

# Implications for Reionization of a Search for Local Group Dwarfs with HI and Optical: Is There a Missing dIrr Problem?

**Erik Tollerud**

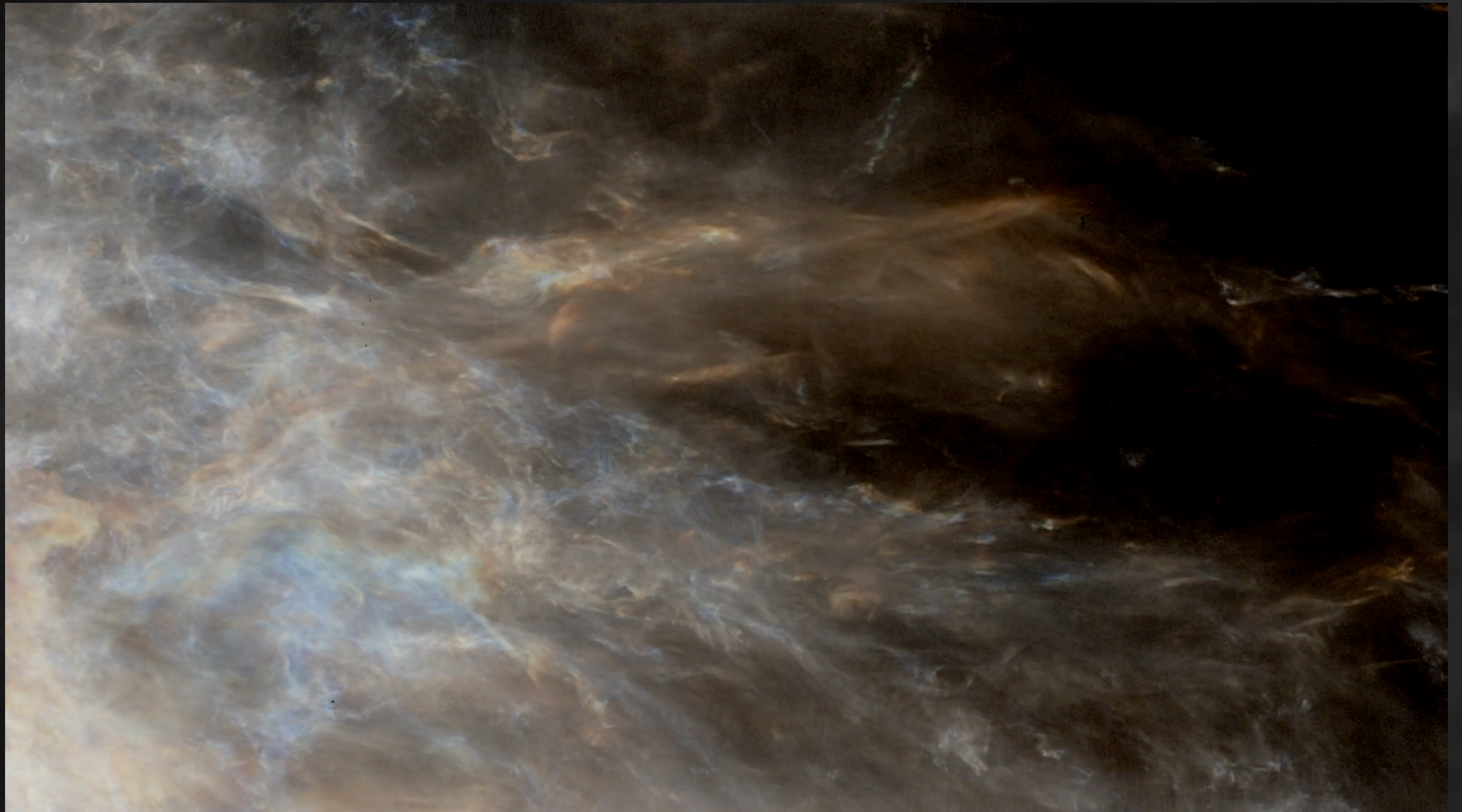
@eteq

Space Telescope Science Institute

w/ Josh Peek, Mary Putman, Jana Grcevich





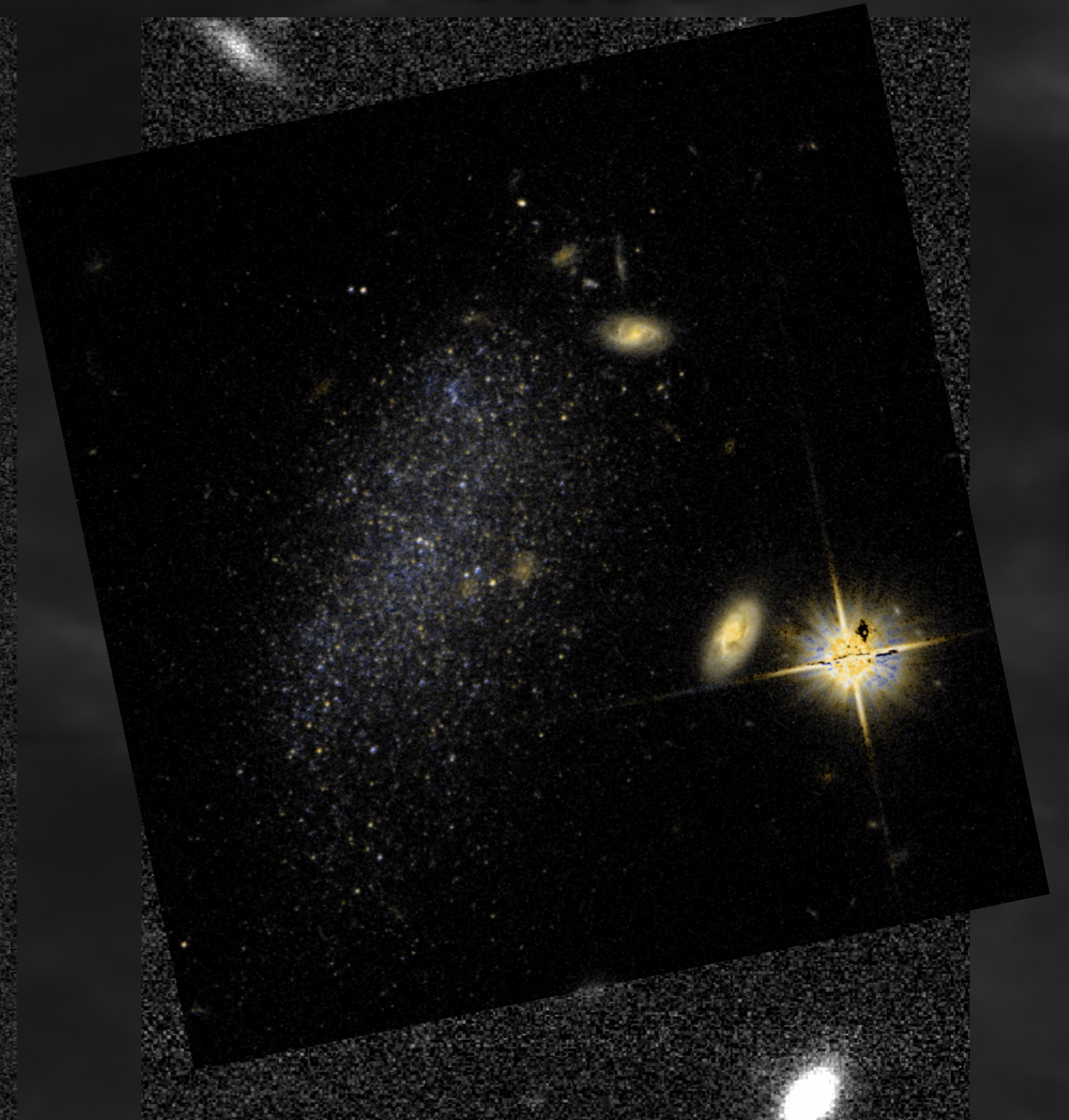
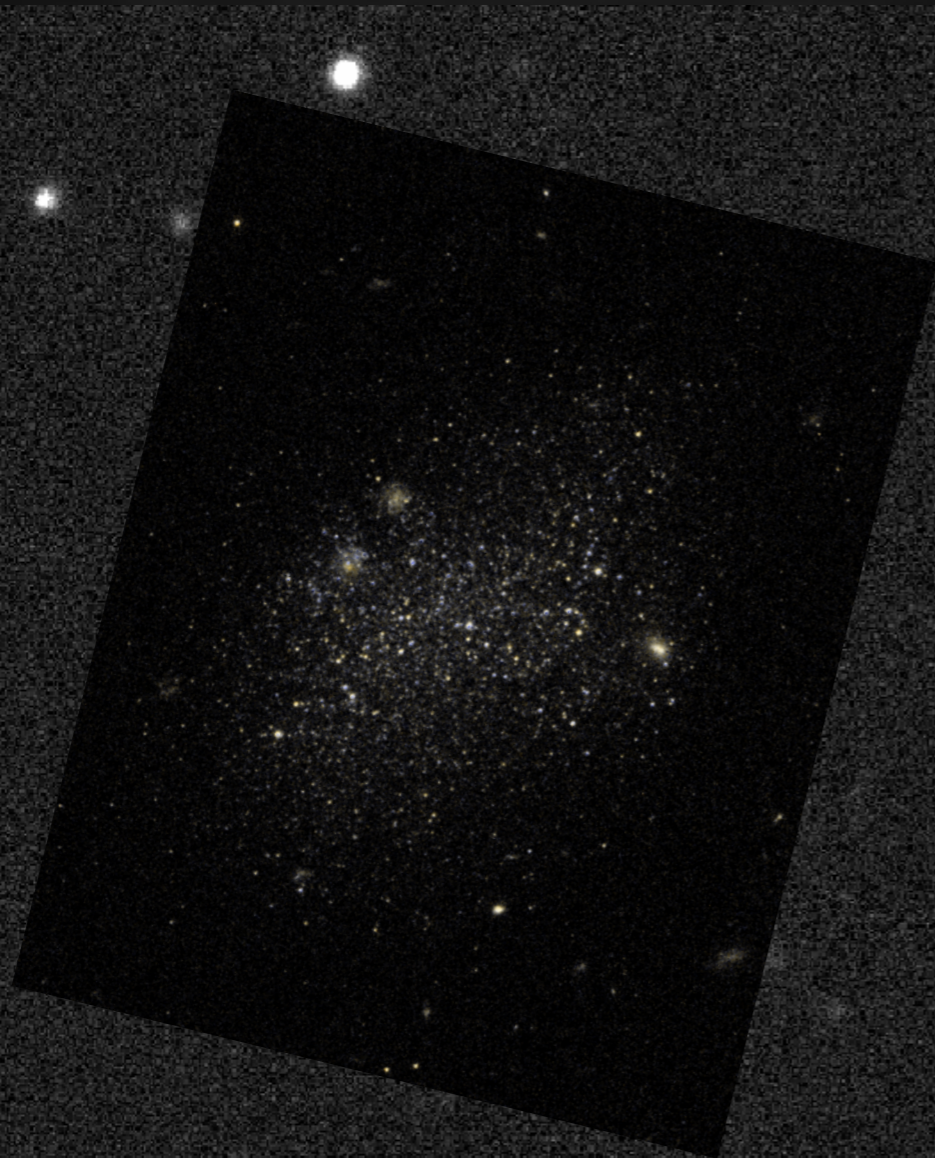




# Four (of ~ 30 candidates) have dwarf optical counterparts

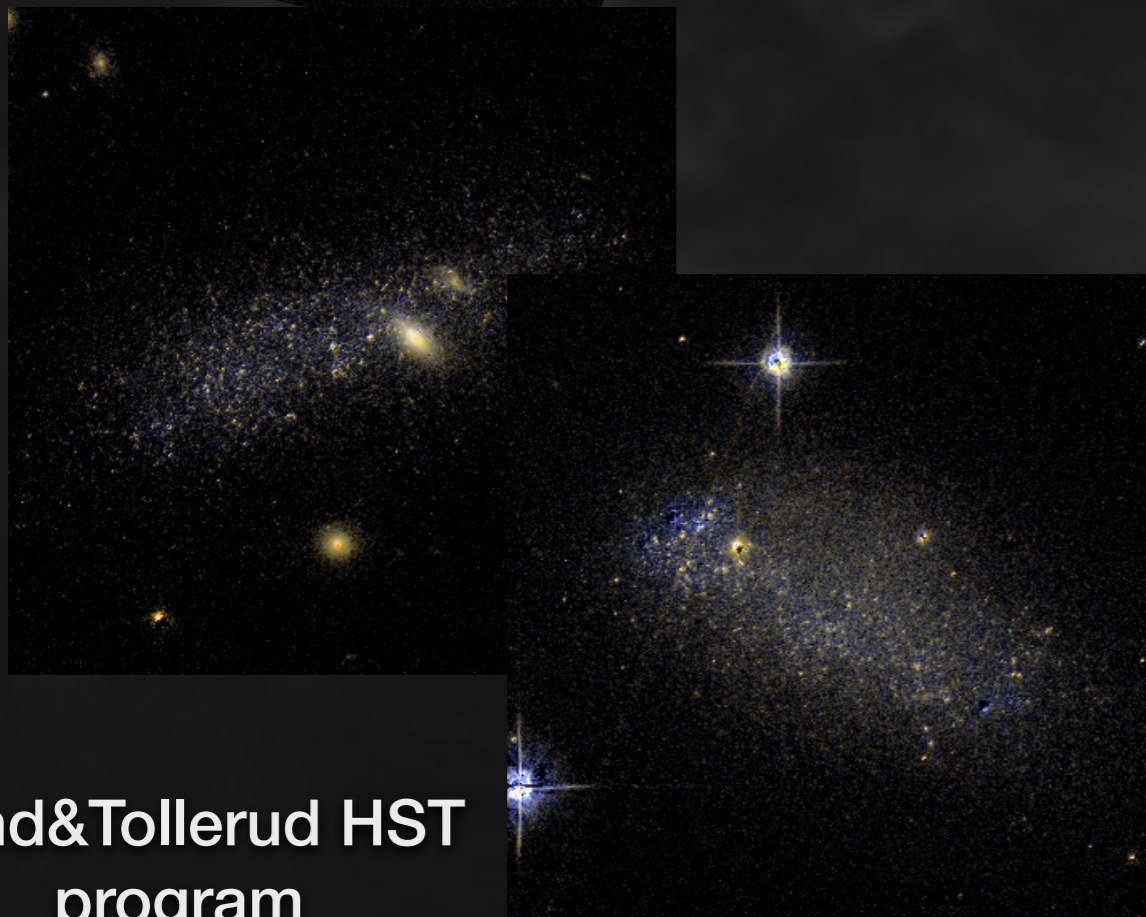
Pisces A

Pisces B

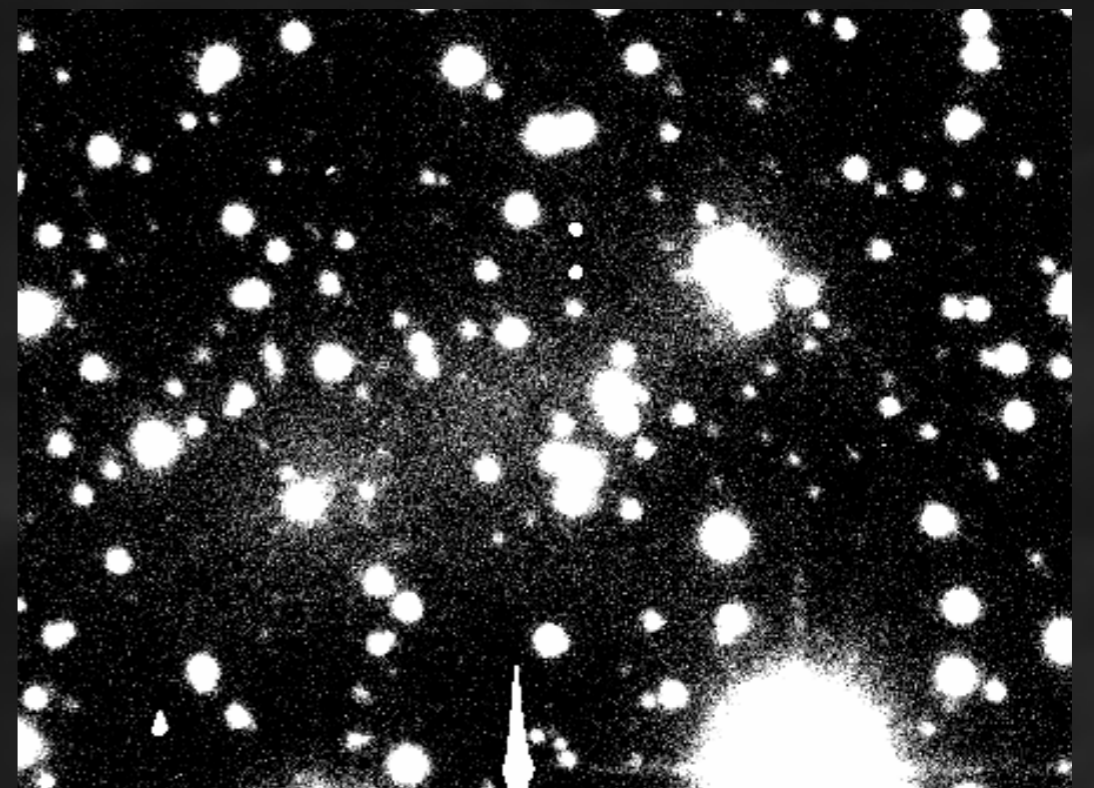




Four (of  $\sim 30$  candidates) have  
dwarf optical counterparts



Sand&Tollerud HST  
program



(And one candidate in the  
Galactic Plane)  
Tollerud et al. in prep

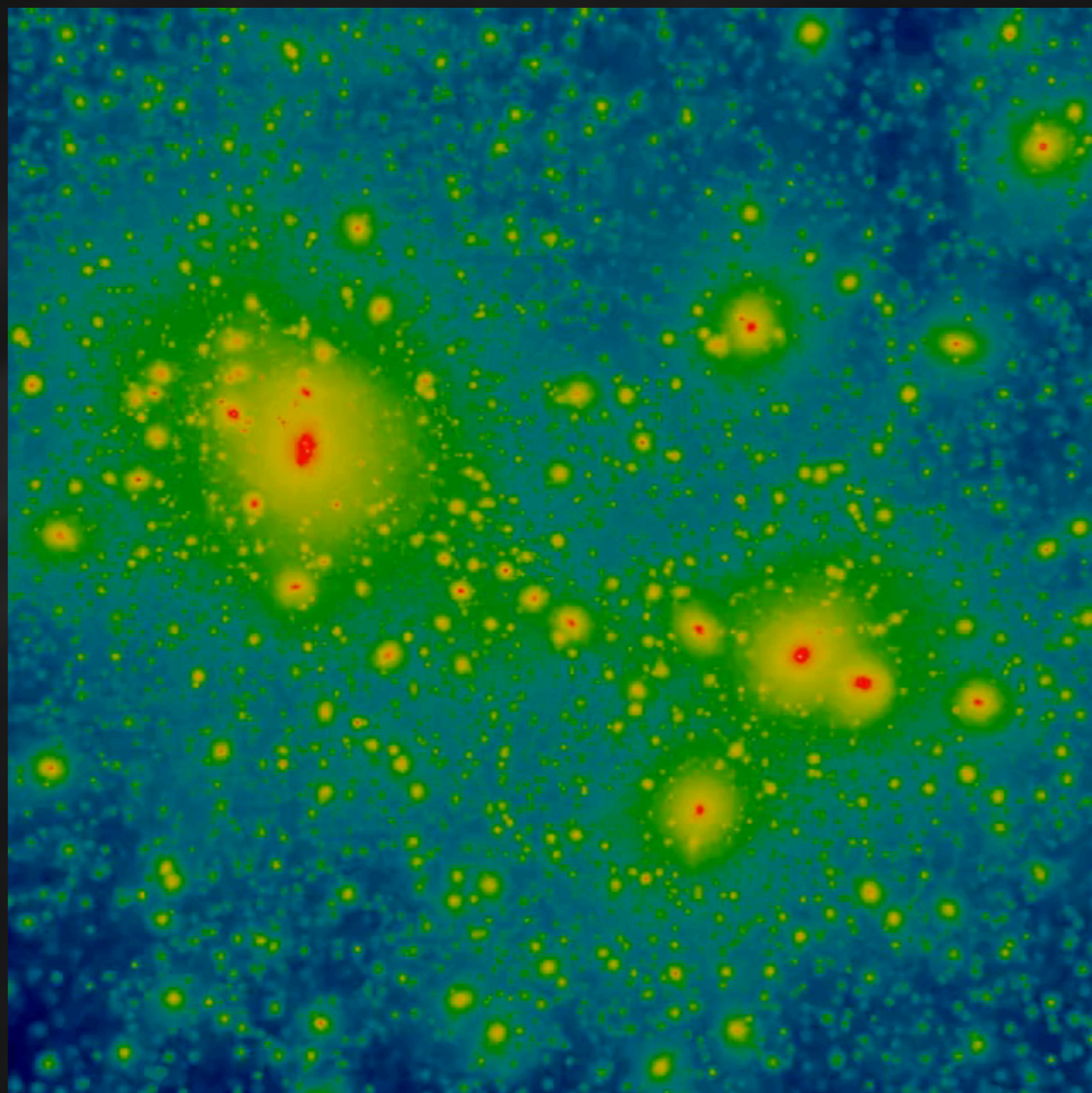


But none in the Local  
Group (i.e.,  $\approx 2$  Mpc)!

How many *should* we  
have found?



# Let's start with Local Group-like DM-only simulations



We put the MW in one (-> Earth and GALFA-HI)



**ELVIS**

(Garrison-Kimmel+ 14)

$$x12 * 2 = 24$$

And M31  
in the  
other



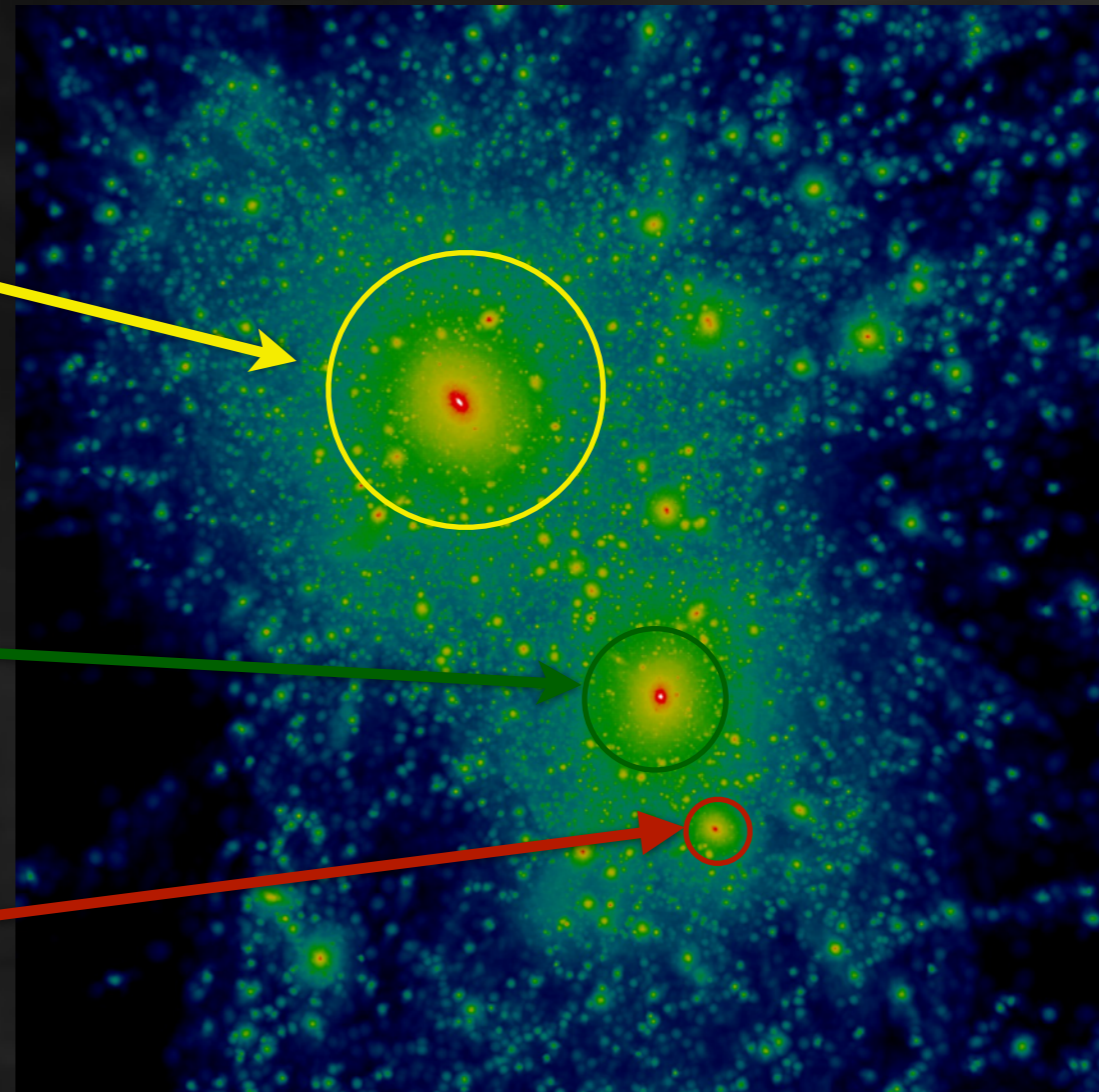
That's halo mass. But  
we need gas mass...



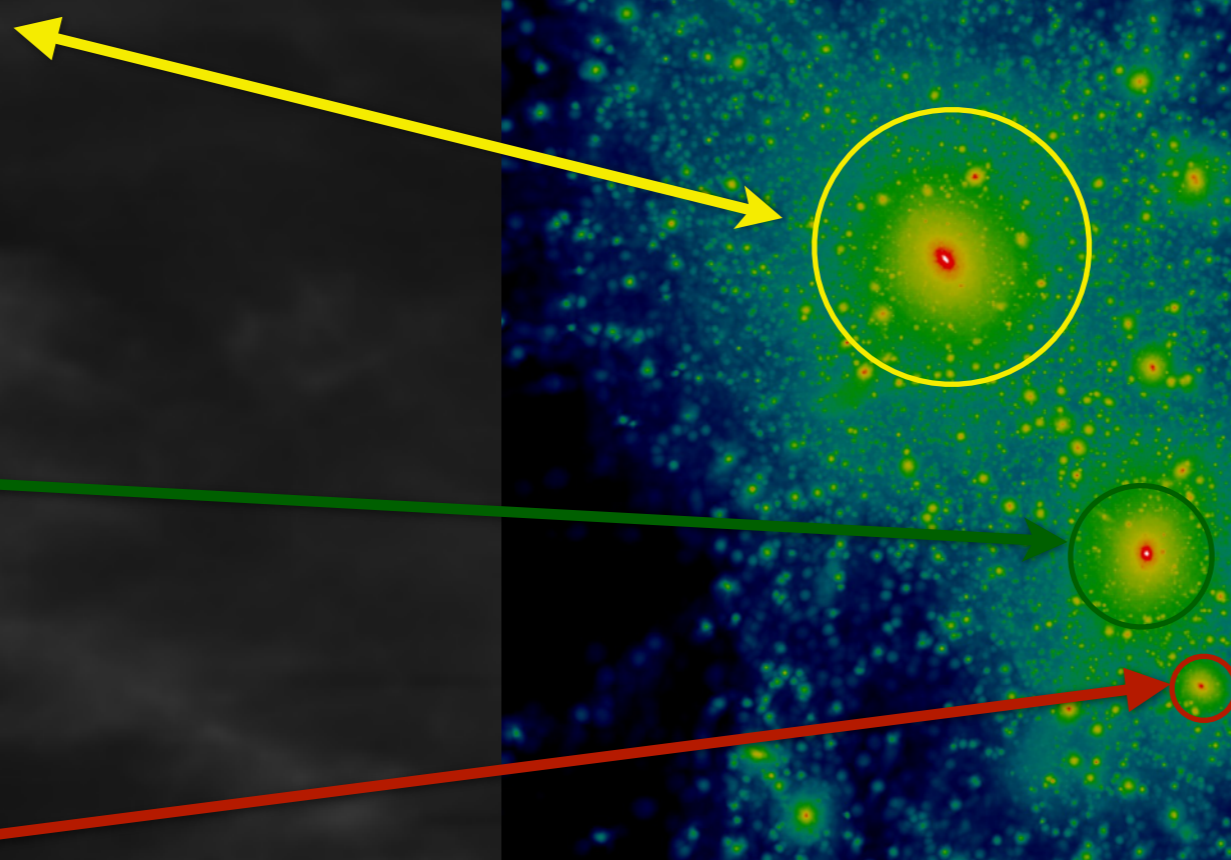
# Get to $M_*$ from $M_{\text{halo}}$ using Abundance Matching



Observed  
Galaxies



Simulations/  
Theory

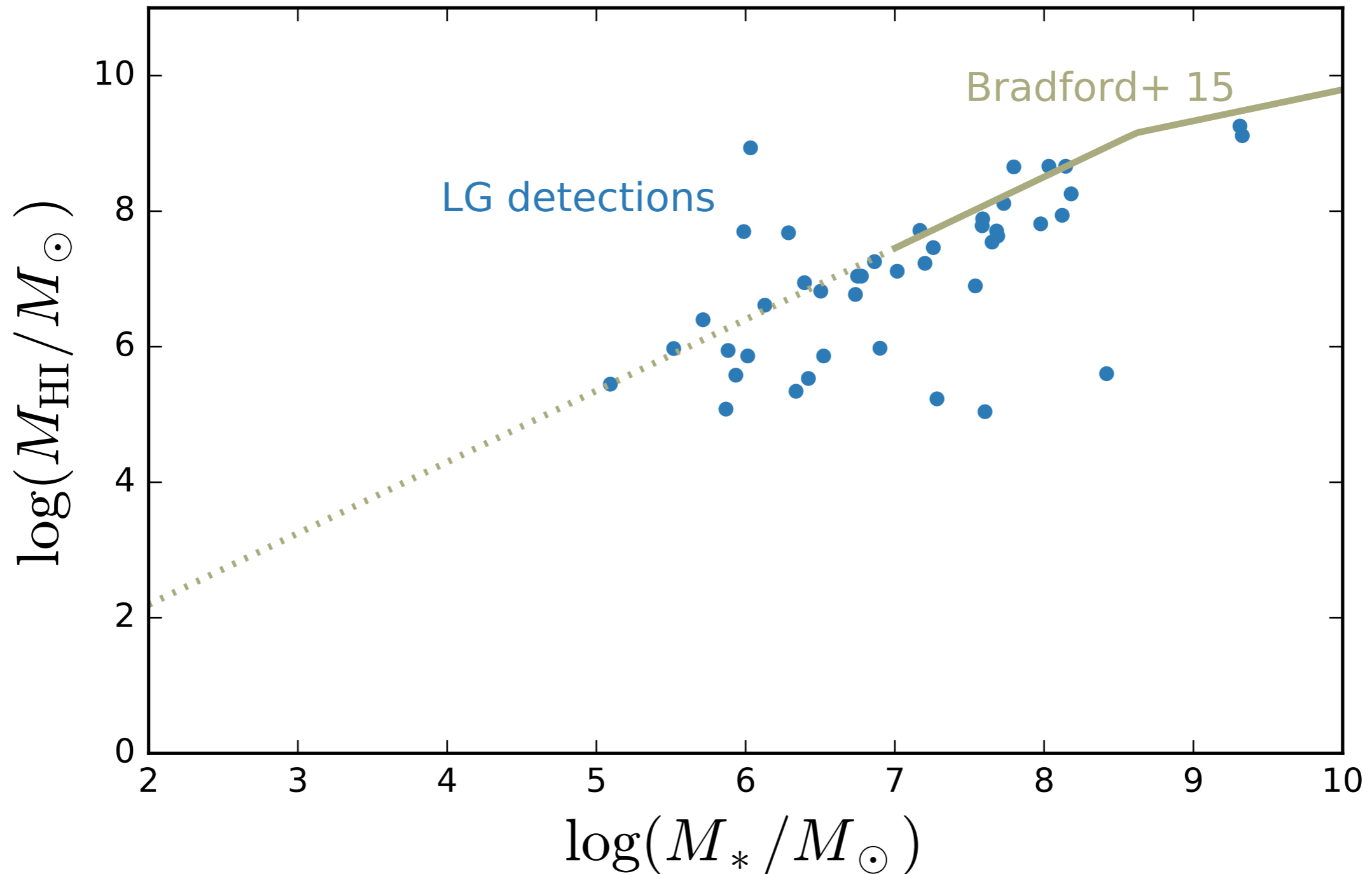




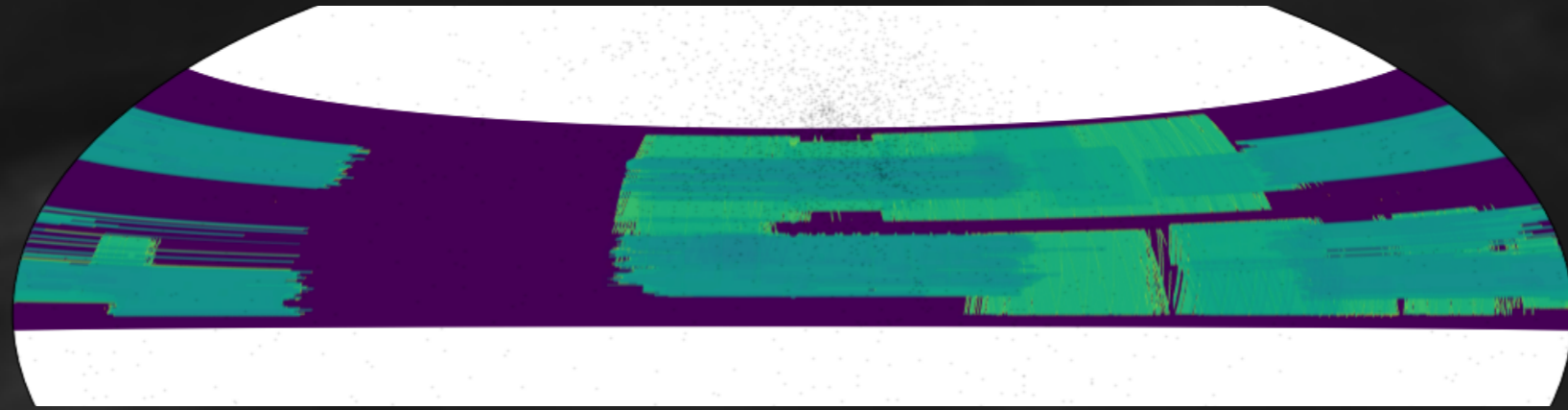
OK, that's  $M_*$ . But we  
wanted  $M_{\text{gas}}$  ...



OK, that's  $M_*$ . But we wanted  
 $M_{\text{gas}}$  : That's observed!



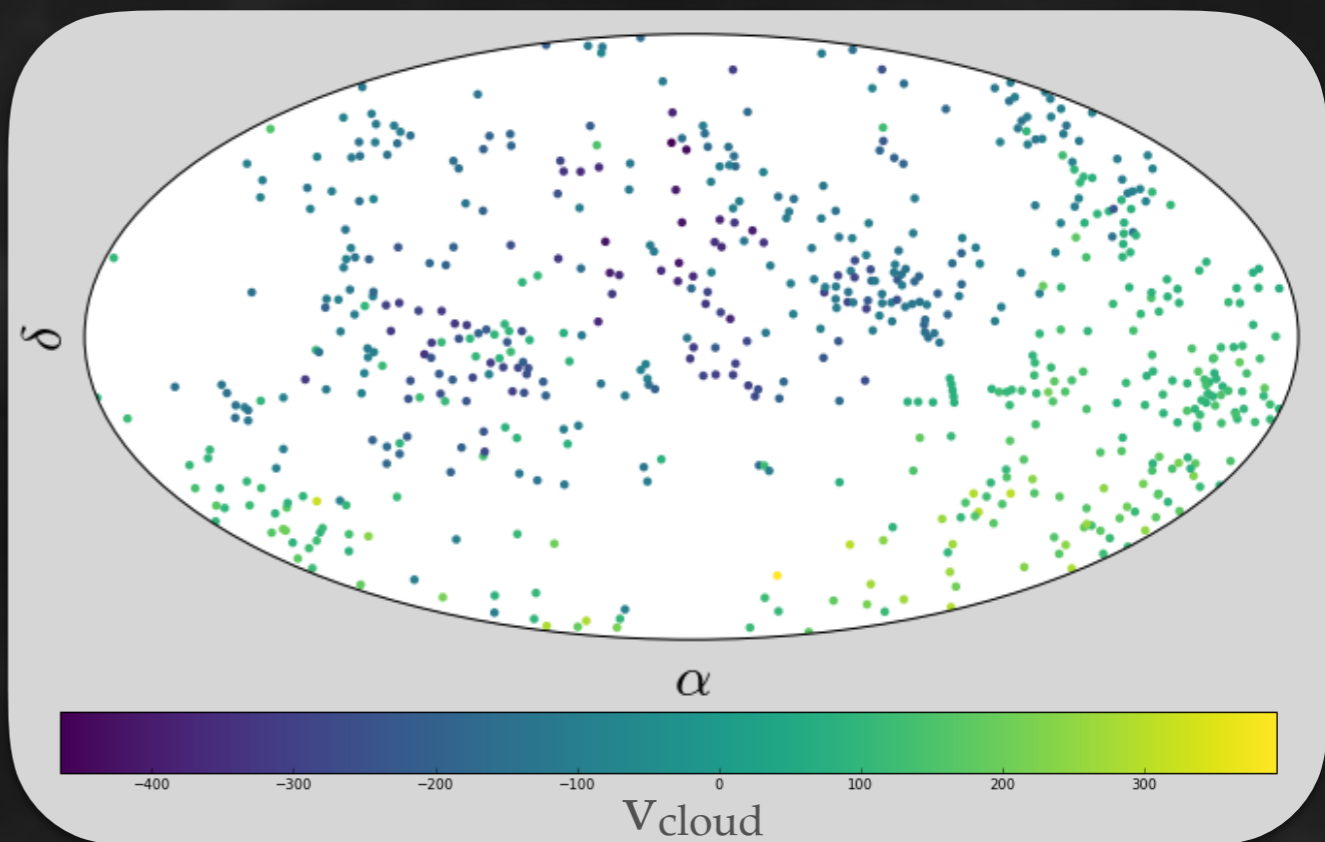
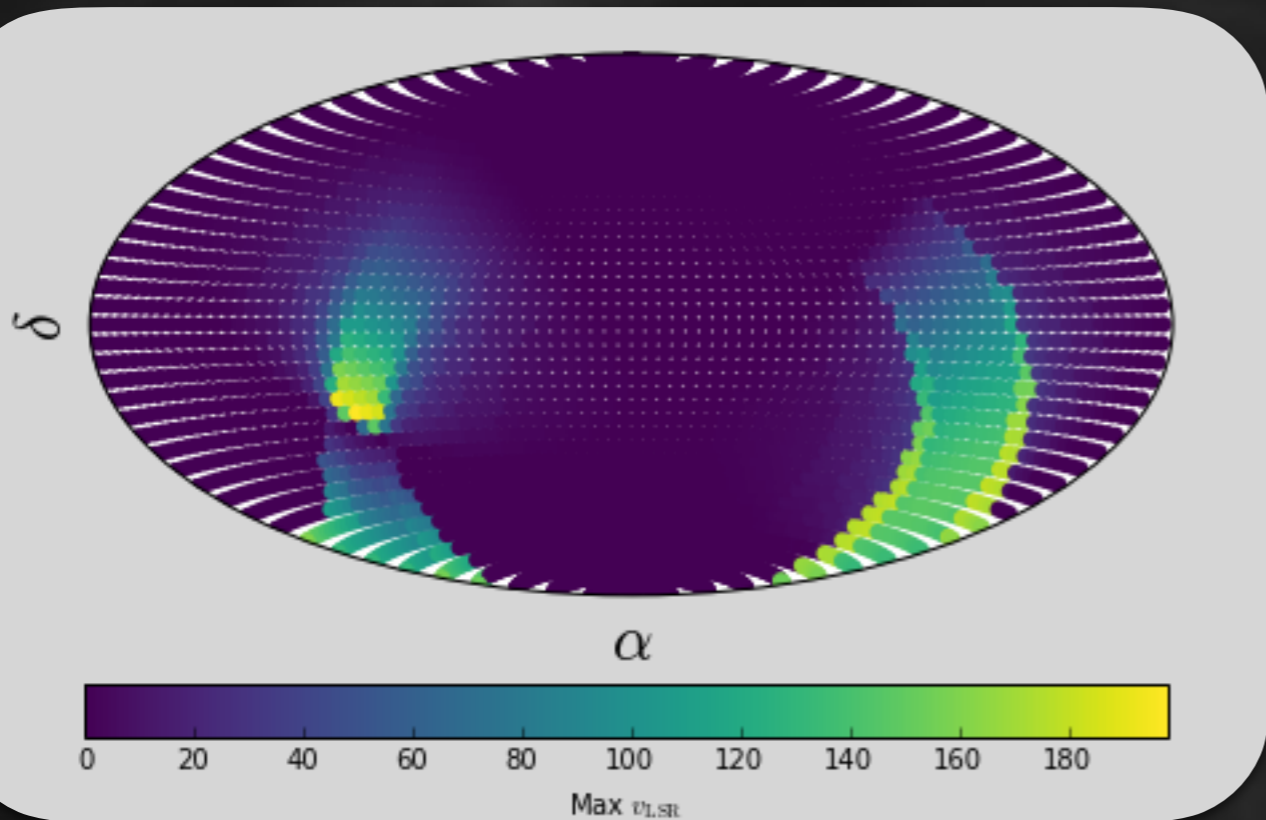
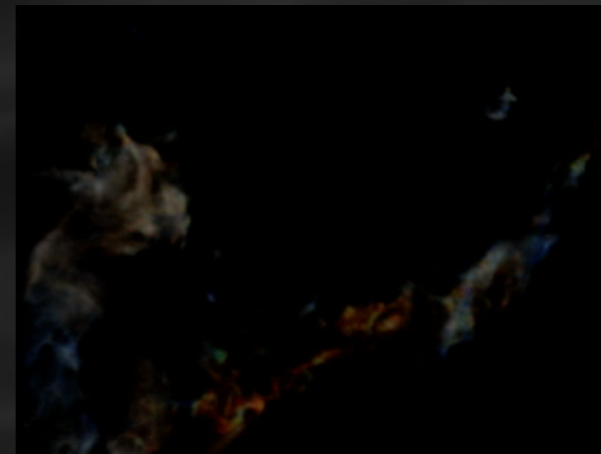
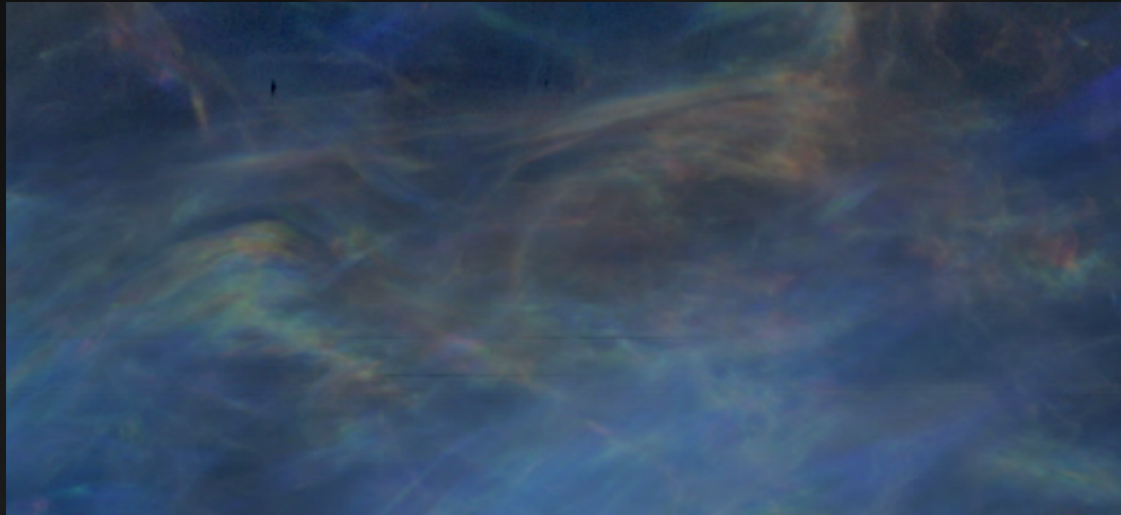
We've now got  $M_{\text{gas}}$  for each halo, so we know if it's detectable by GALFA-HI.



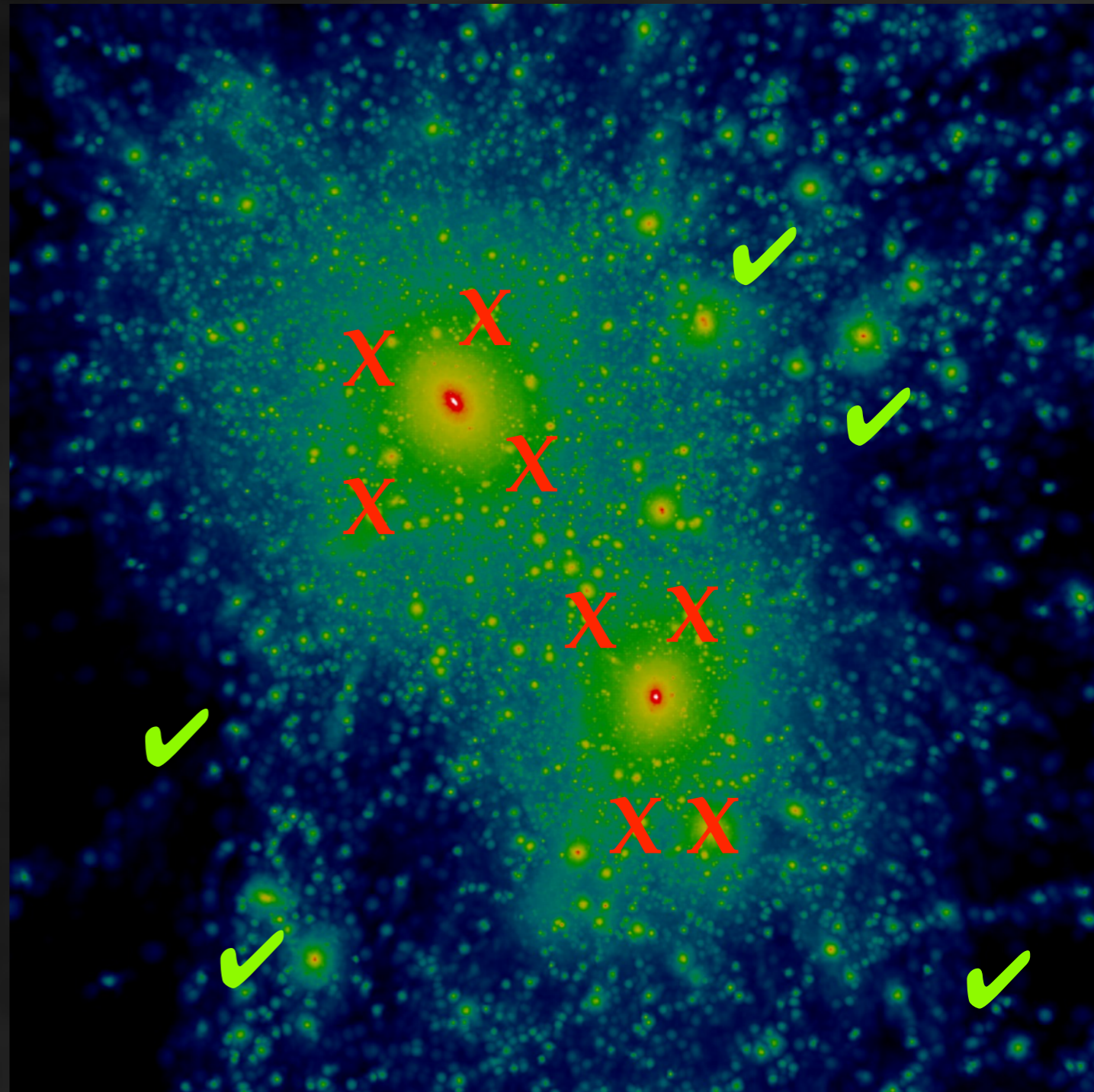
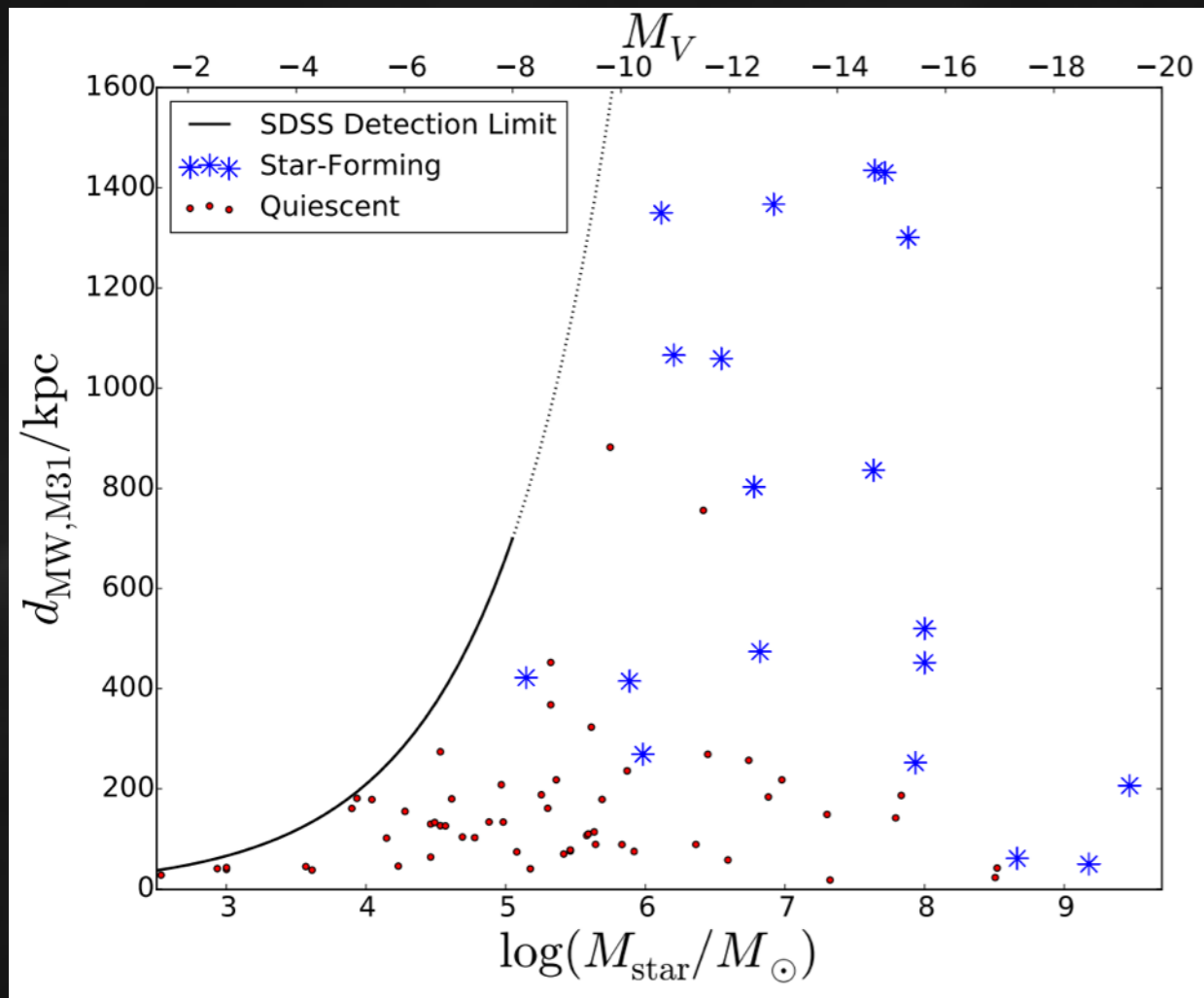
What else is in the observations that's not in the simulations?



# Need to remove Galactic crud exactly as in Observations



# Also need to account for satellite quenching



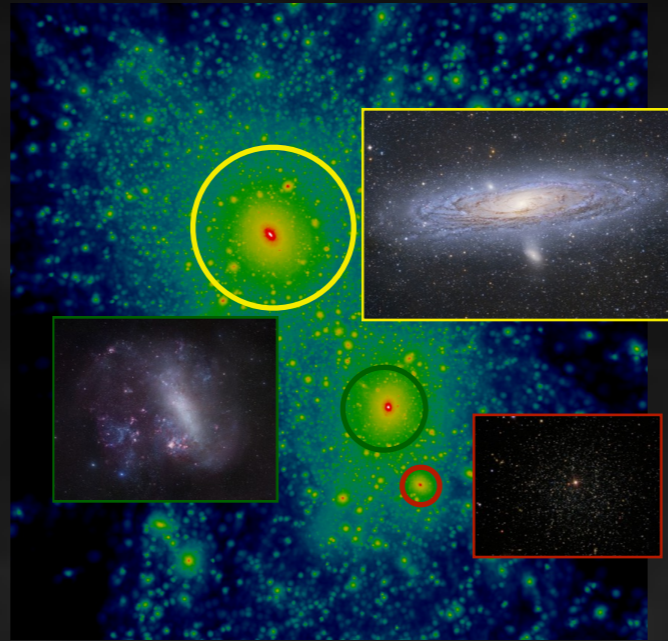


# Formula for a mock GALFA-HI

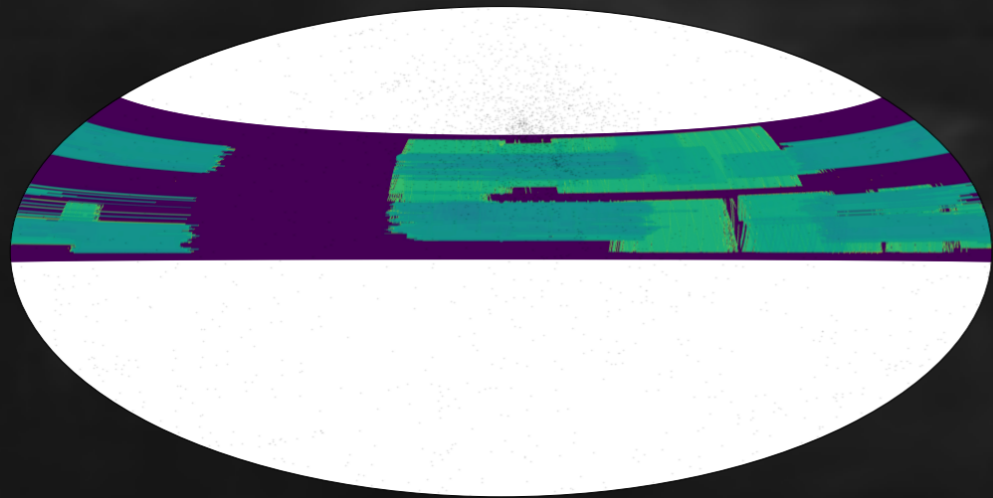
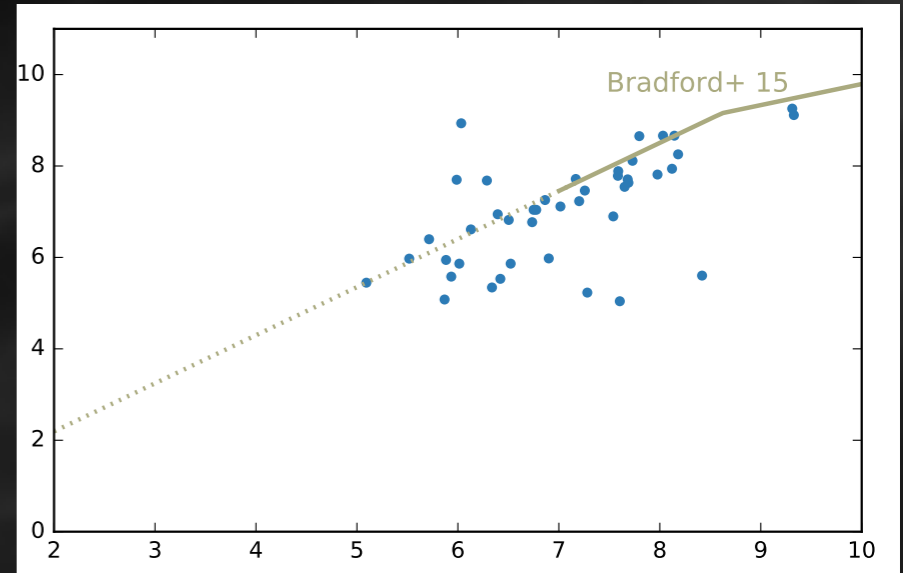
x24



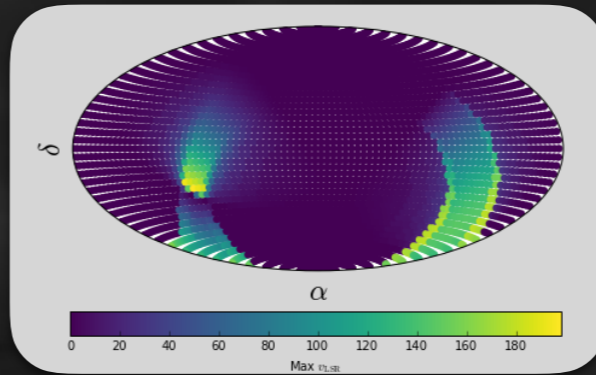
+



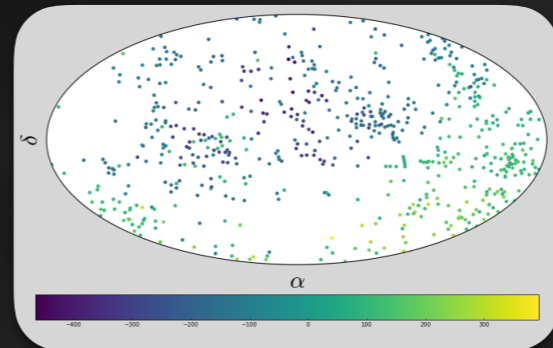
+



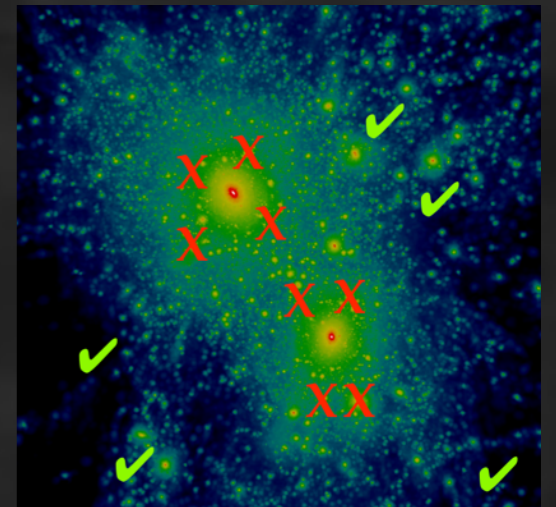
-



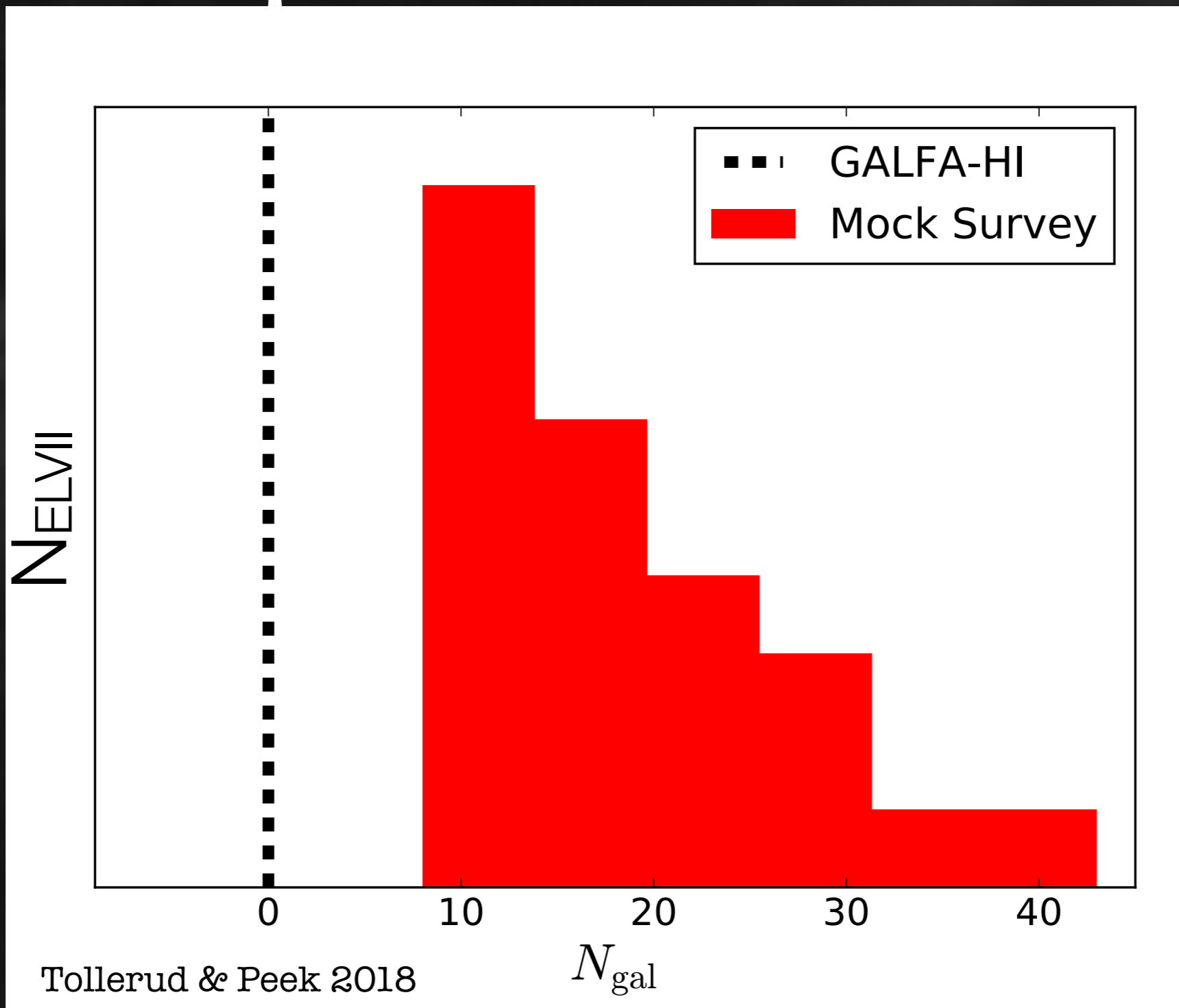
-



-



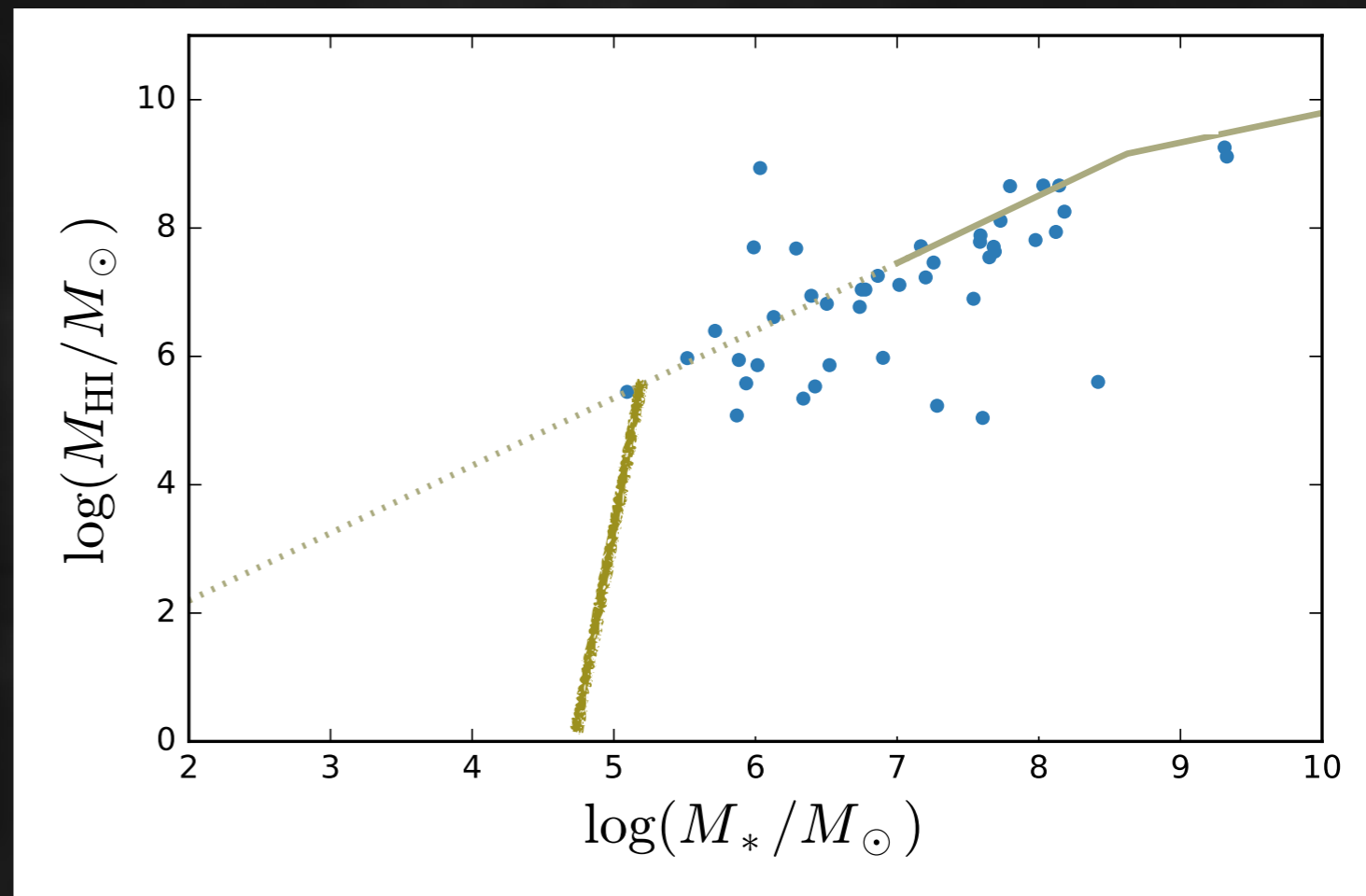
# Which gives us... quite a few!



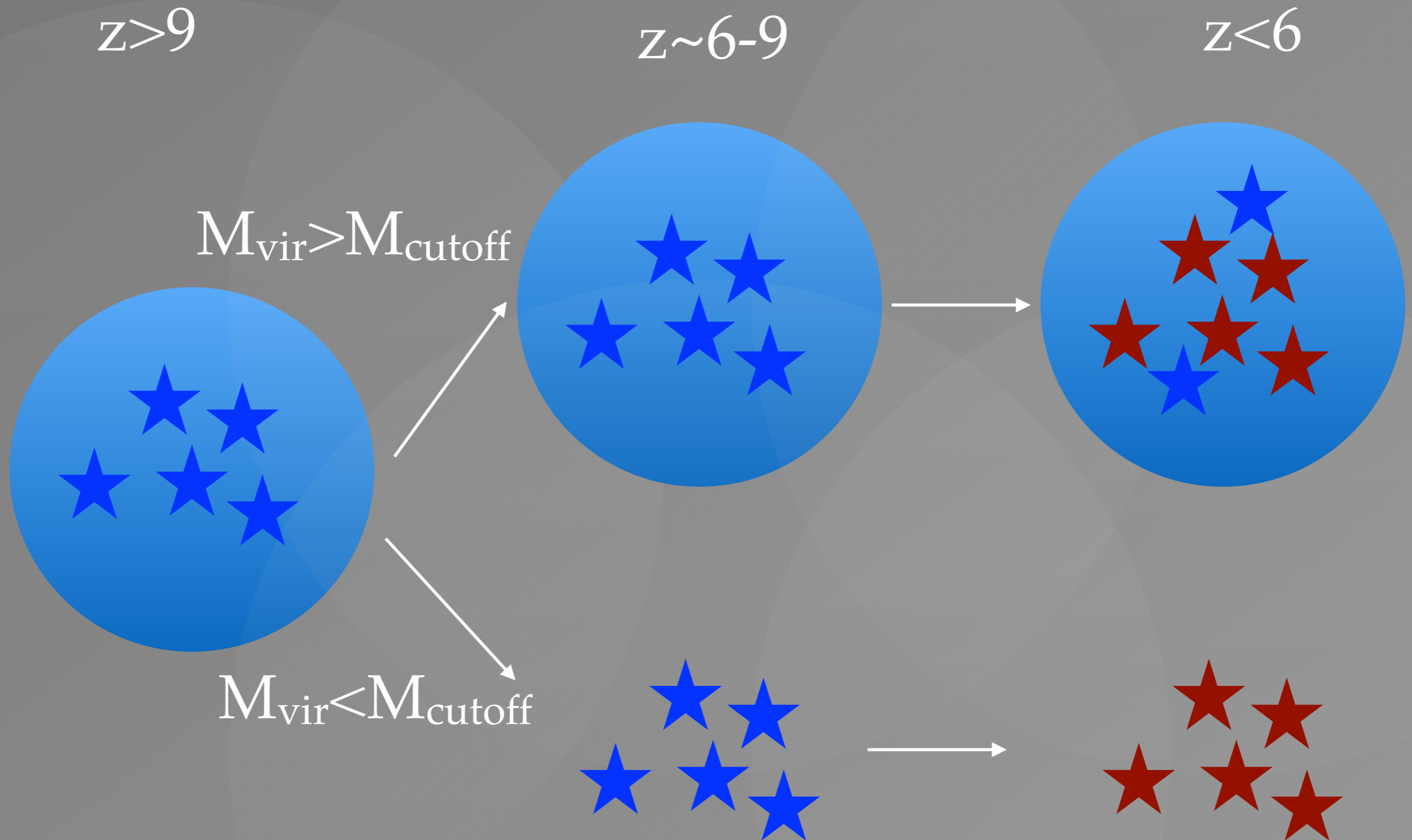


# Why might there be a “Missing dIrr Crisis- Problem-isaster-portunity”?

1. The LG is a  $\approx 1/24$  outlier (SAGA?)
2.  $\Lambda$ CDM is wrong
3. The  $M_{\text{HI}}-M_{\text{Halo}}$  relation has a break
  - Reionization?

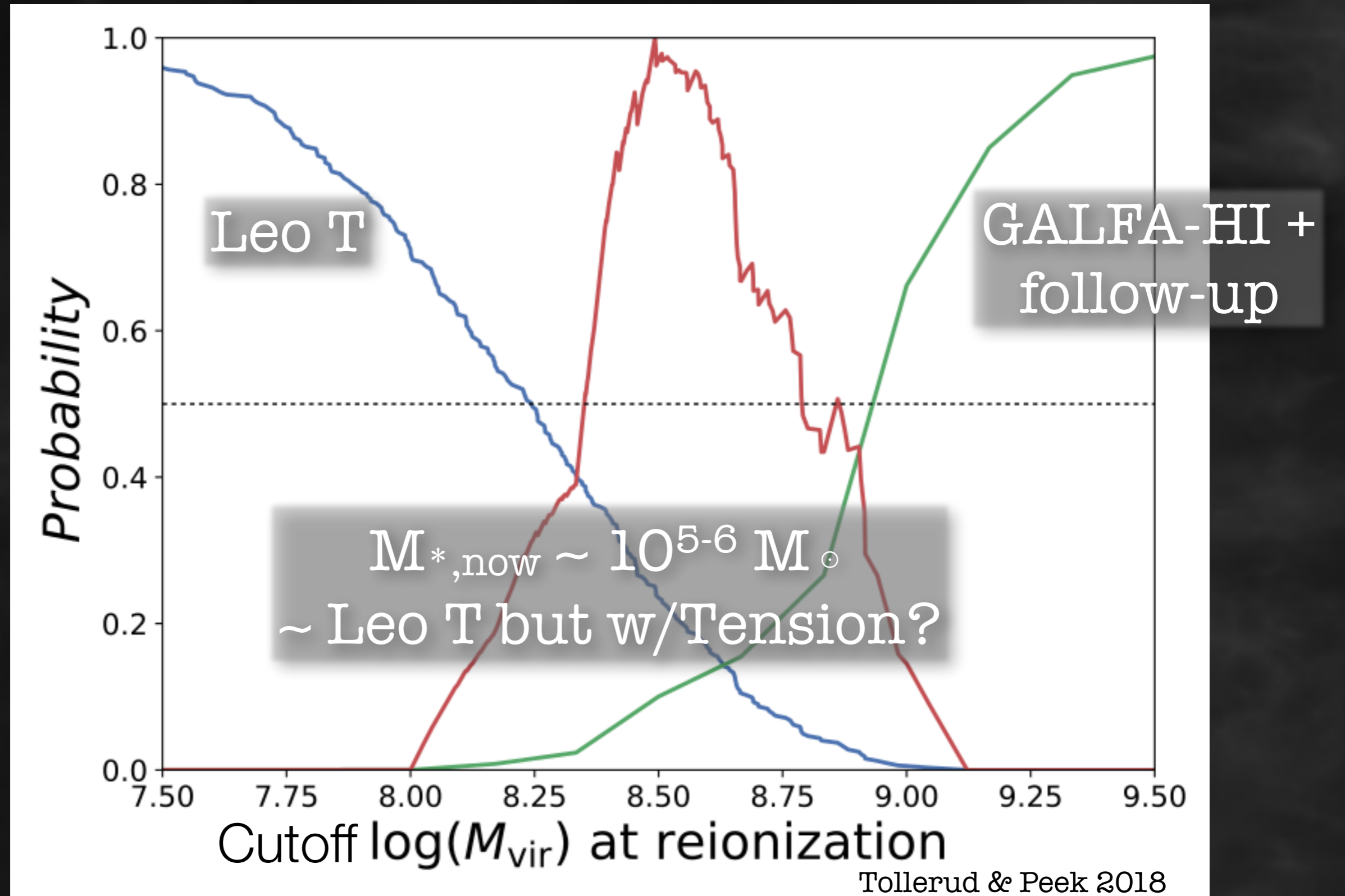


# How does Reionization Help?

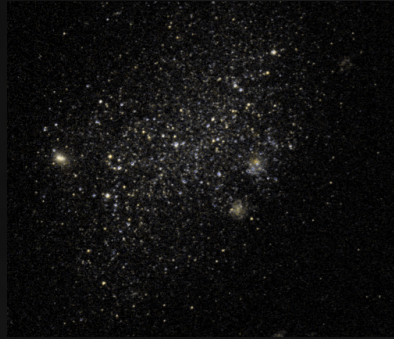




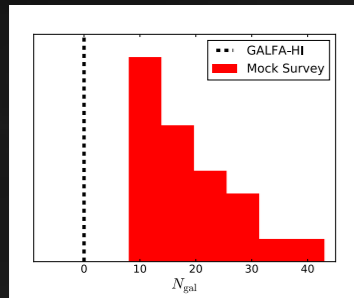
# At what mass does Rel zap dwarfs?



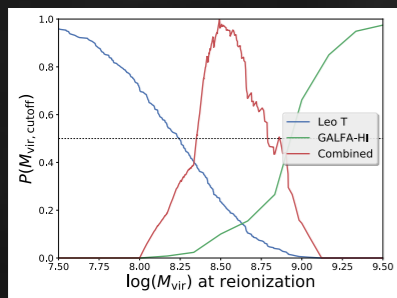
# Summary



- ◆ Local Dwarf Galaxies can be found by HI surveys + deep follow-up (but none in LG)



- ◆ This +  $\Lambda$ CDM simulations suggest a missing dLrr (gas-rich) dwarfs problem



- ◆ This may be a limit on the reionization halo/galaxy mass scale (independent of star formation histories):

- ◆  $M_{\text{vir, reionization}} \sim 10^{8.5-9} M_{\odot}$
- ◆  $M_{*, \text{ now}} \sim 10^{5-6} M_{\odot}$
- ◆ *Maybe* some tension w/ Leo T?