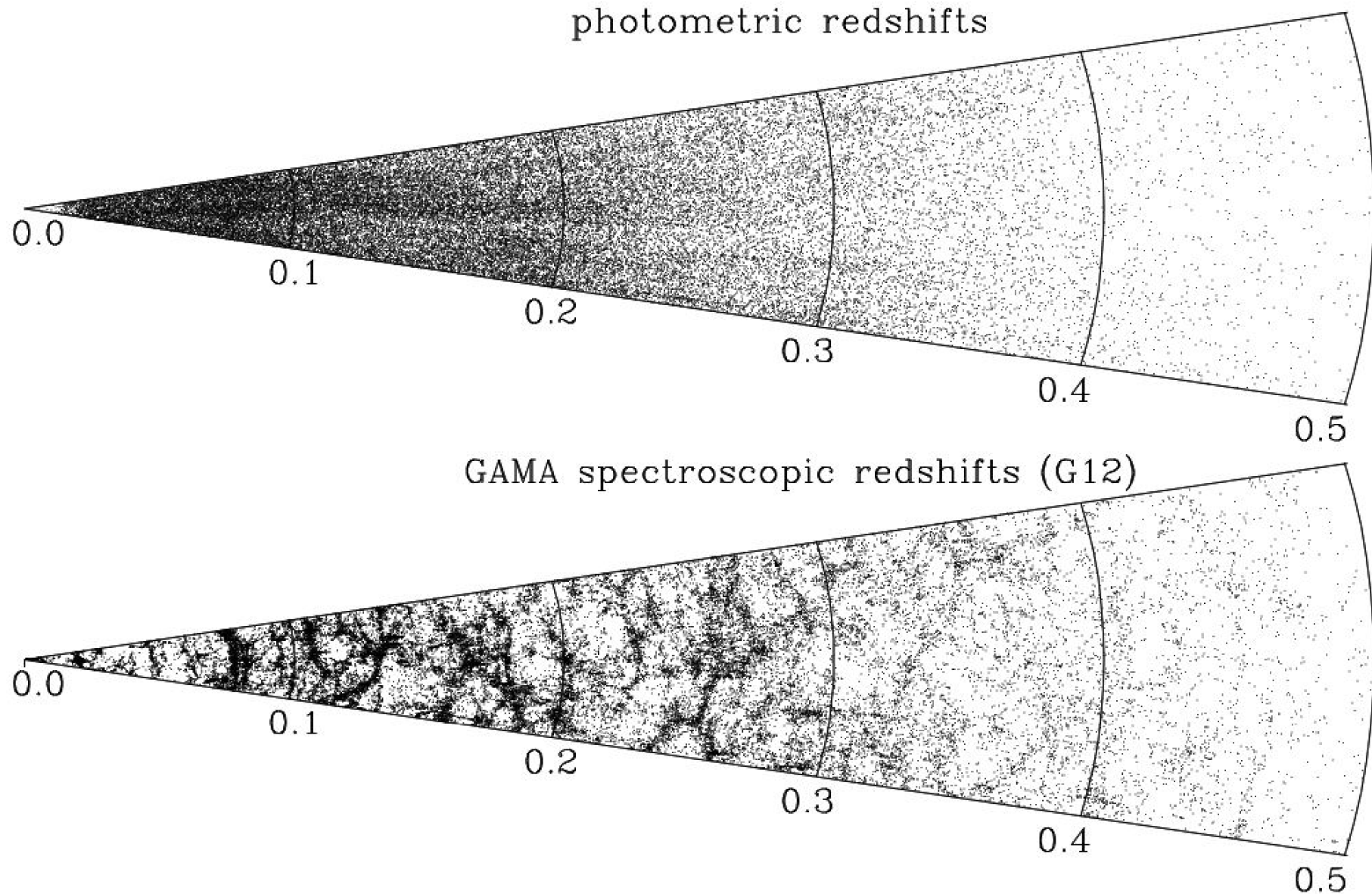


Recent addition:

G02 (W1)
with:
 - CFHTLenS
 - XMM-XXL

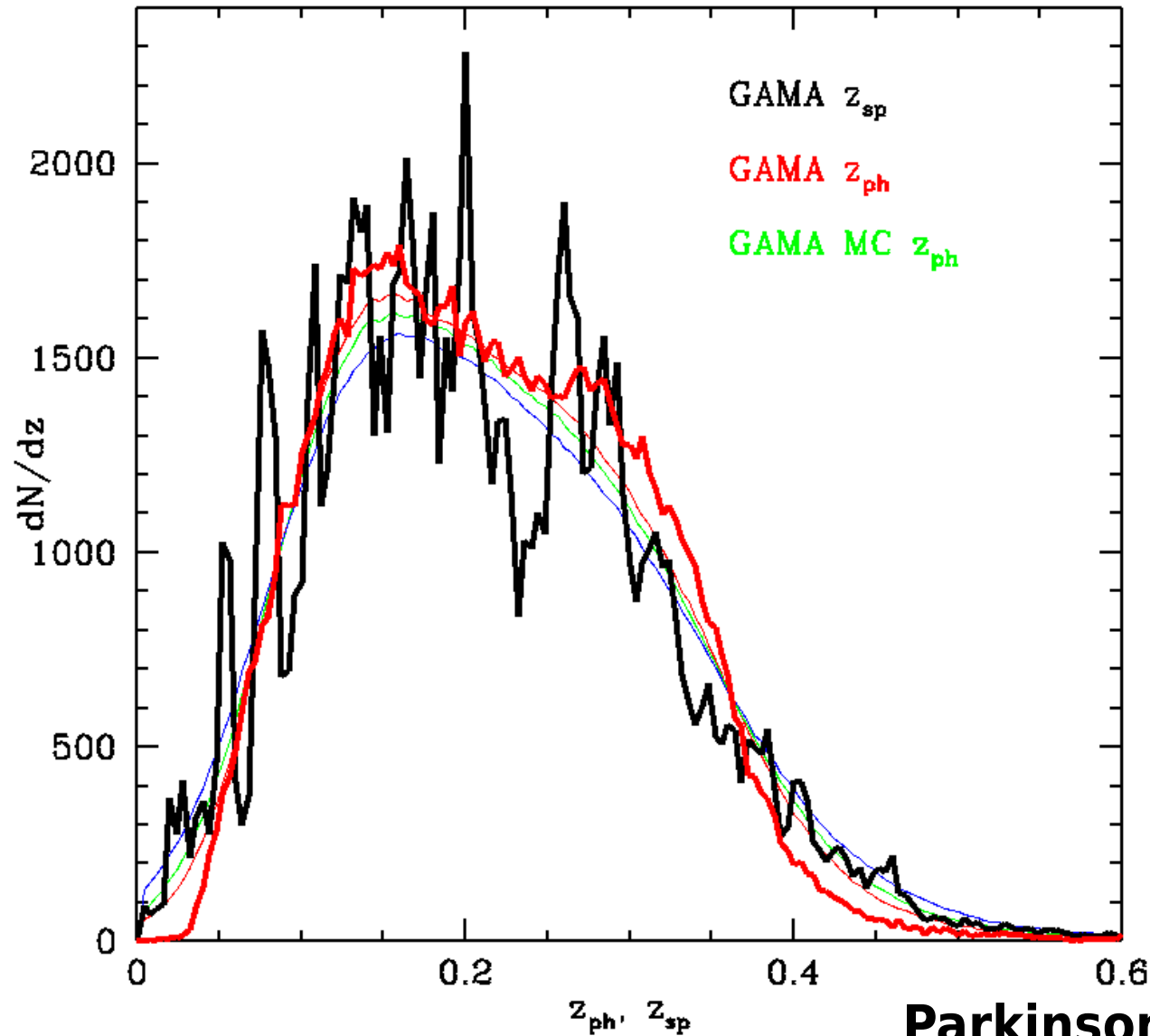
GAMA: Preliminary Results

improved photometric redshifts



GAMA: Preliminary Results

$N(z)$ for z_{spec} and z_{photo}

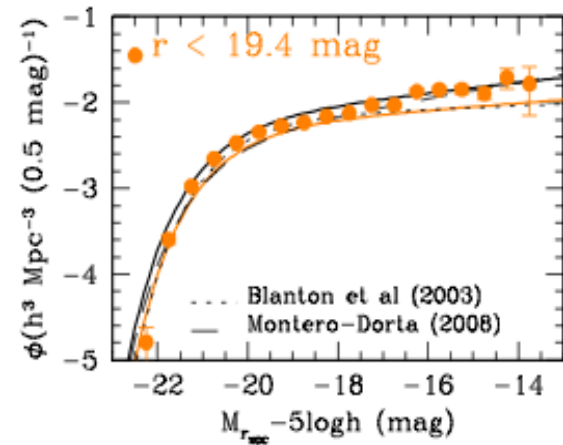
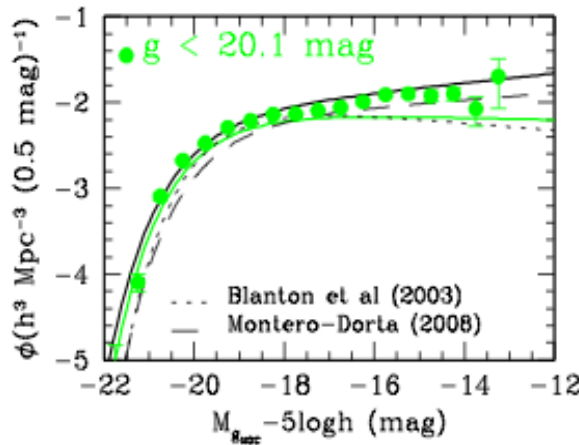
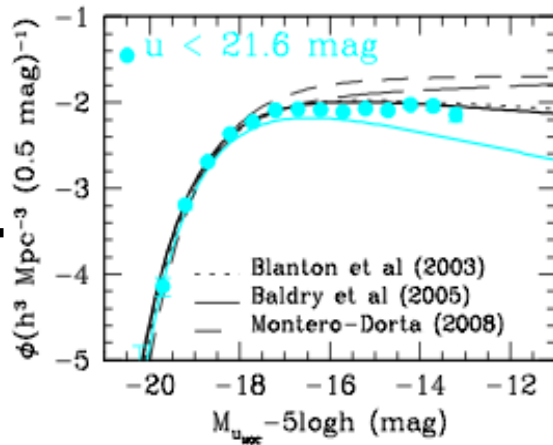
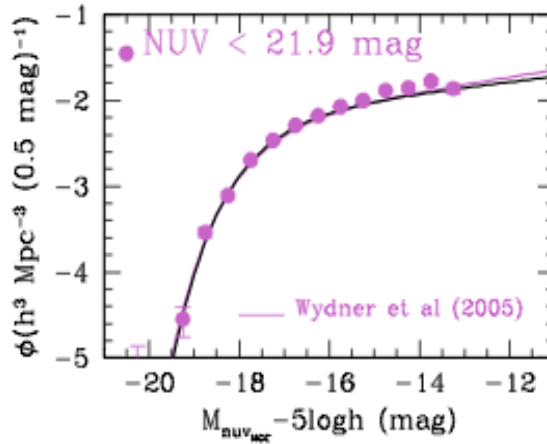
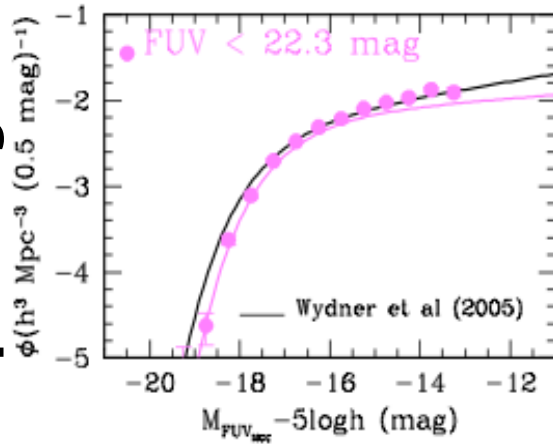


Parkinson et al. (in prep)

GAMA: Preliminary Results

11-band galaxy luminosity function ($z < 0.1$)

#galaxies per Volume per Magnitude



11 band luminosity functions to $z < 0.1$ over the common GAMA regions (115.14sq deg) to $r < 19.4$ corrected for colour bias and incompleteness

Bright

Abs. Mag.

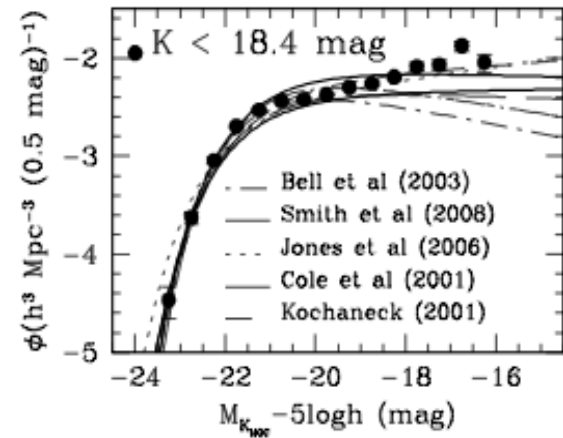
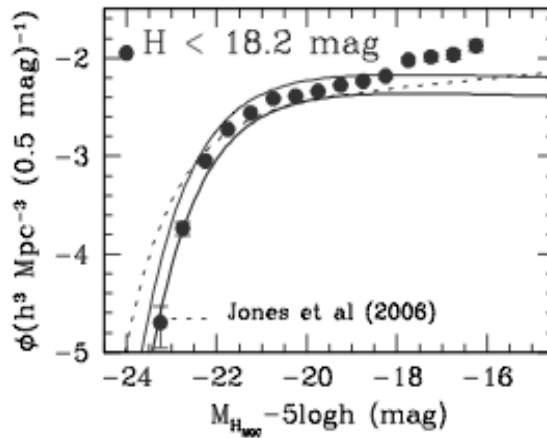
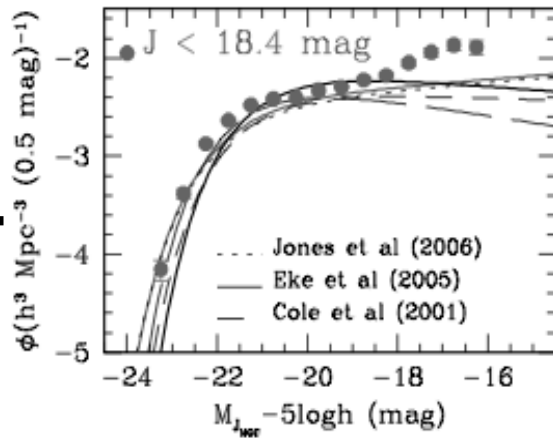
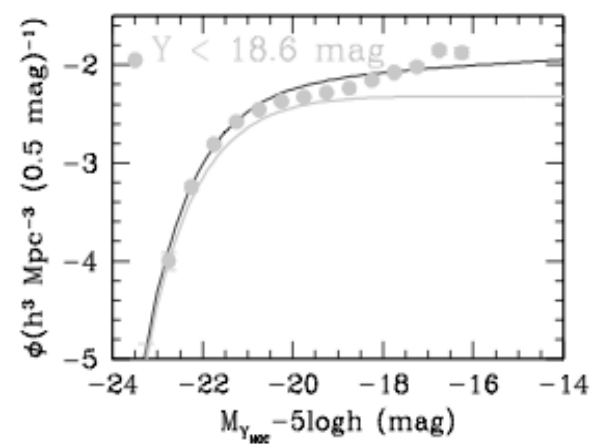
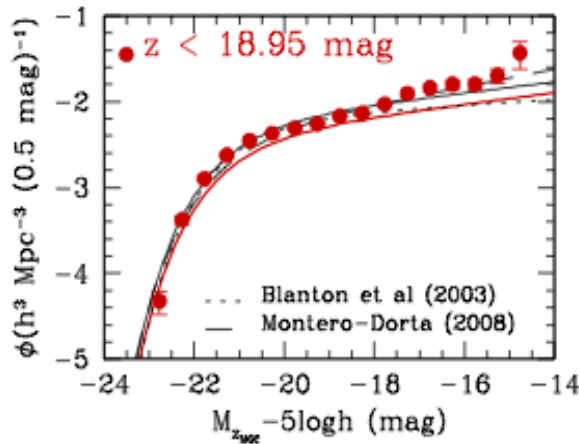
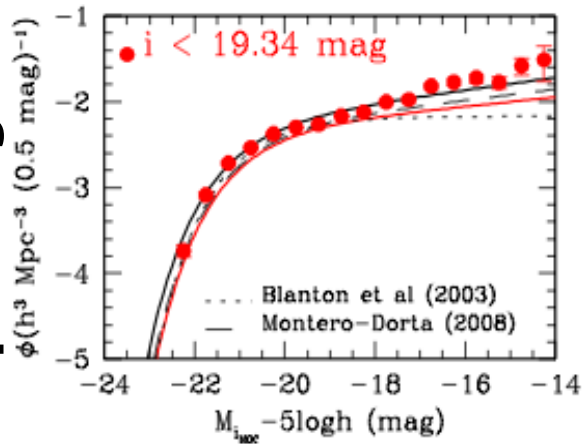
Faint

Driver et al. (in prep)

GAMA: Preliminary Results

11-band galaxy luminosity function ($z < 0.1$)

#galaxies per Volume per Magnitude



Bright

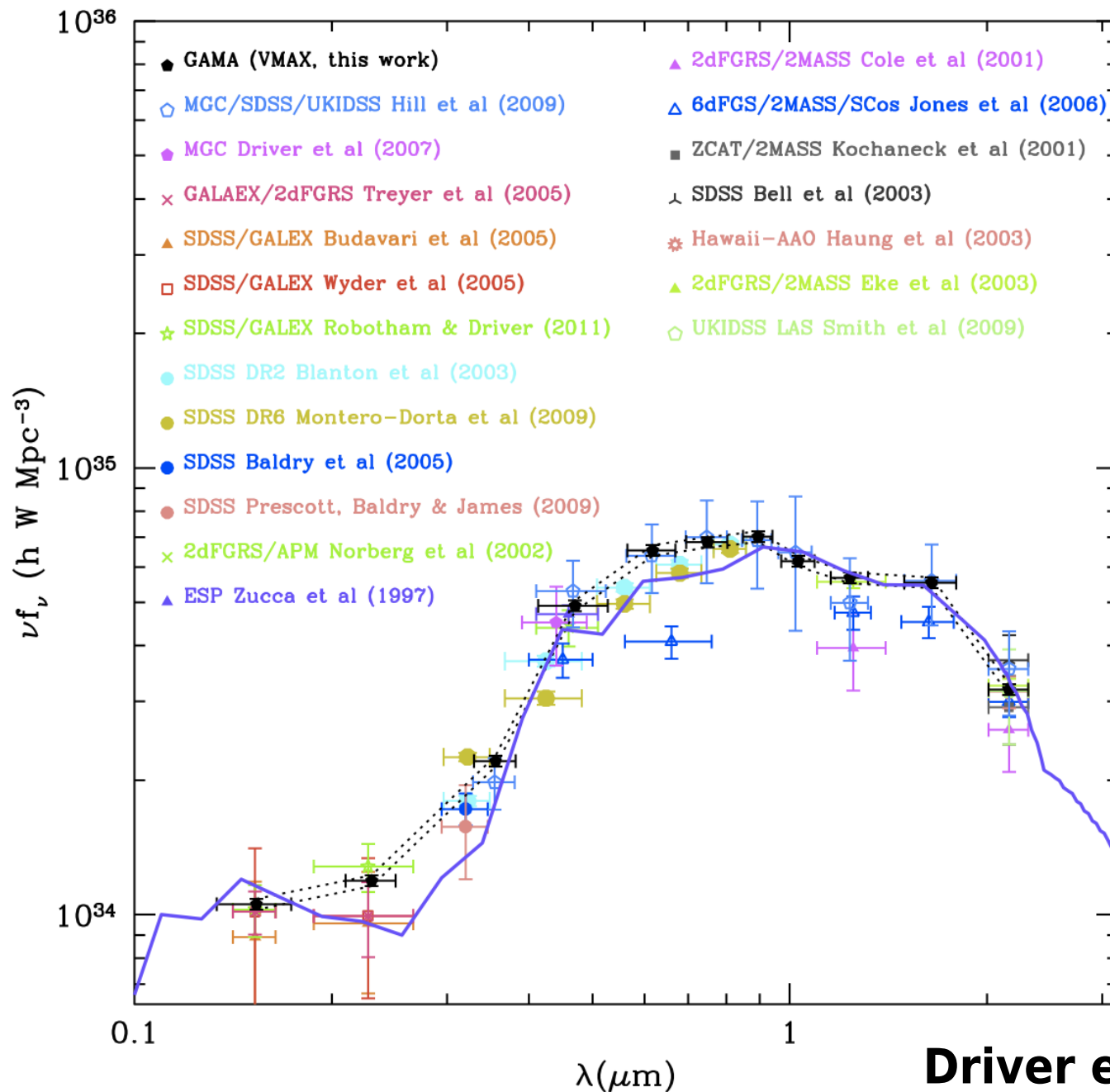
Abs. Mag.

Faint

Driver et al. (in prep)

GAMA: Preliminary Results

Cosmic Spectral Energy Distribution ($z < 0.1$)

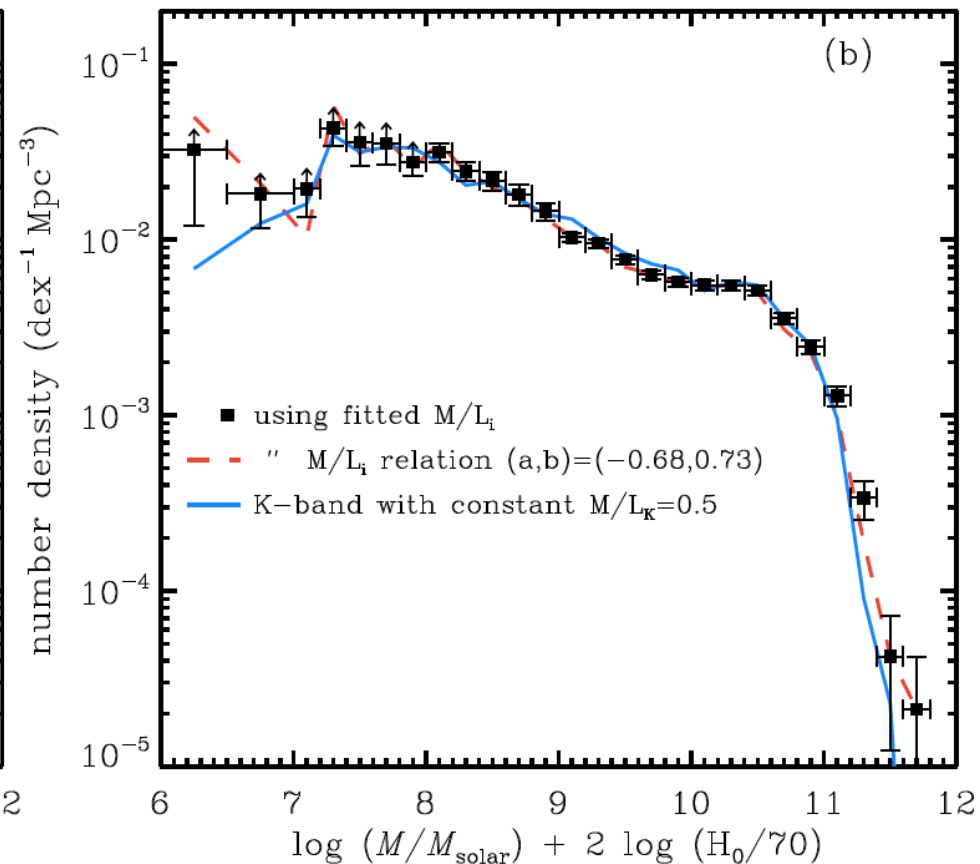
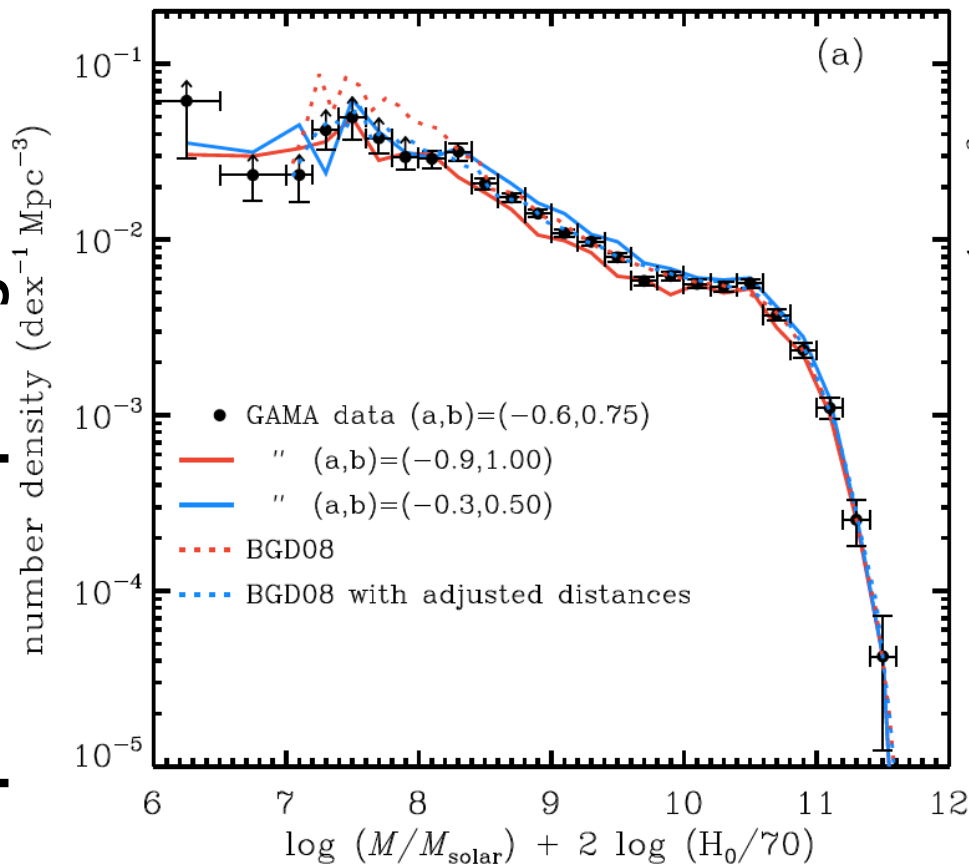


Driver et al. (in prep)

GAMA: Preliminary Results

stellar mass function ($z < 0.06$)

#galaxies per Volume per log Stellar Mass



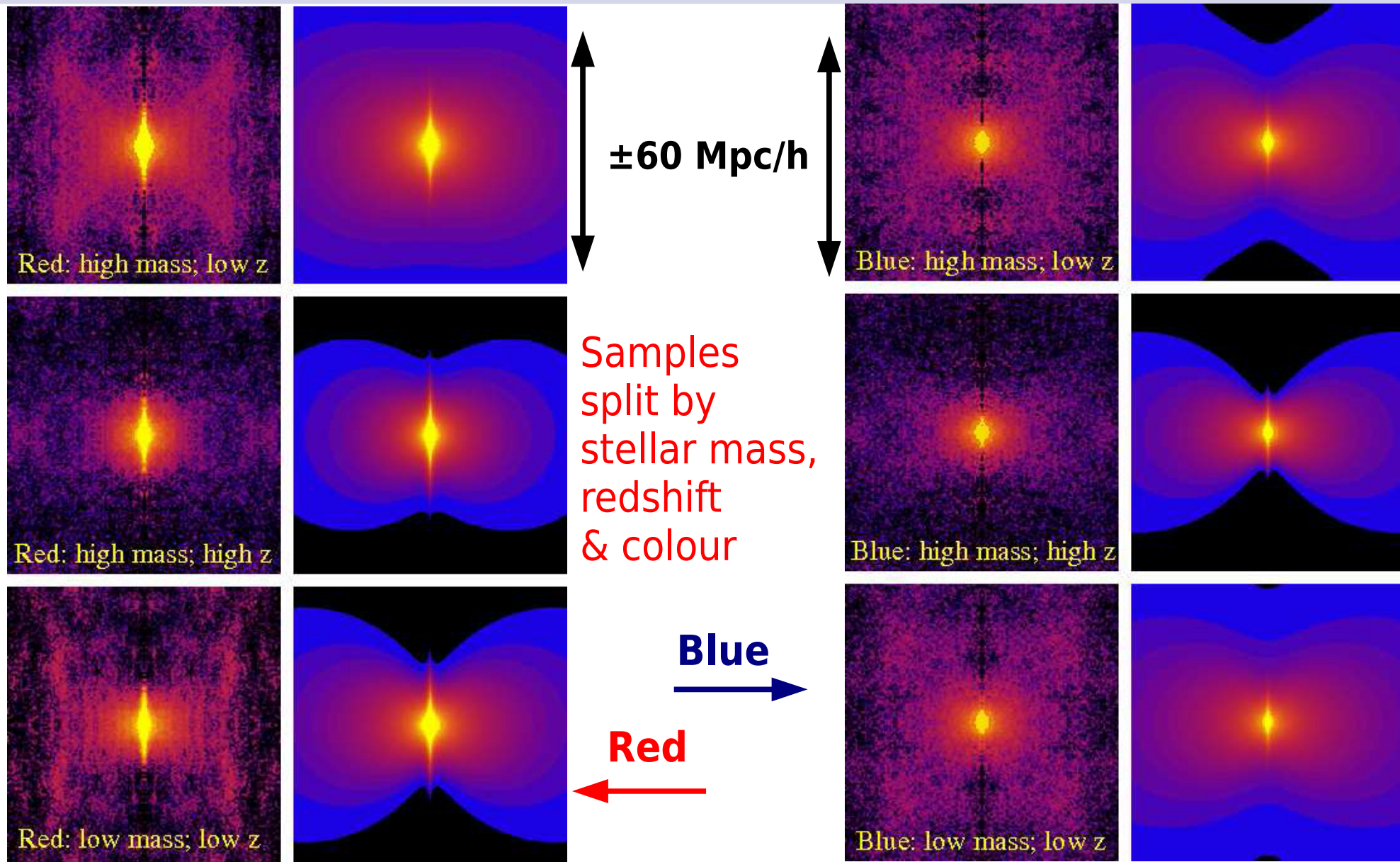
Low Stellar Mass High

$$\text{Log}_{10} (M/L)_i = a + b (g-i)$$

Baldry et al. (in prep)

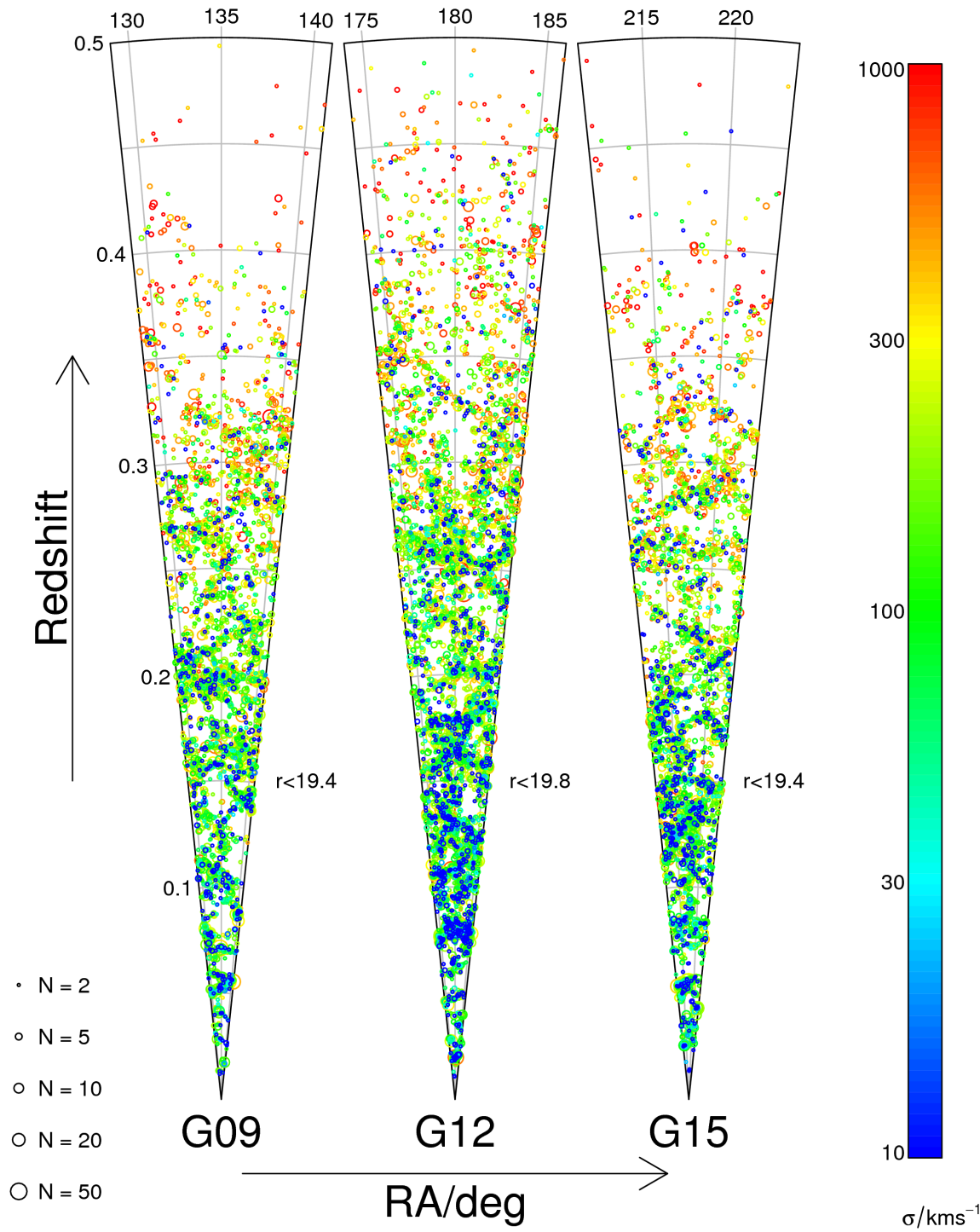
GAMA: Preliminary Results

Clustering & Redshift Space Distortions



GAMA: G³Cv1 coneplot

- 142 sq.deg. to $r=19.4$
- 12k $N_g \geq 2$ groups
- 1.5k $N_g \geq 5$ groups



Robotham et al. (2011)

GAMA: THE DATABASE (I)

All (~250k):

General: GAMA ID : SDSS ID : z (heliocentric) : z quality

Flux: UV : optical : near-IR : mid-IR : far-IR : Radio (20,rest-21,30,40,90cm)

Shape: CAS : Sersic index: half-light radii : b/a : PA in *ugrizYJHK*

Opacity: $\tau_{UV,ugriz,YJHK}$

Spectral features: Emission: $H\alpha, H\beta, H\gamma, H\delta, OII, OIII, NII$

Abs.: $Dn4000, Ca4227, H\alpha, H\beta, H\gamma, H\delta, Mgb, Fe$

SFR: UV : $H\alpha$: far-IR : Radio continuum

Fossil record: Age : SFH : element abundance

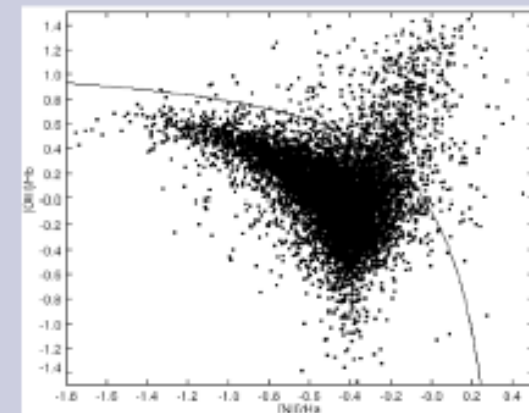
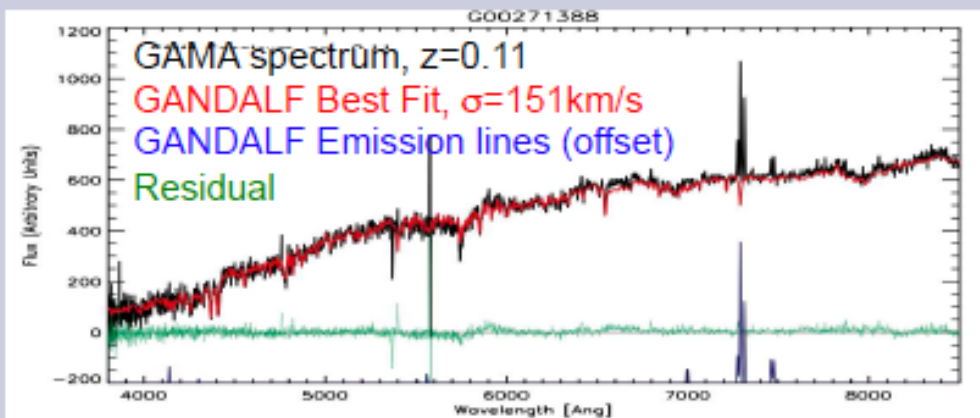
AGN: BPT diagnostics : type : strength : ionisation state

Dynamics: σ_{spec} (GANDALF) : W_{21} : HI line profile

Distances: Tully-Fisher : Faber-Jackson

Masses: Stellar : SMBH : HI : Dust : Baryon : Dynamical

Environment/Halo: Local density : Group membership : Group halo mass



GAMA: THE DATABASE (II)

For $z < 0.1$ (~30k):

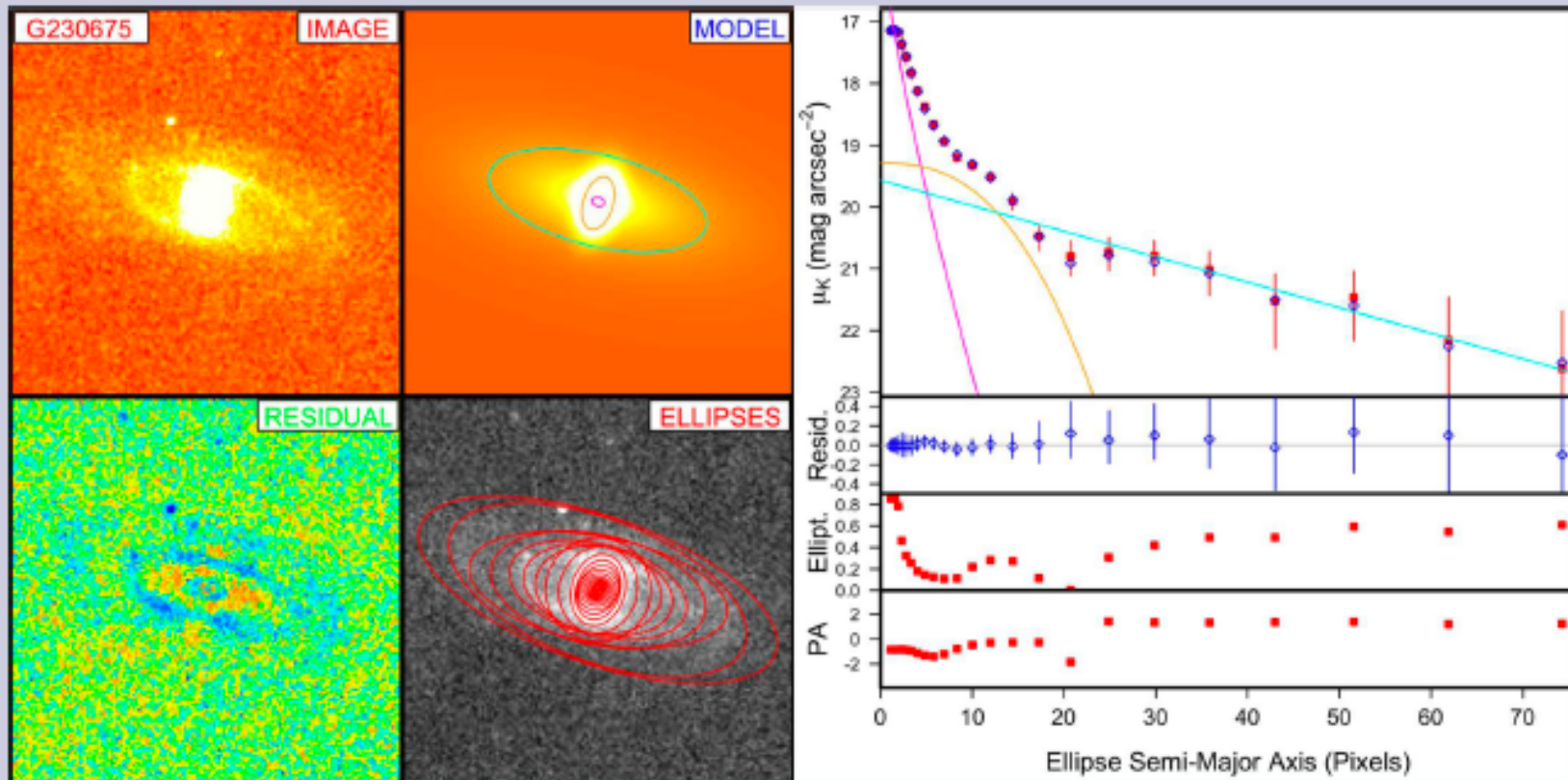
Structural: Bulge/Bar/Disc decomp. in *ugrizYJHK* (GALFIT3)

Bulge: Sersic index, half light radius, Pos. Angle, Ellipticity

Bar: Sersic index, half light radius, scale-length

Disc: Scalelength, PA, b/a

SMBH Mass: via M- σ , M-L, M-n relations



GAMA Team



<http://www.gama-survey.org/>